

# Protons, electrons and muons

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Many thanks to my supervisors:

F. Ptochos (UCy), J. Pasternak and K. Long (Imperial College London),  
Y. Papaphilippou (CERN), R. Bartolini and A. Seryi (Oxford University)

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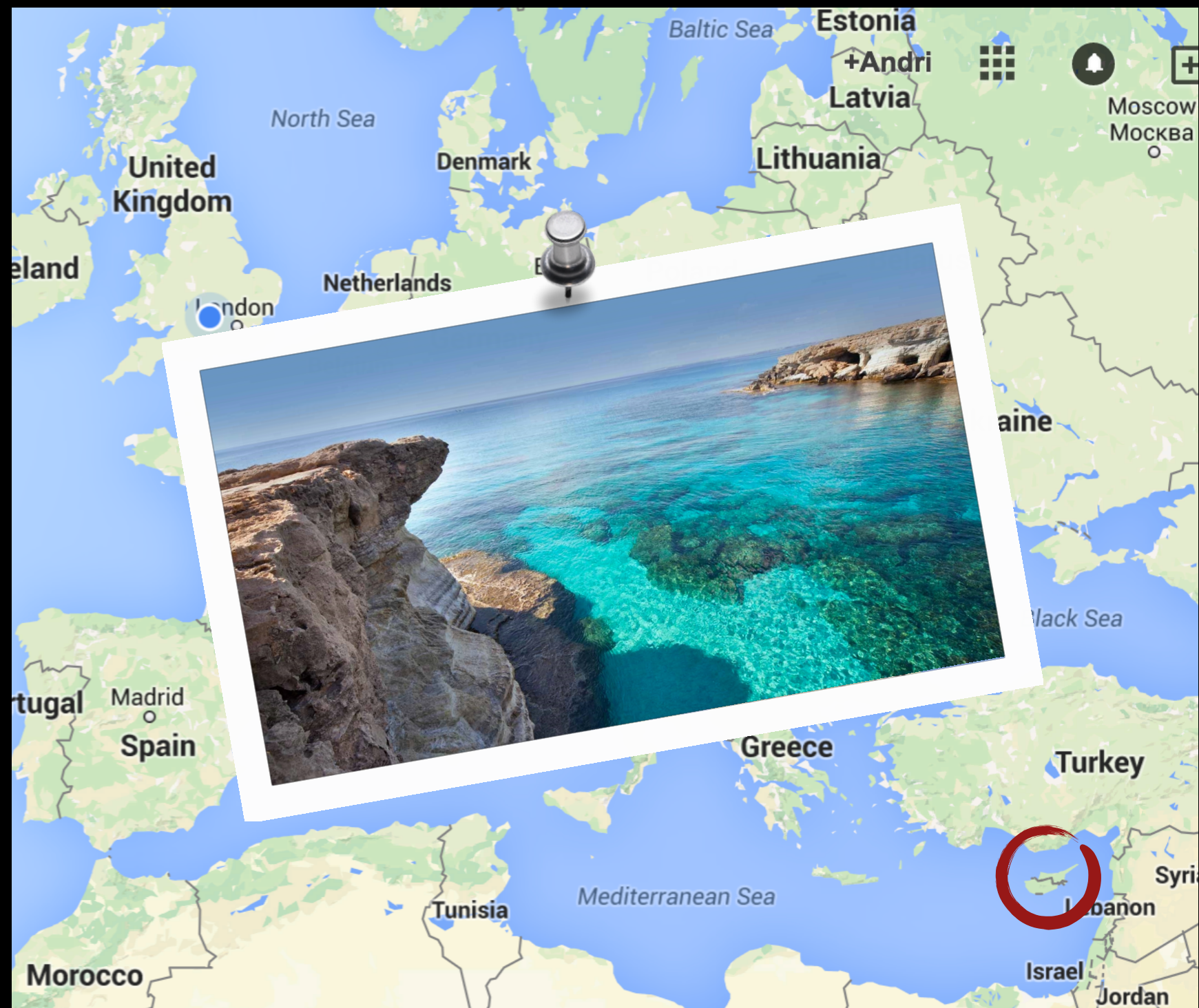




















2004

2008



BSci Physics  
University of  
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\* Final year thesis,  
FNAL, CDF,  
“Upsilon Meson  
Polarisation”

● R&D and simulations

\* running experiment

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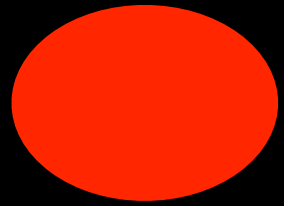
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# Neutrino Factory

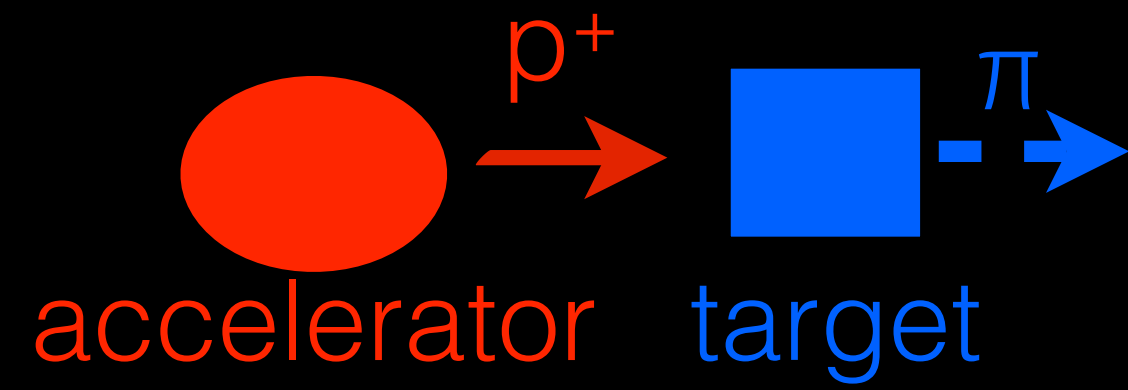
# Neutrino Factory

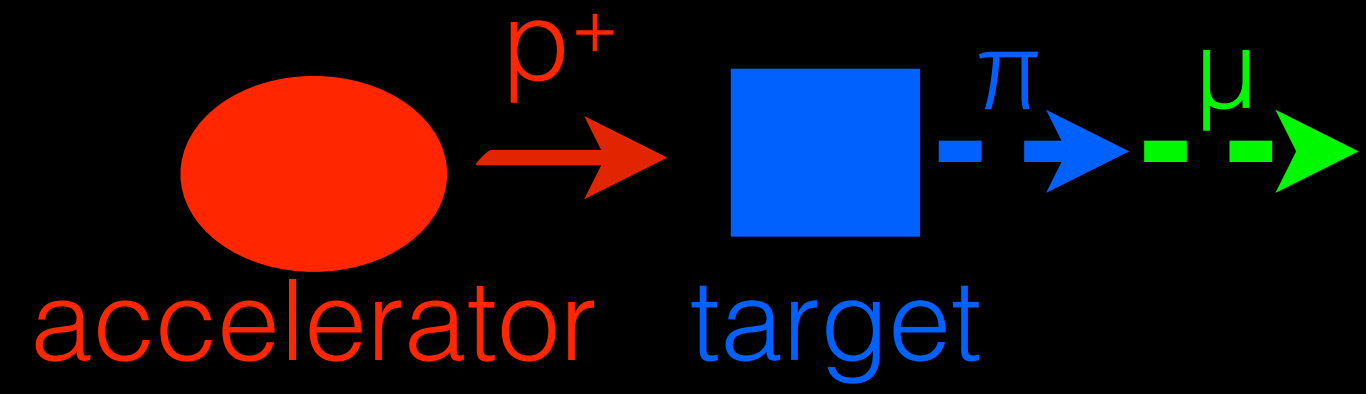
Future project. Aims to produce high intensity neutrino beams to determine mass hierarchy and measure neutrino parameters in unprecedented precision [1]

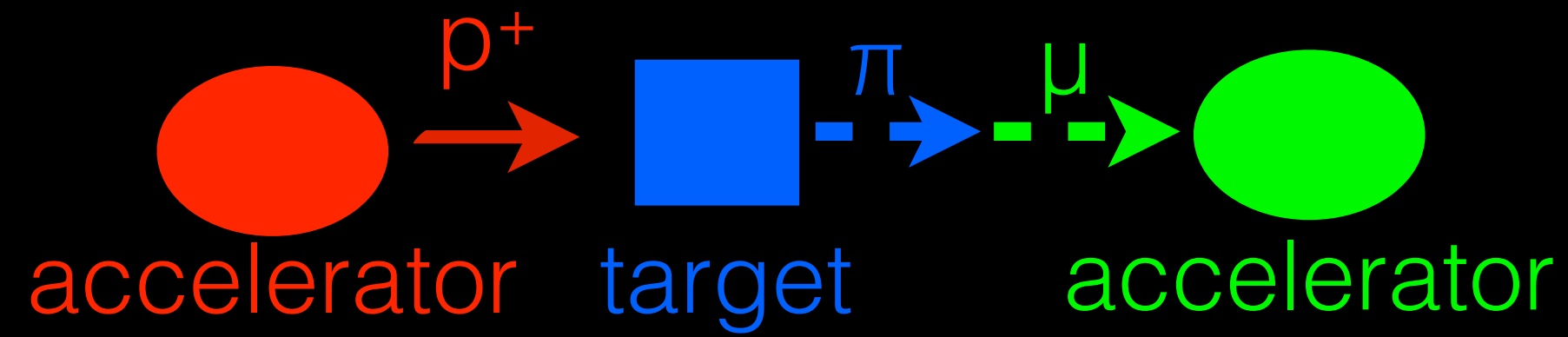


accelerator

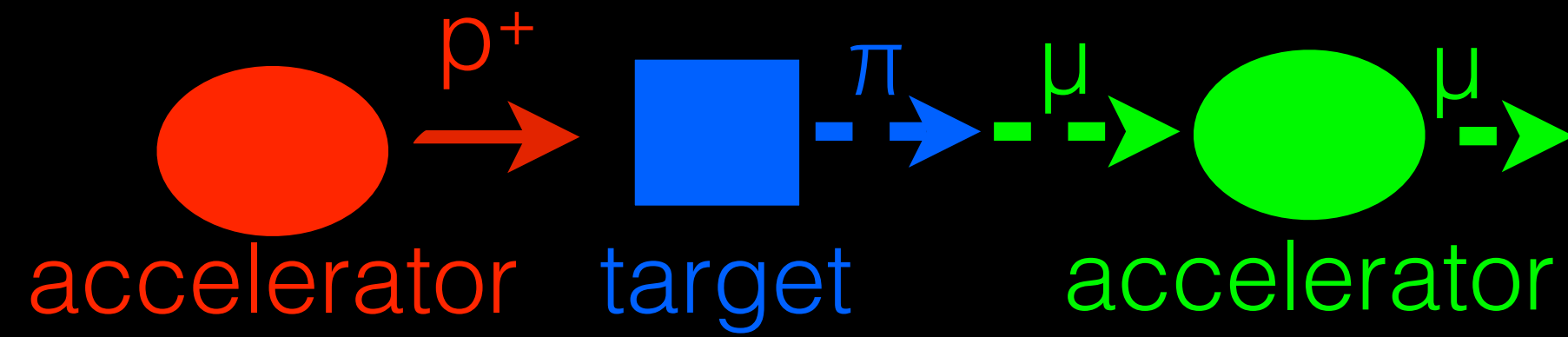


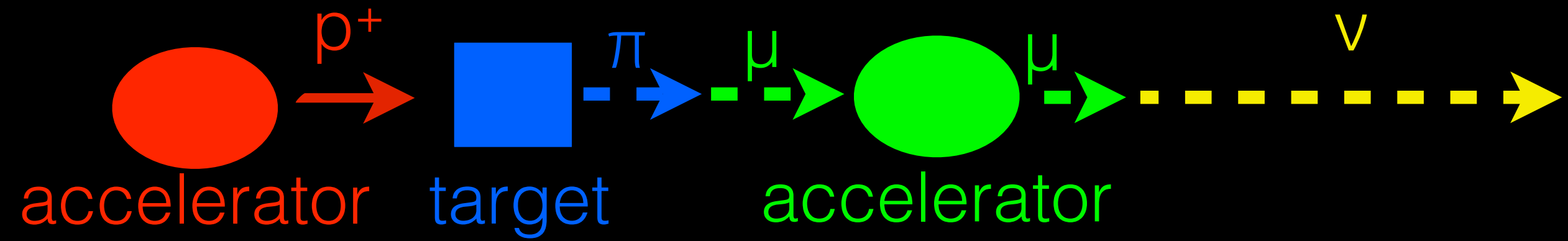


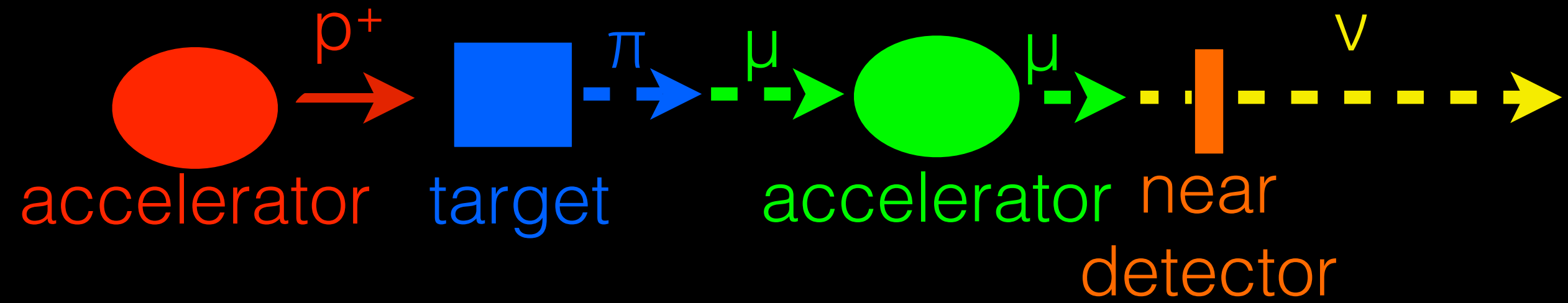


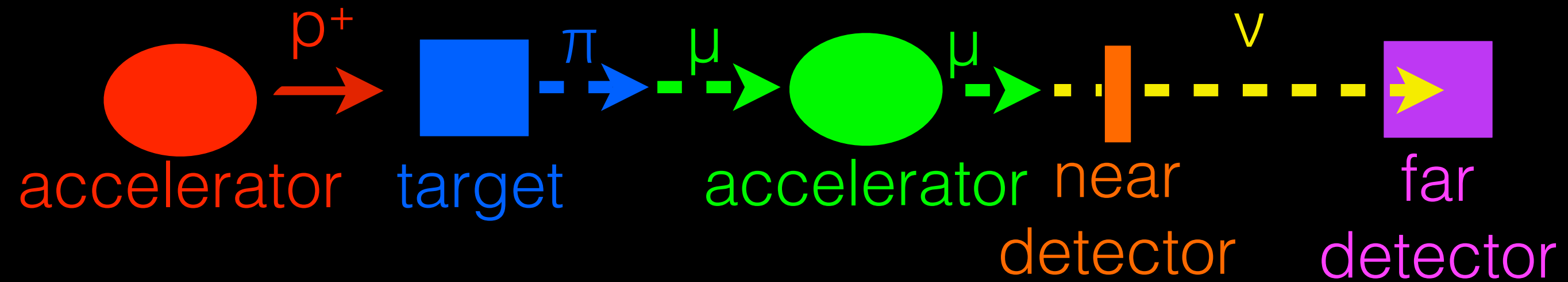


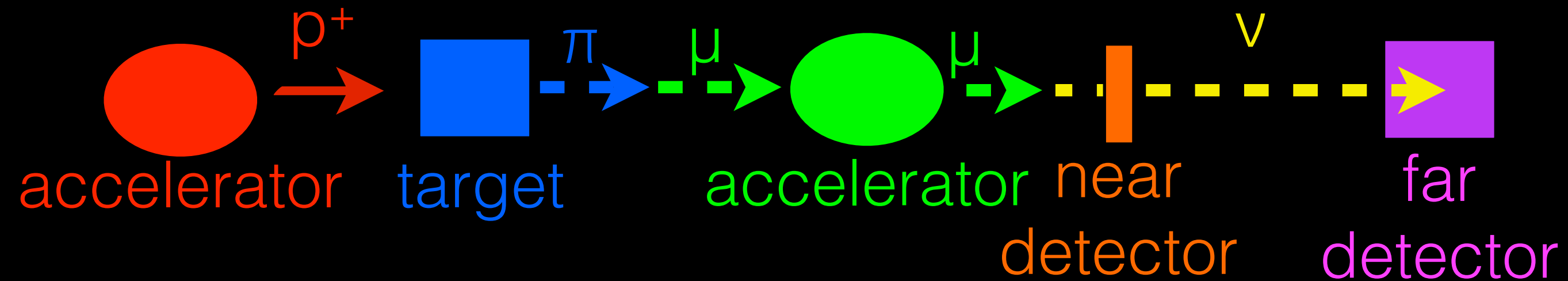




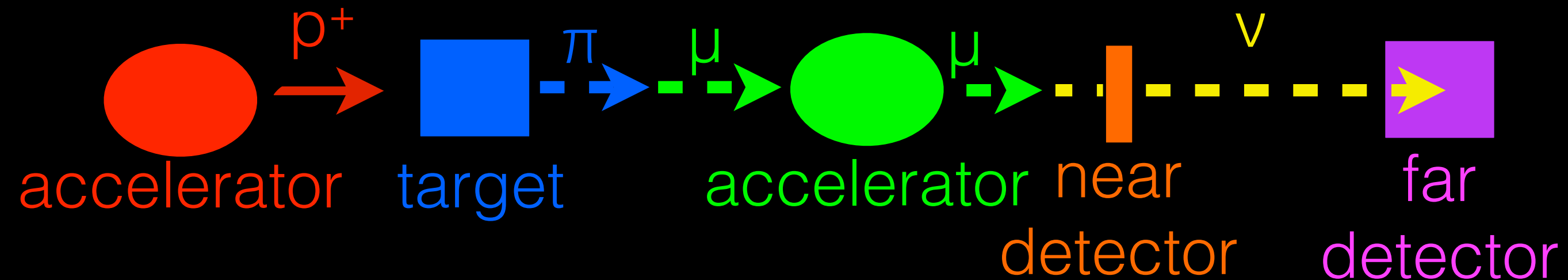






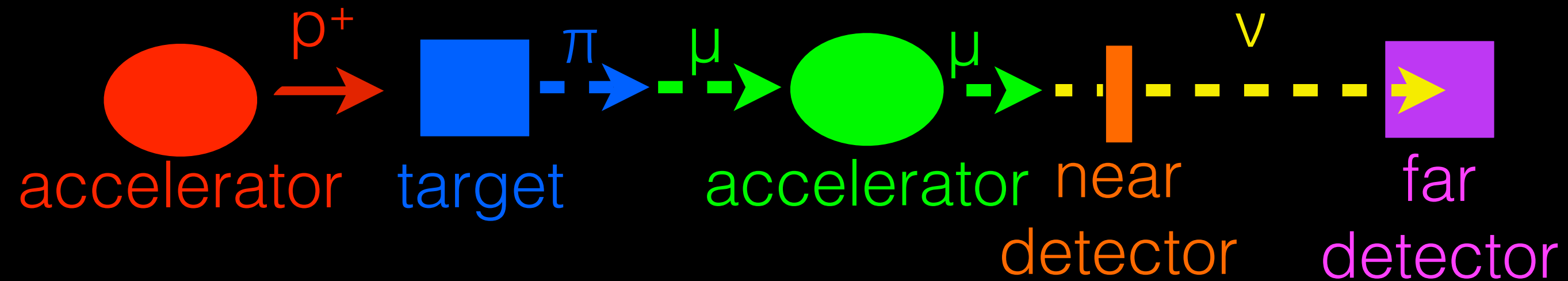


*Problem:* when muons are produced they occupy a large transverse phase-space

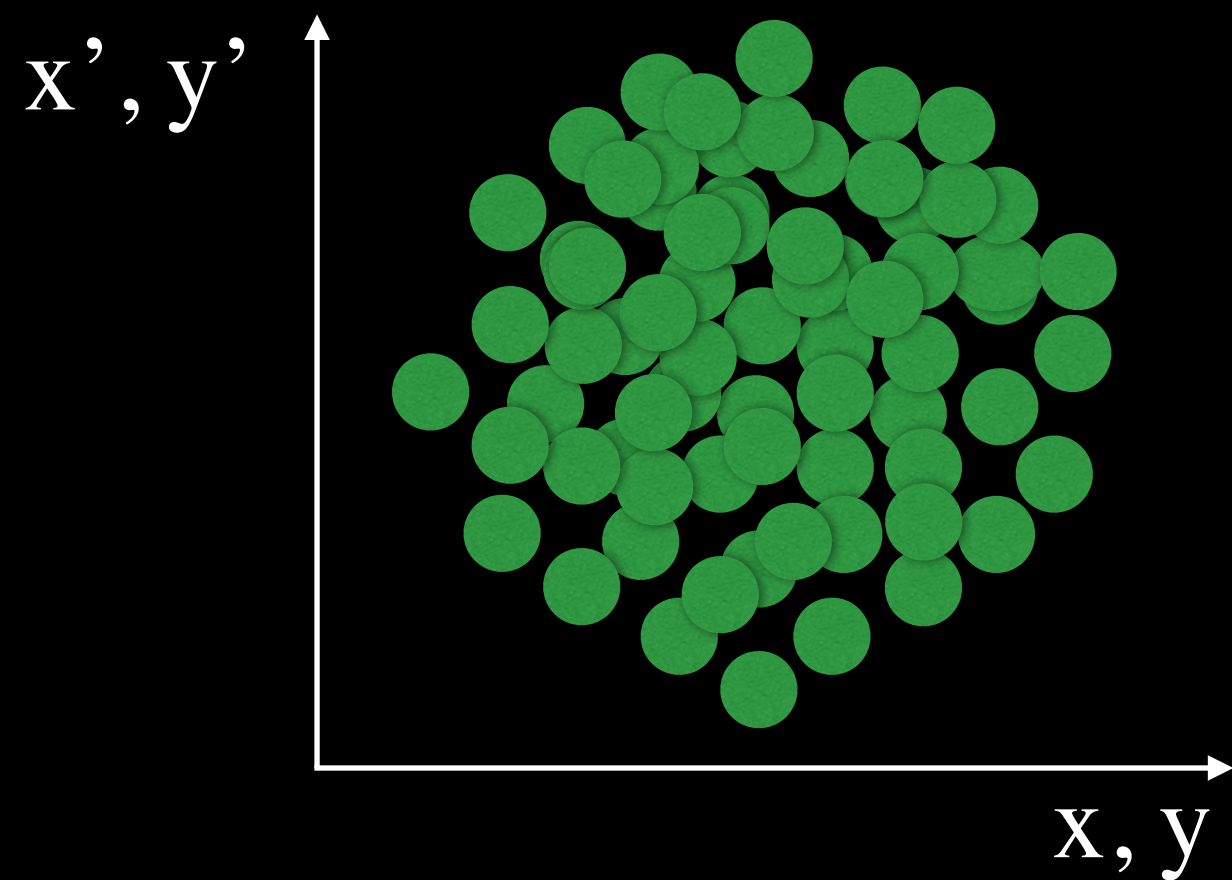


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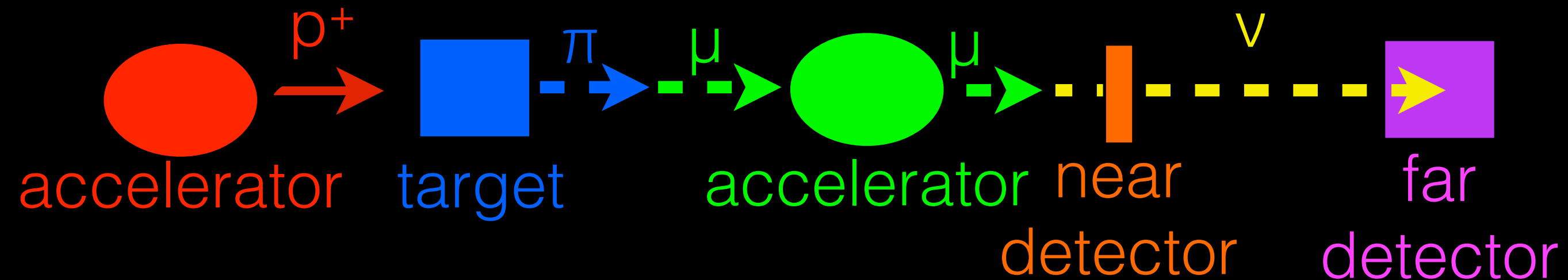
In order for muons to fit within acceptance of downstream accelerator, phase-space needs to reduce



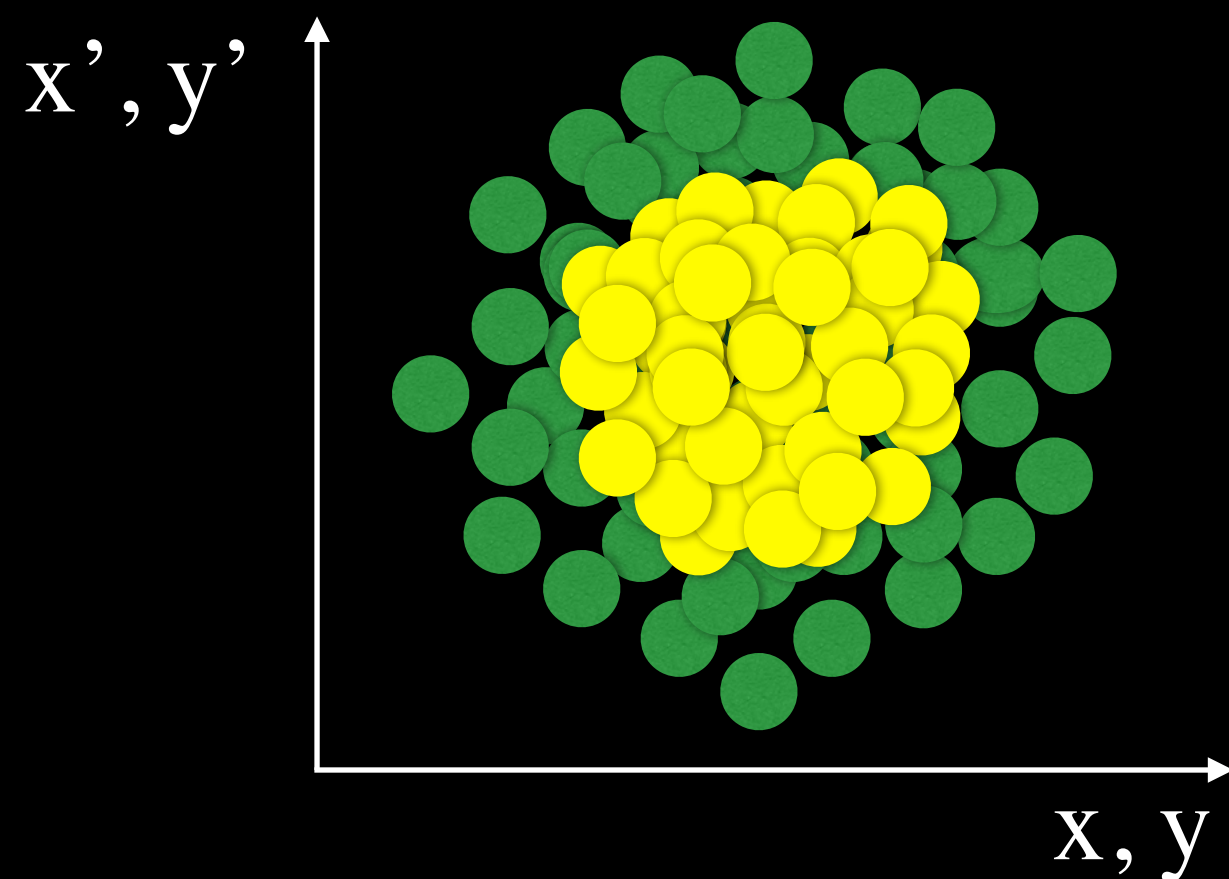
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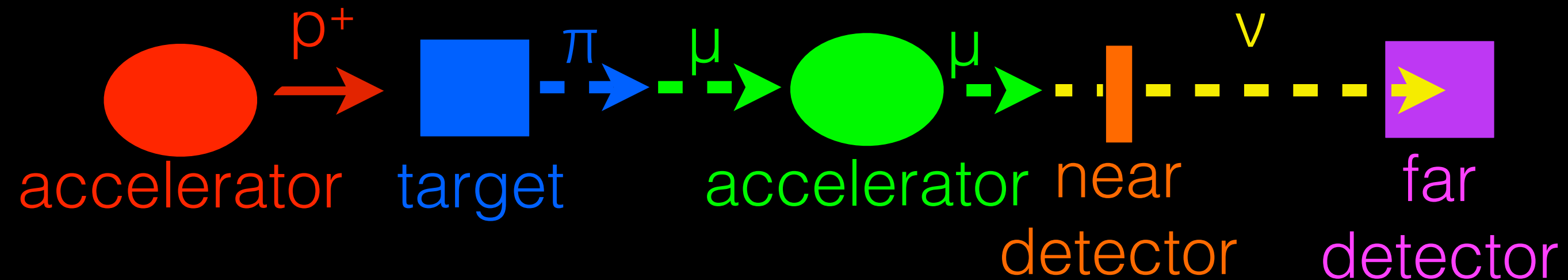


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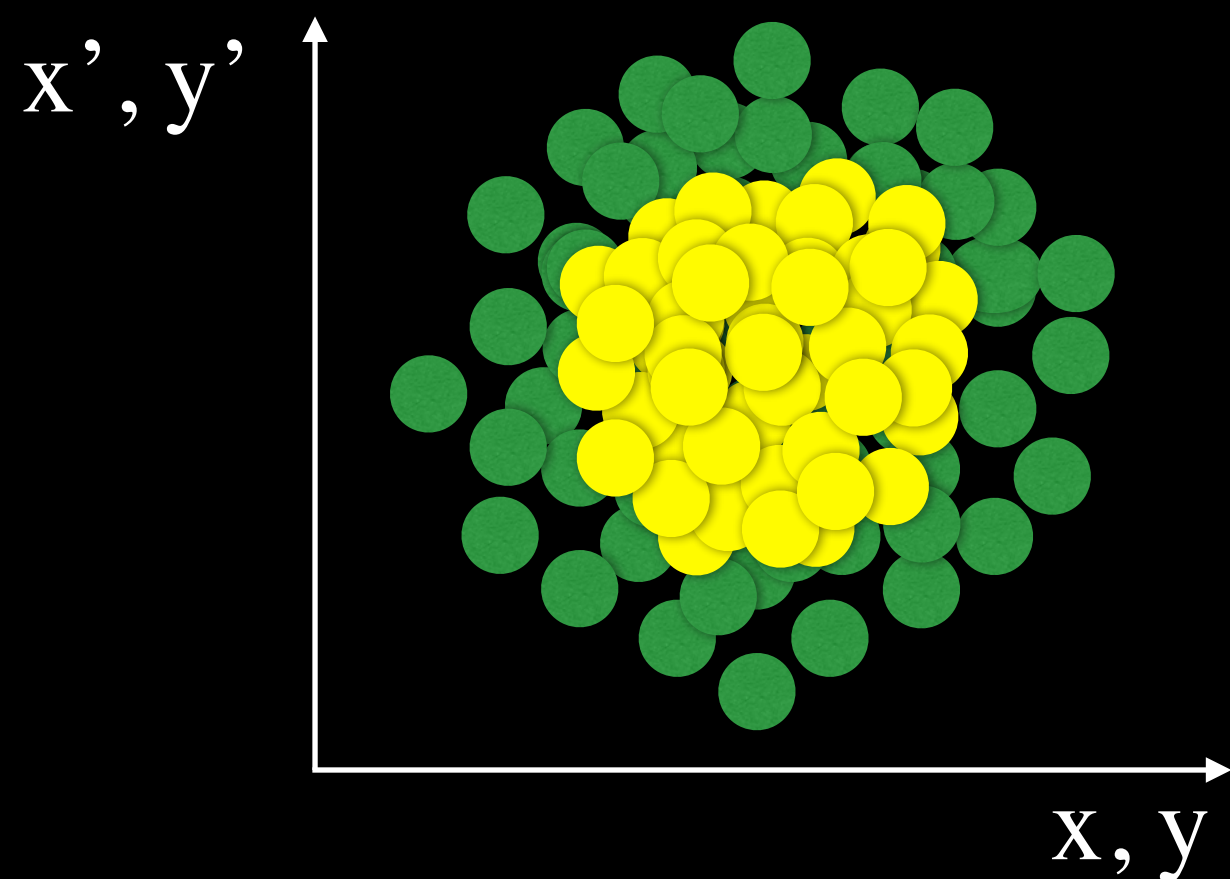


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**Problem:** when muons are produced they occupy a large transverse phase-space



In order for muons to fit within acceptance of downstream accelerator, phase-space needs to reduce

...but muons decay very fast  
so only viable technique is  
**ionisation cooling**

# Ionisation Cooling

# Ionisation Cooling

Muons pass through absorbers;  
momentum decreases in all directions

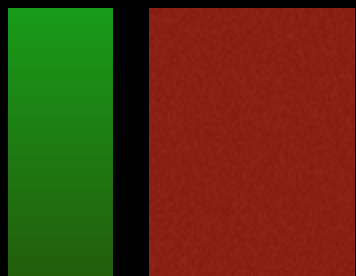


Absorber

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Muons pass through absorbers;  
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Muons pass through RF cavities; momentum  
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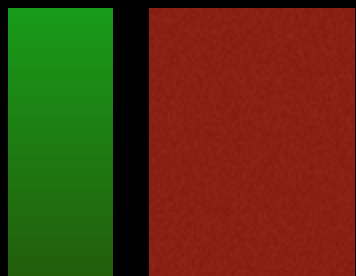
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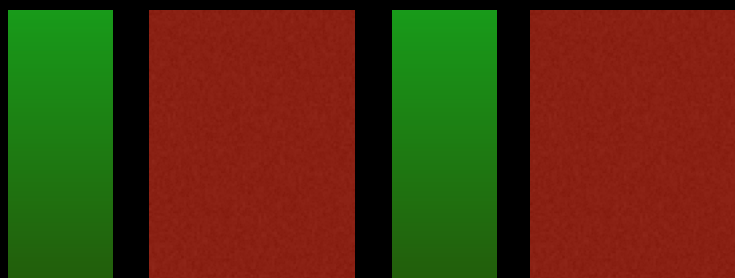
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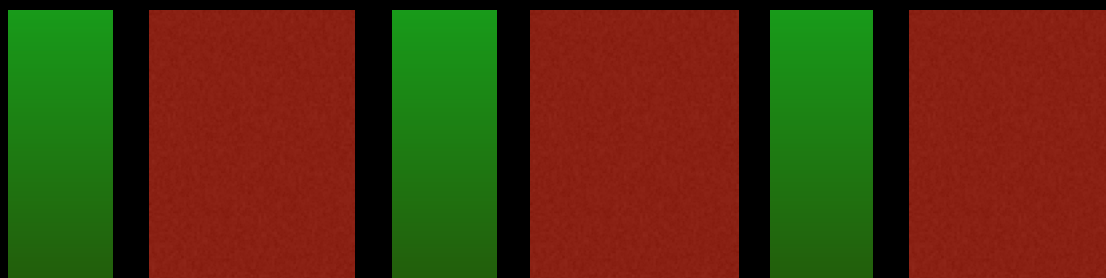
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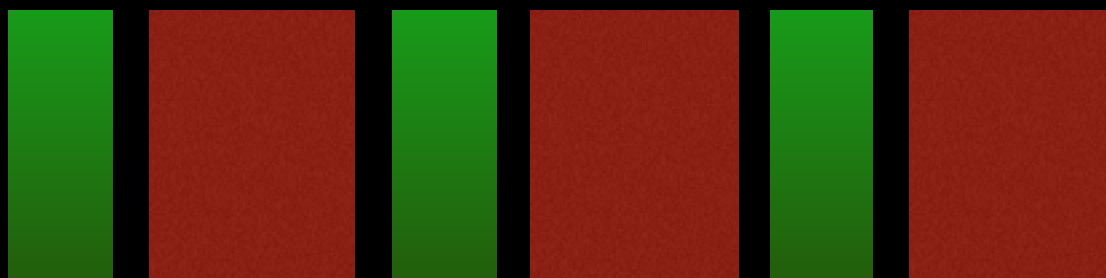
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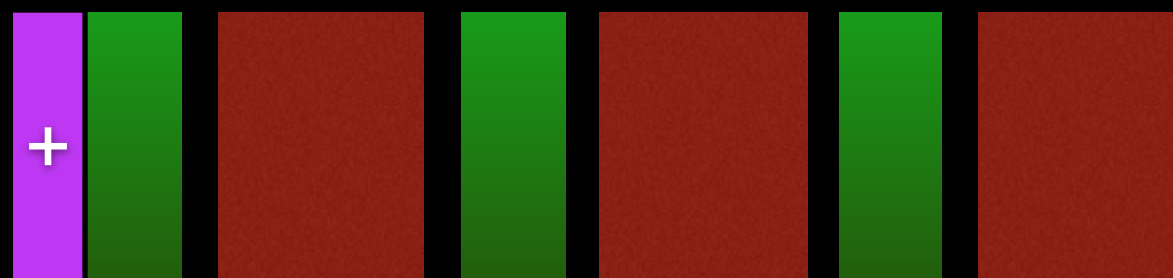
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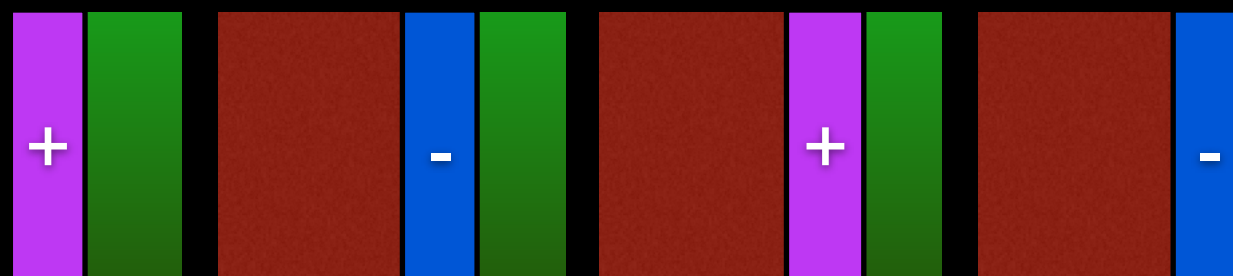
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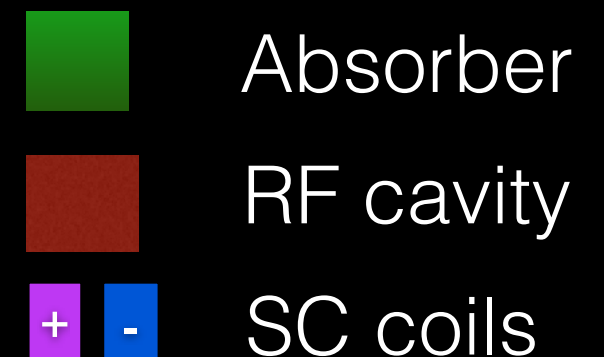
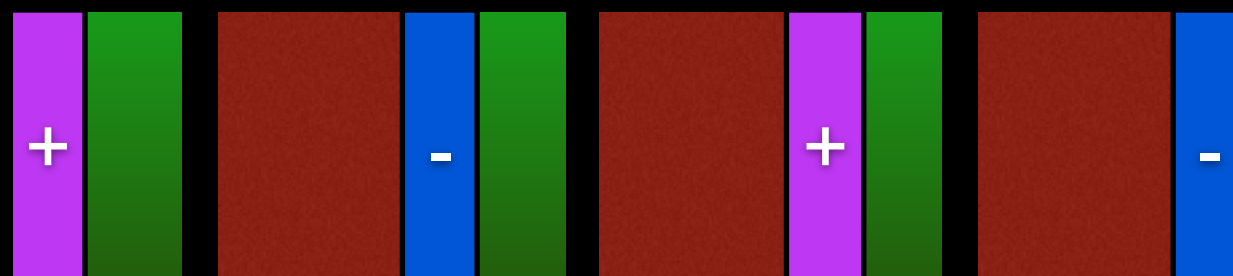
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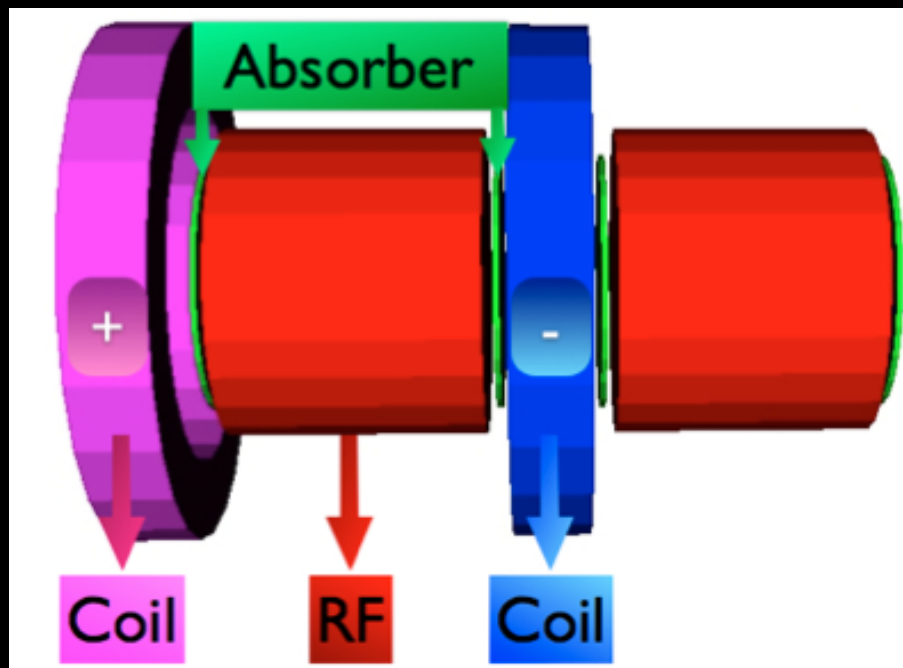
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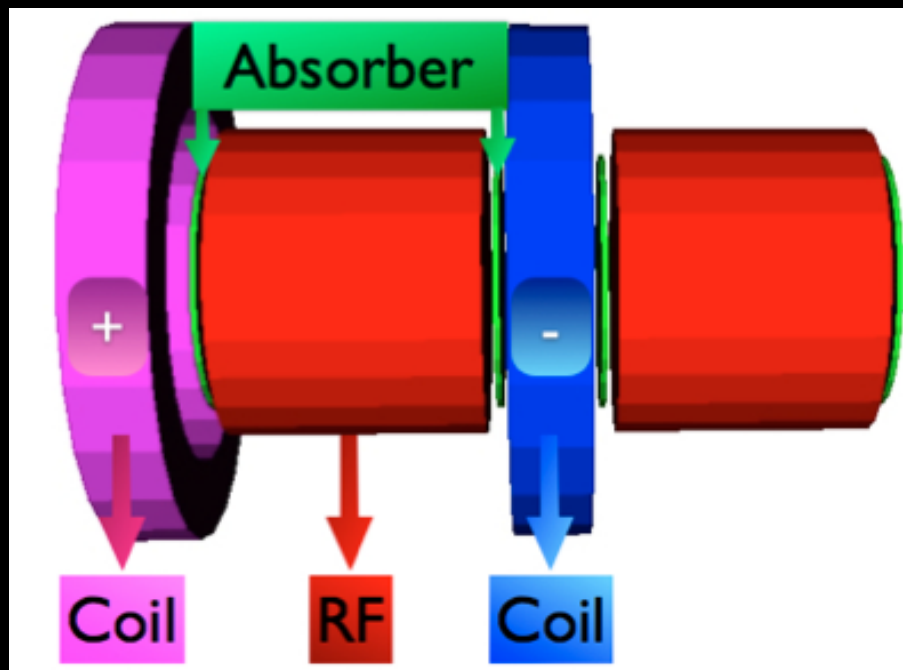


# Reference Cooling Lattice



# Reference Cooling Lattice

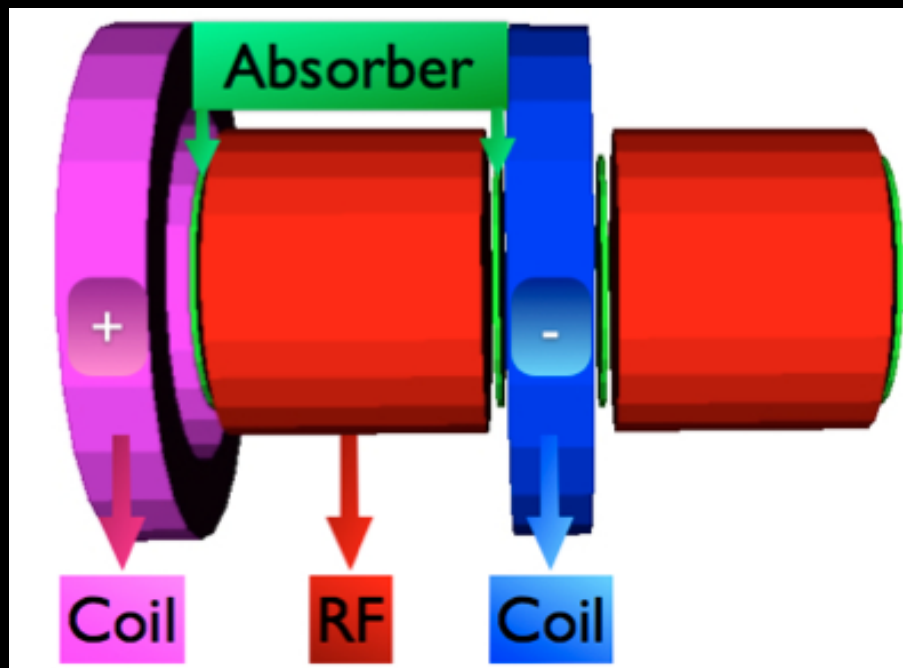
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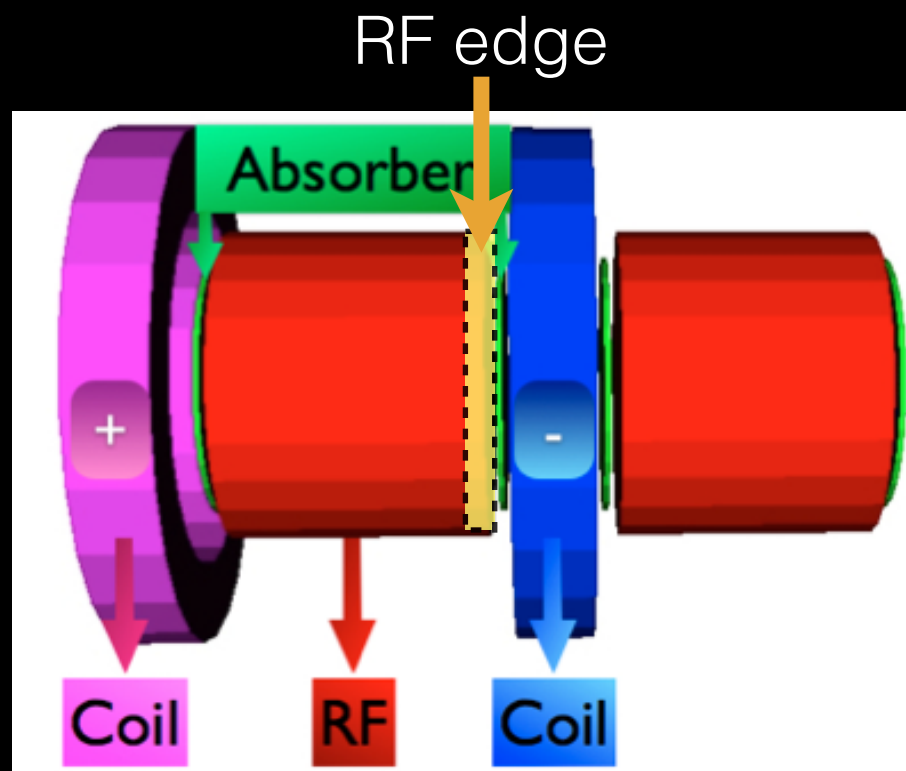




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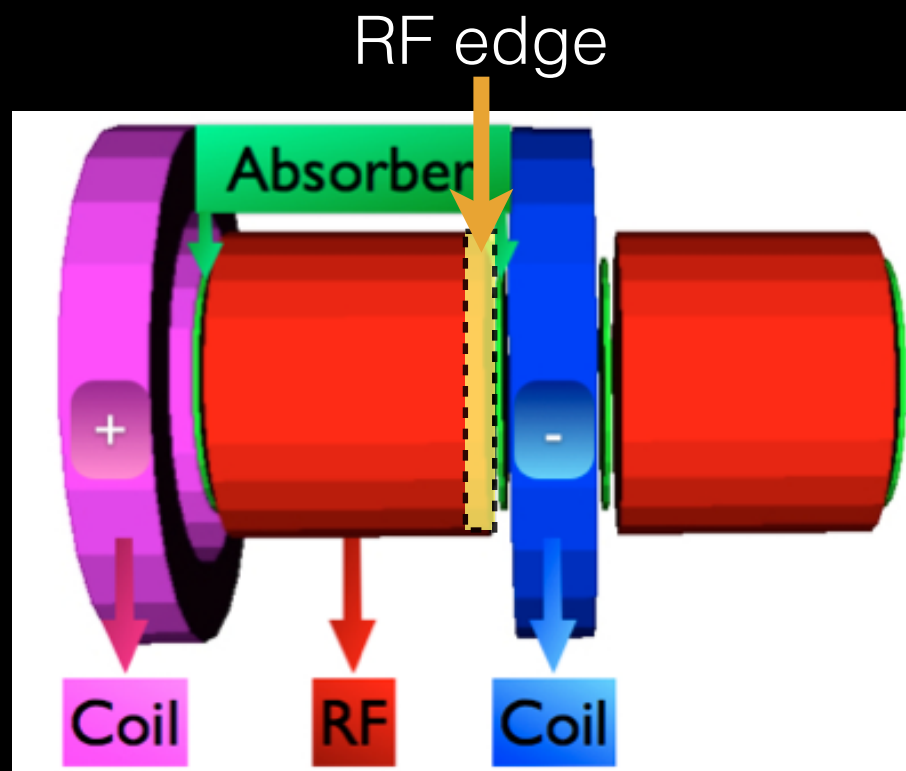


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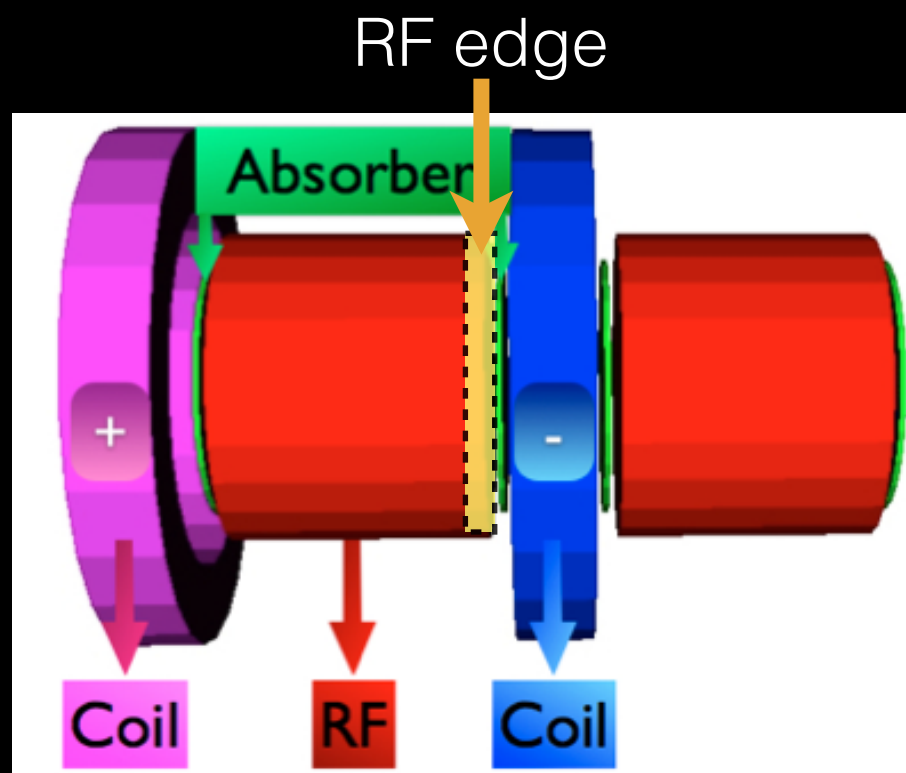


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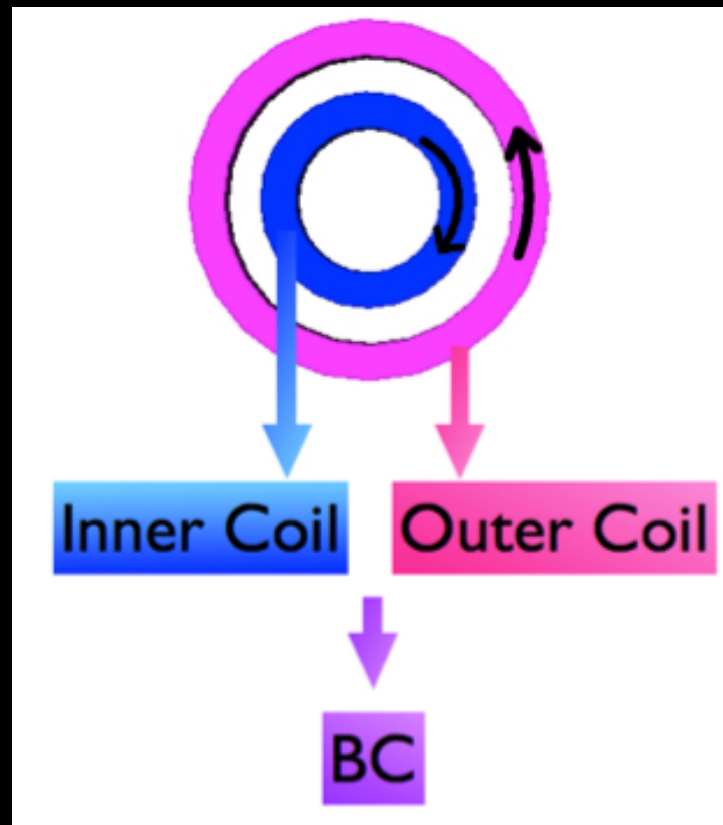
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Need to find alternative lattice that:  
a) reduces significantly magnetic field at RFs  
b) without compromising emittance reduction and muon transmission

# Bucked Coils Lattice

- **IDEA:** Proposed and designed a new lattice that uses a pair of co-axial and opposite polarity coils, called Bucked Coils (BC), rather than a single one

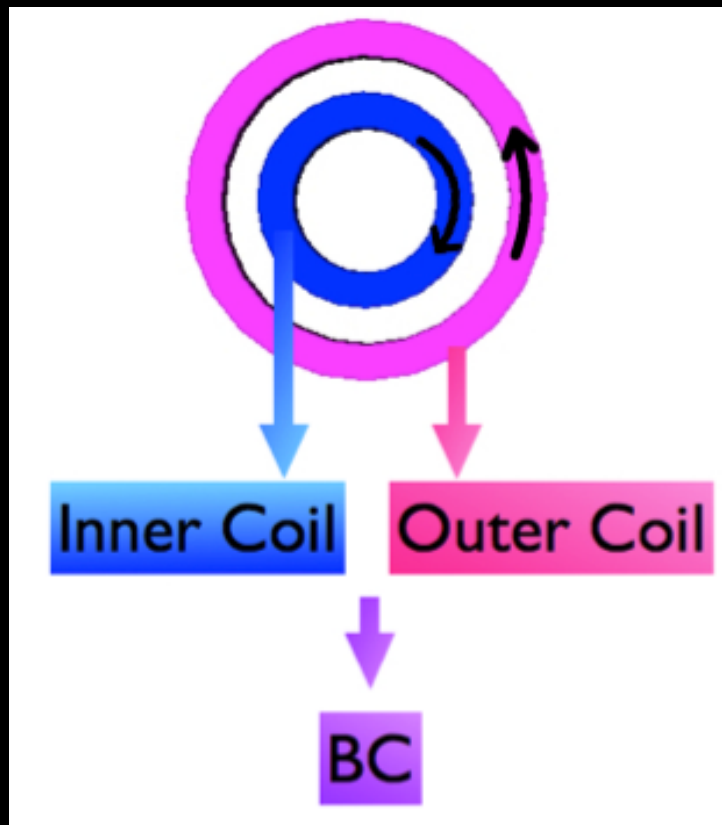


Pair of Bucked Coils (BC):  
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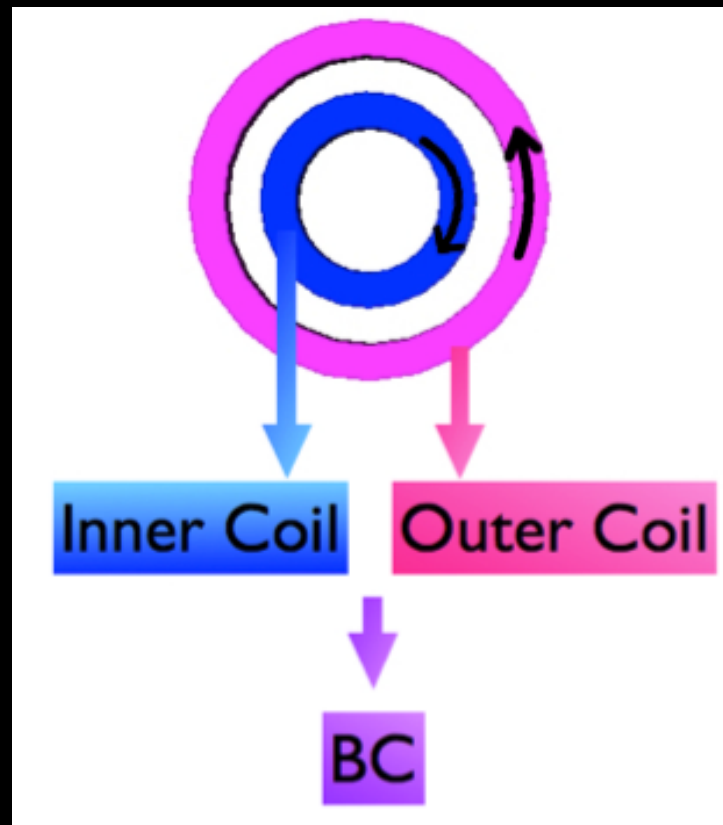


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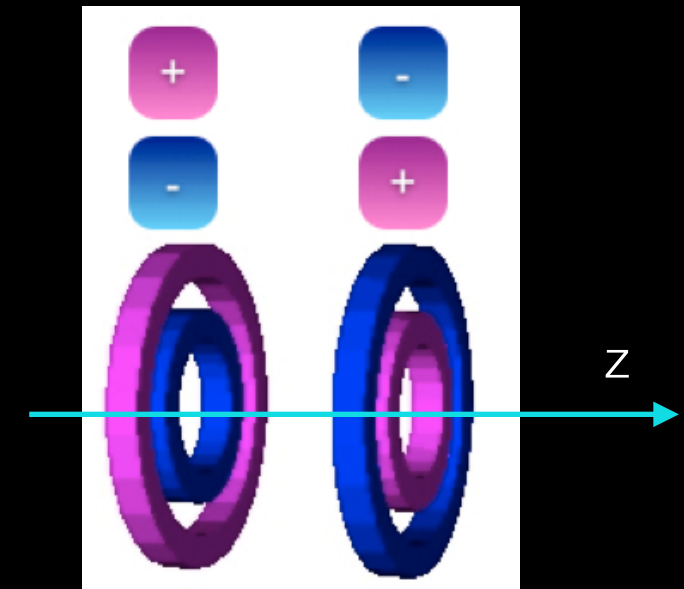
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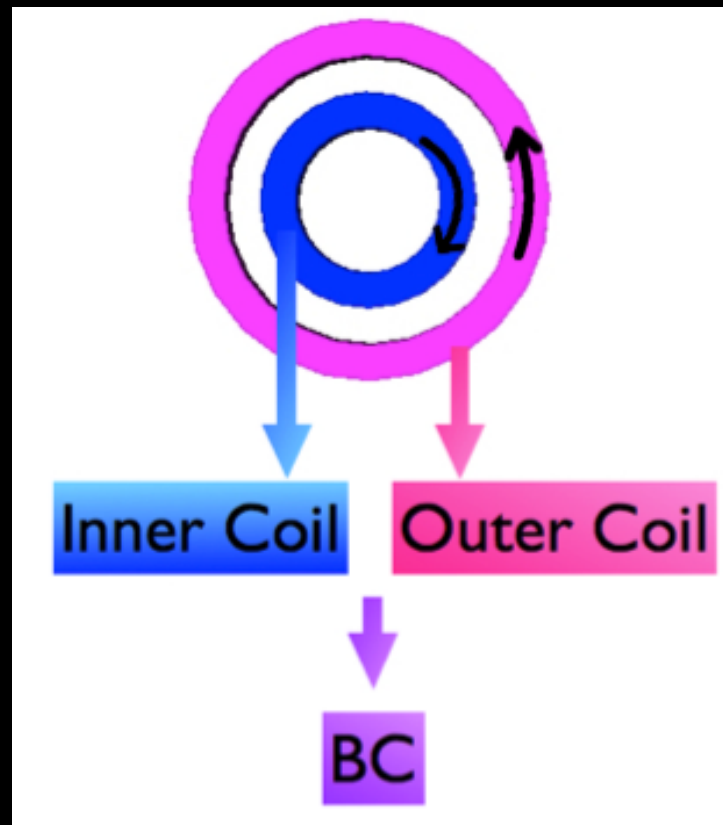


Polarity of BC  
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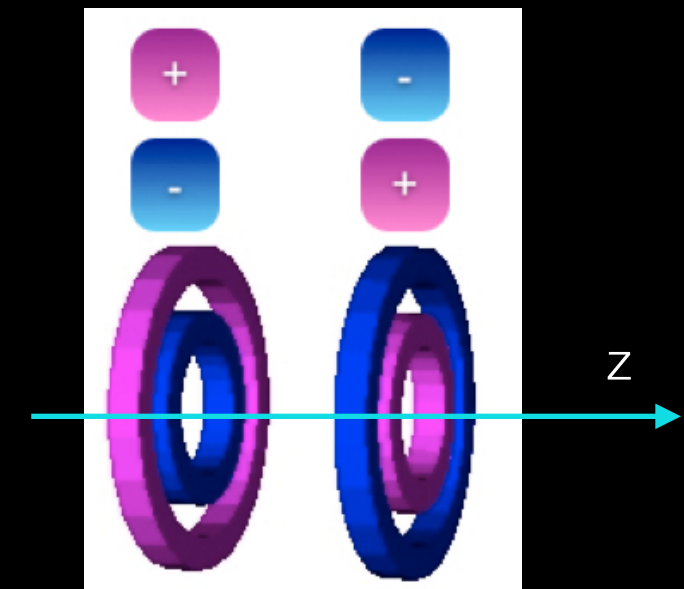
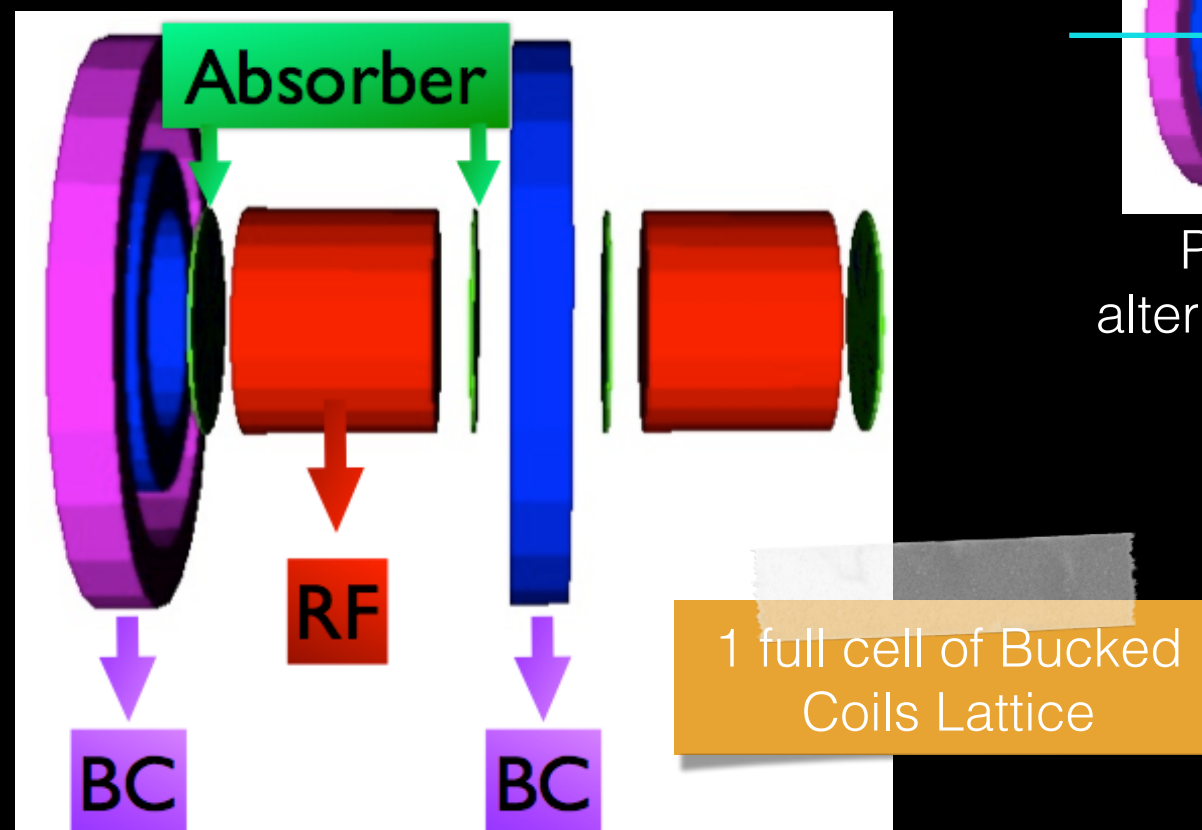
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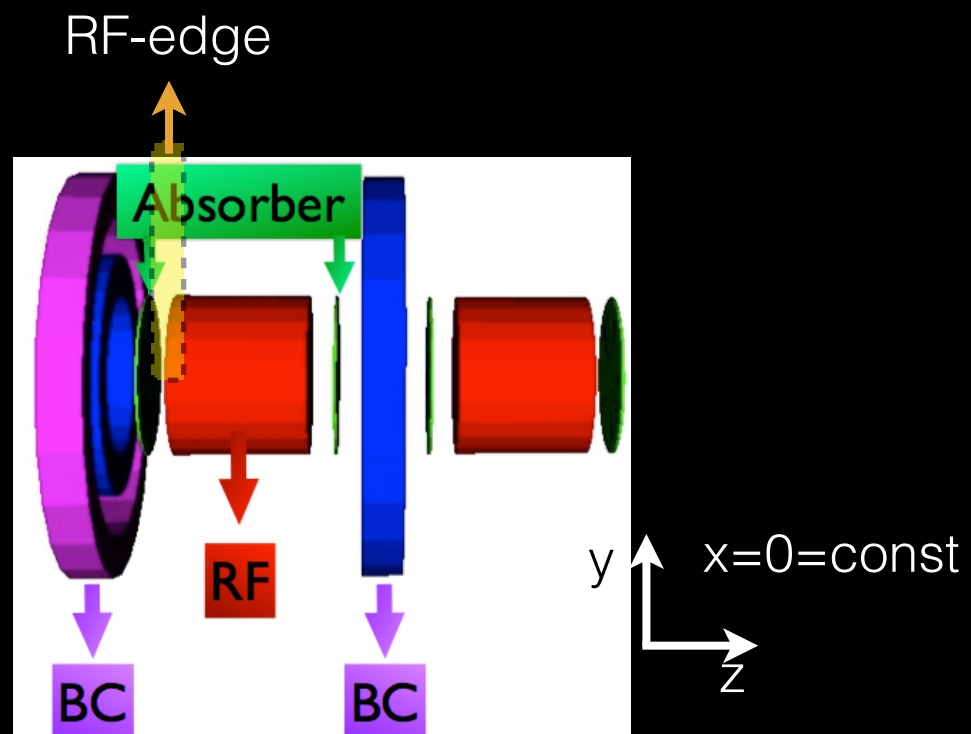


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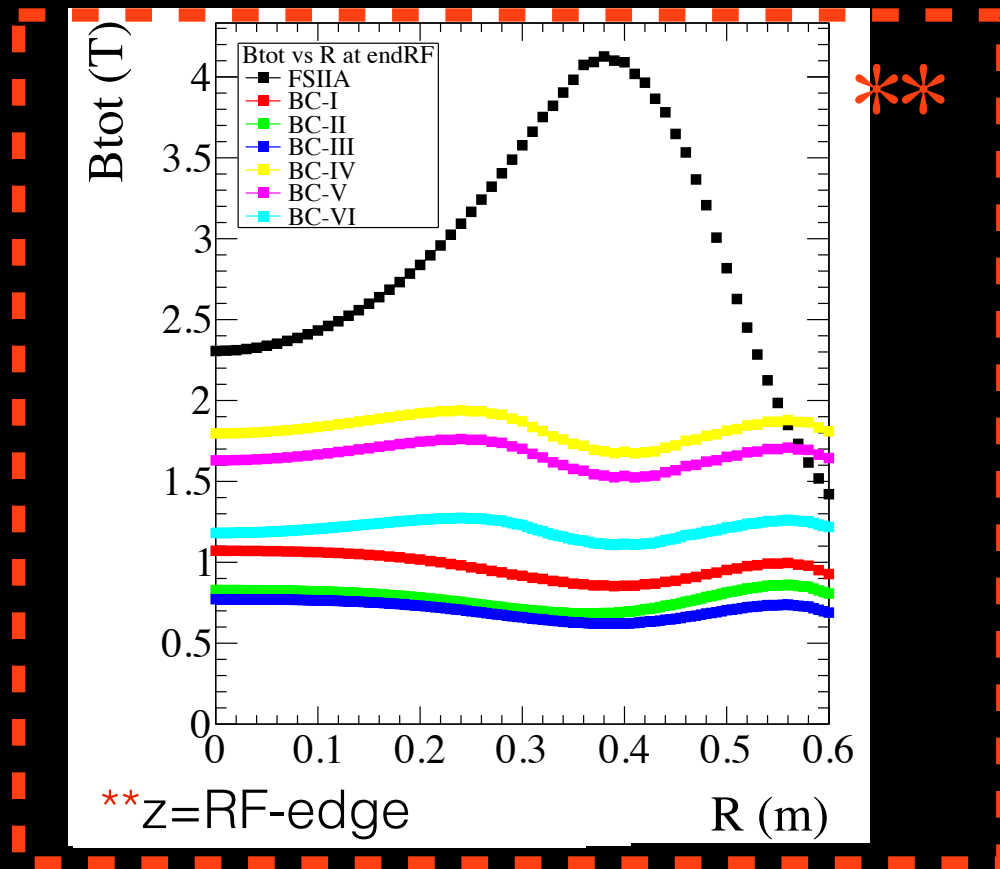
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# Magnetic field

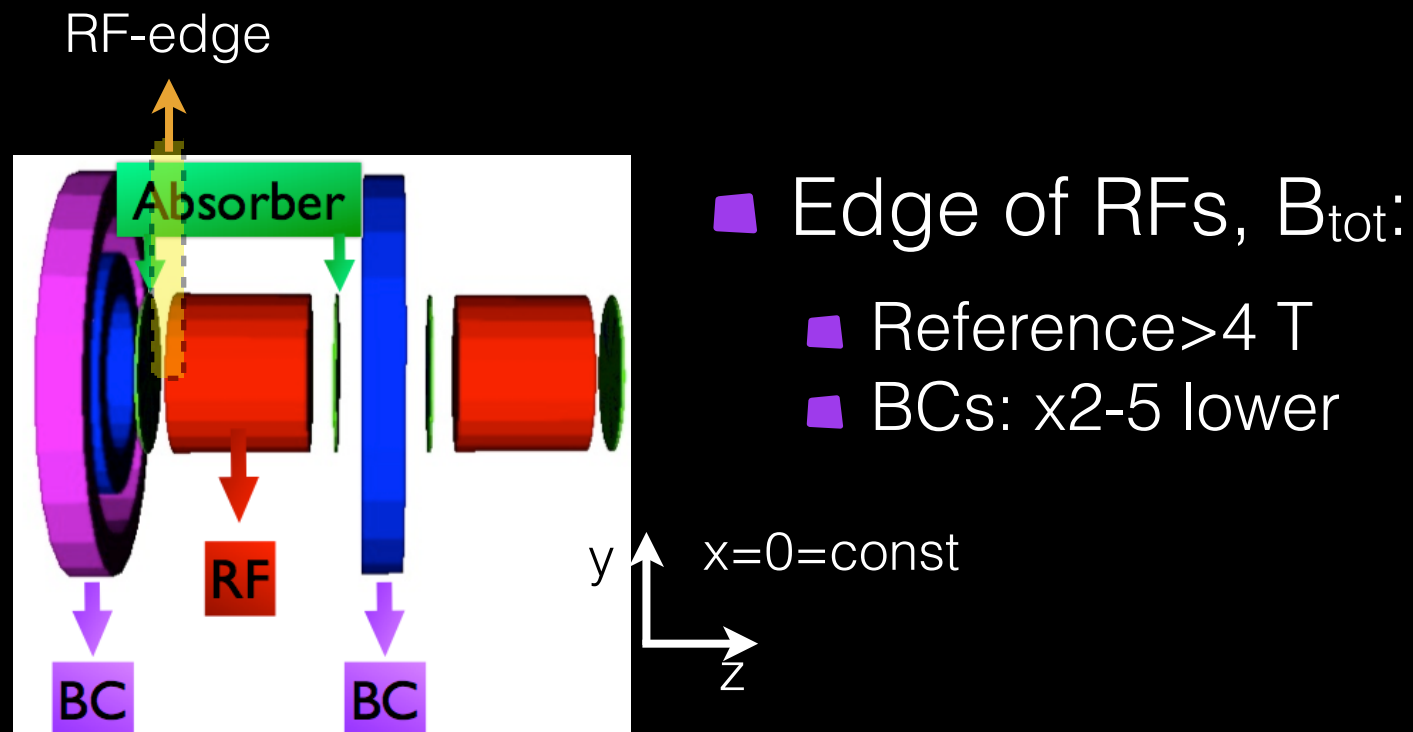




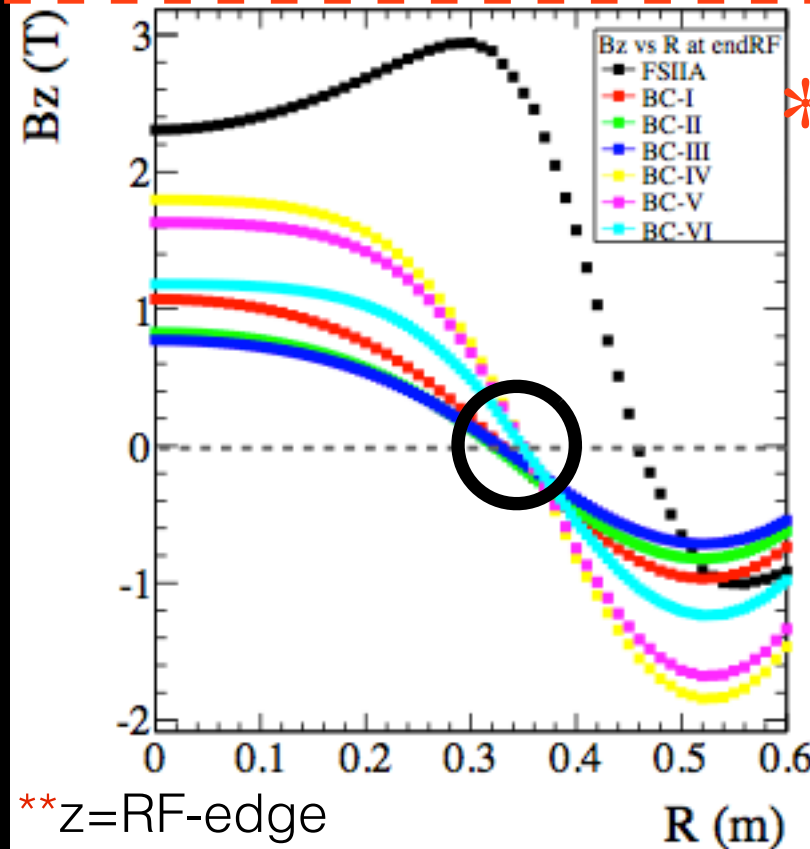
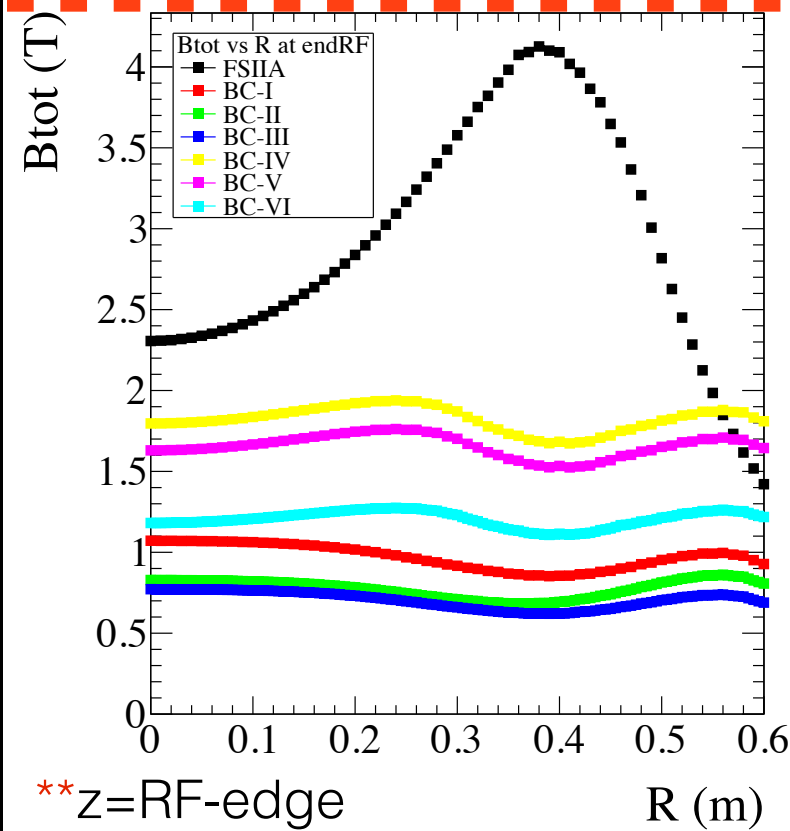
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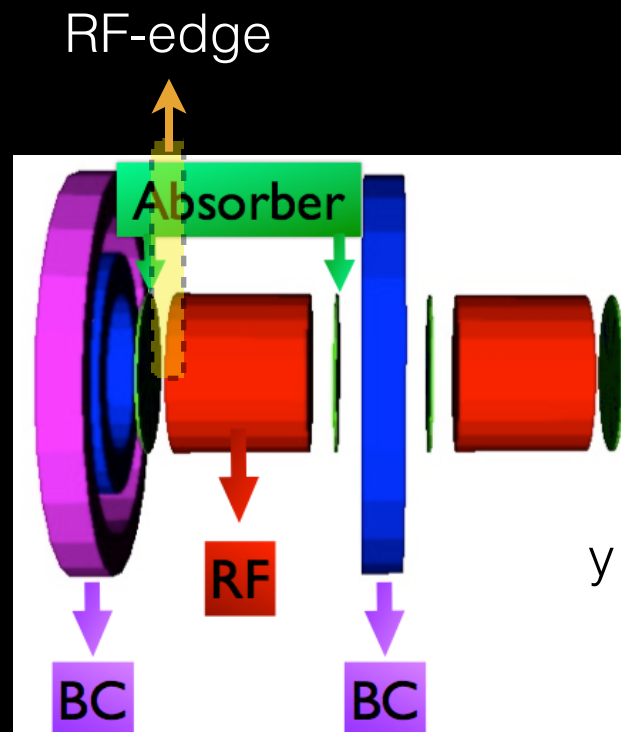
Reference  
BC versions



# Magnetic field



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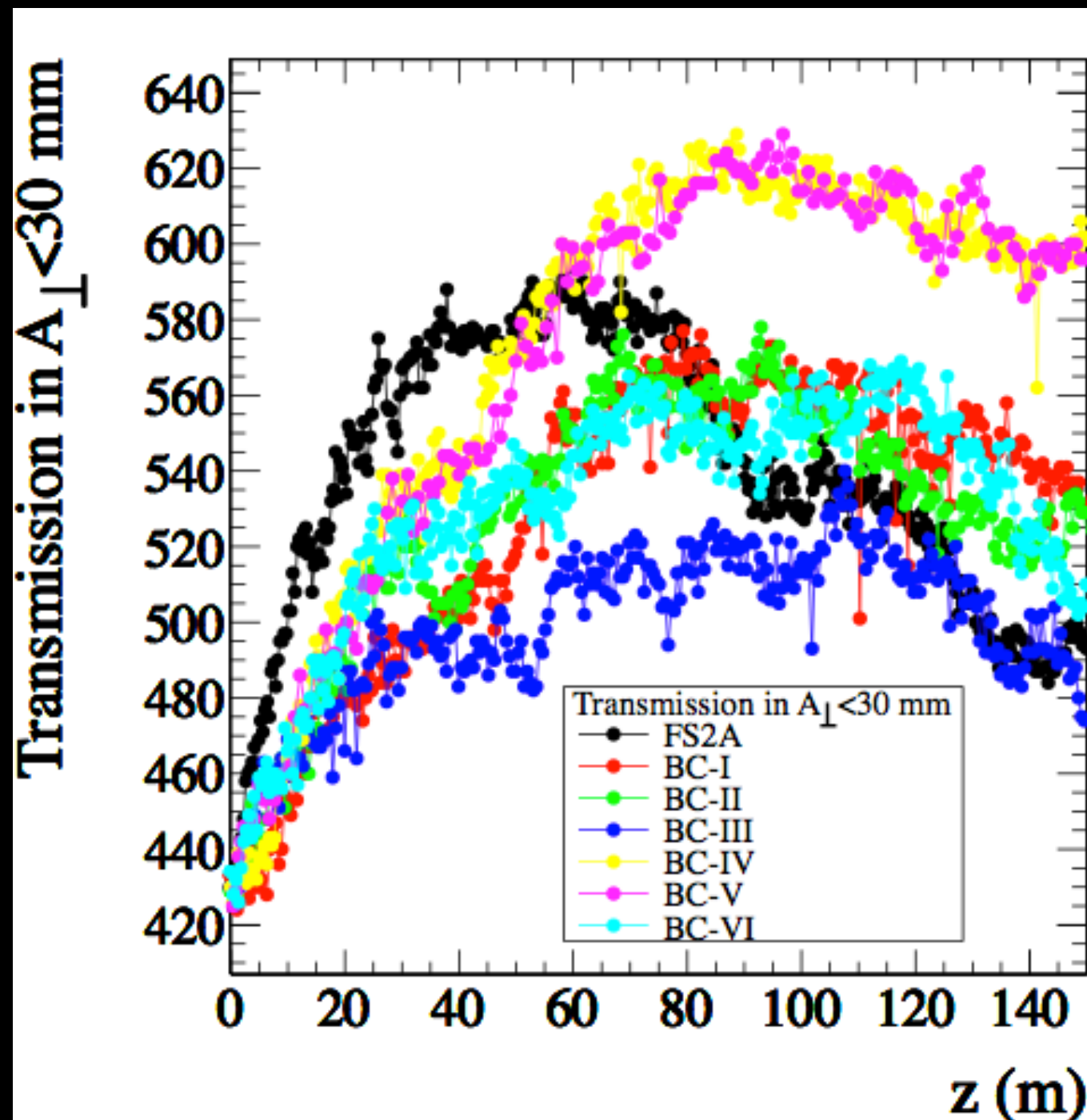


- Edge of RFs,  $B_{\text{tot}}$ :
  - Reference > 4 T
  - BCs: x2-5 lower

- Edge of RFs,  $B_z$  at  $R=35$  cm (RF iris):
  - Reference ~ 2.5 T
  - BCs: 0 T**

# Cooling efficiency

- BC: best transmission within 30 mm of  $A_T^*$

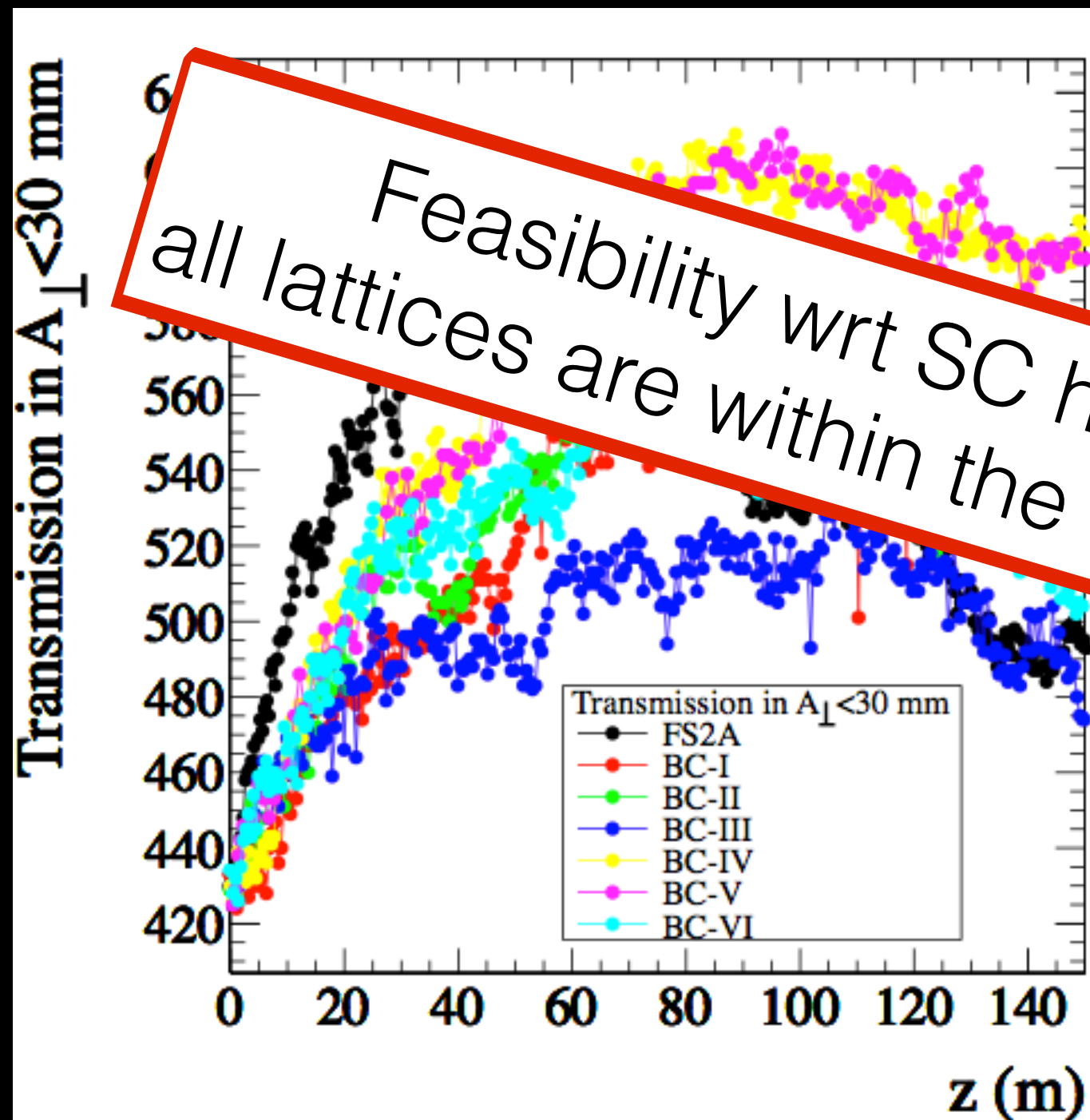


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\*Transverse acceptance of downstream accelerator system

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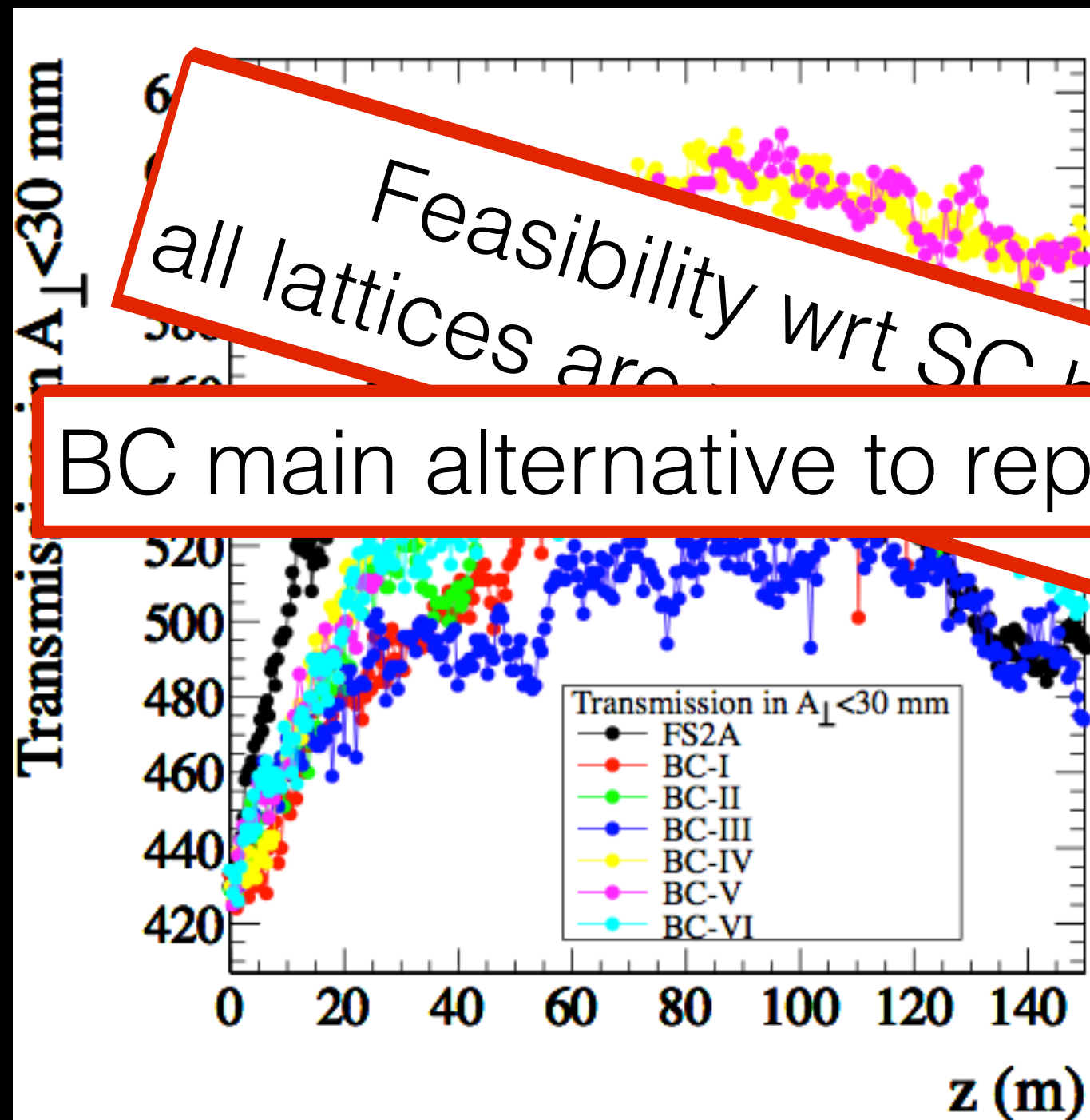


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# Cooling efficiency

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Reference  
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BC main alternative to replace the reference lattice

limits of SC operation

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# Cooling efficiency

- **BC: best transmission within 30 mm of  $A_T^*$**



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Results of SC operation

Jinst

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## Bucked Coils lattice: a novel ionisation cooling lattice for the Neutrino Factory

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<sup>b</sup>STFC, RAL, ISIS,  
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2012 JINST

\*Transverse  
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● HP-PS, optics correction and collimation, LAGUNA-LBNO

\* Fast Extraction machine development, SPS

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# LAGUNA-LBNO\*

Main goal: feasibility study of new European research infrastructure able to host a deep ( $\sim 1,5$  km) underground neutrino detector (mass:  $\sim 10^5 - 10^6$  tons) for fundamental research in particle and astroparticle physics

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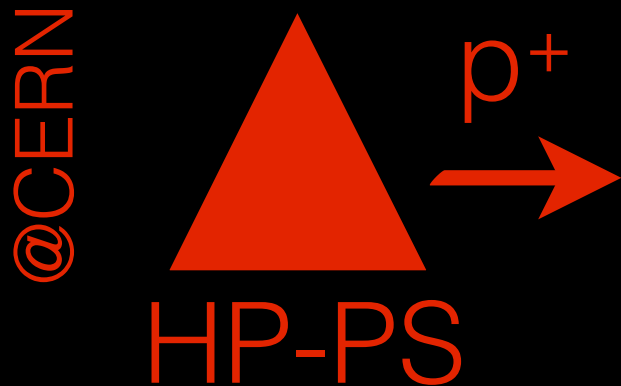
LAGUNA-LBNO will study matter-antimatter asymmetry using neutrinos produced at CERN [6]

\*Large Apparatus Studying Grand Unification, Neutrino Astrophysics and Long Baseline Neutrino Oscillations

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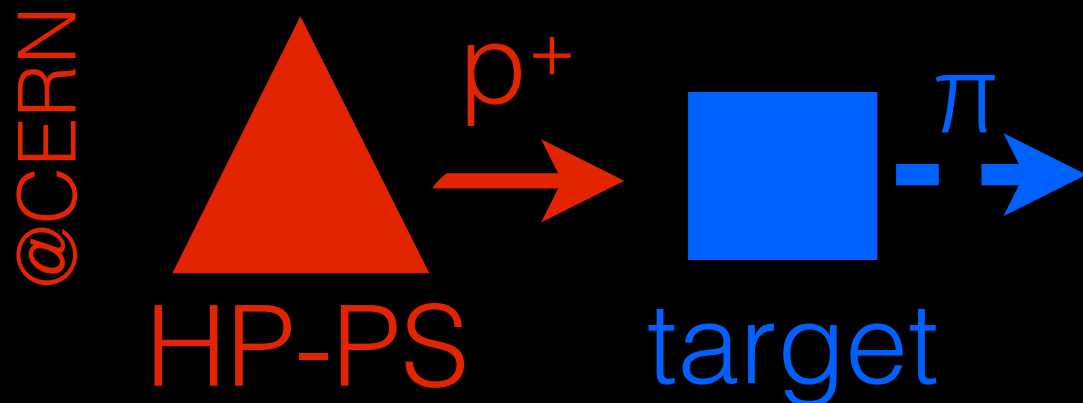


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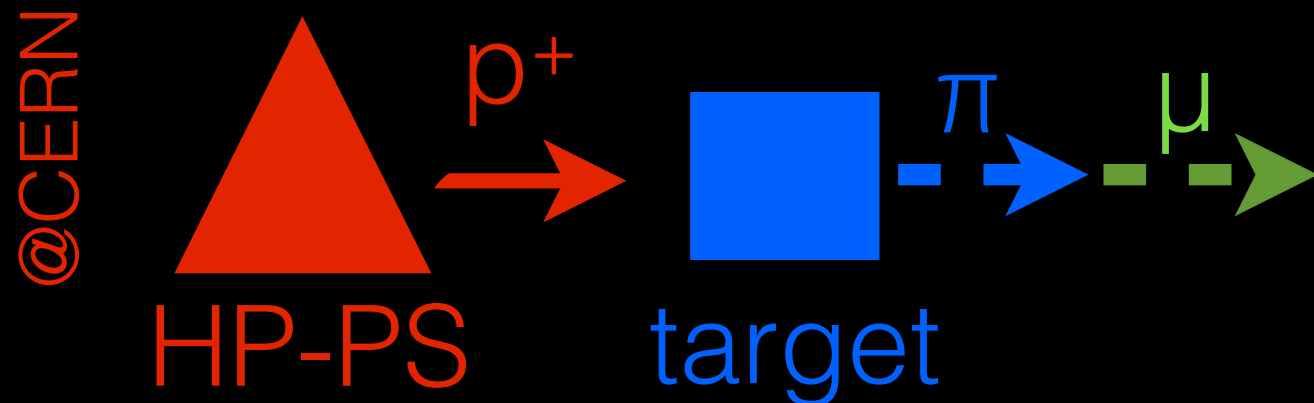


\*Large Apparatus Studying Grand Unification, Neutrino Astrophysics and Long Baseline Neutrino Oscillations

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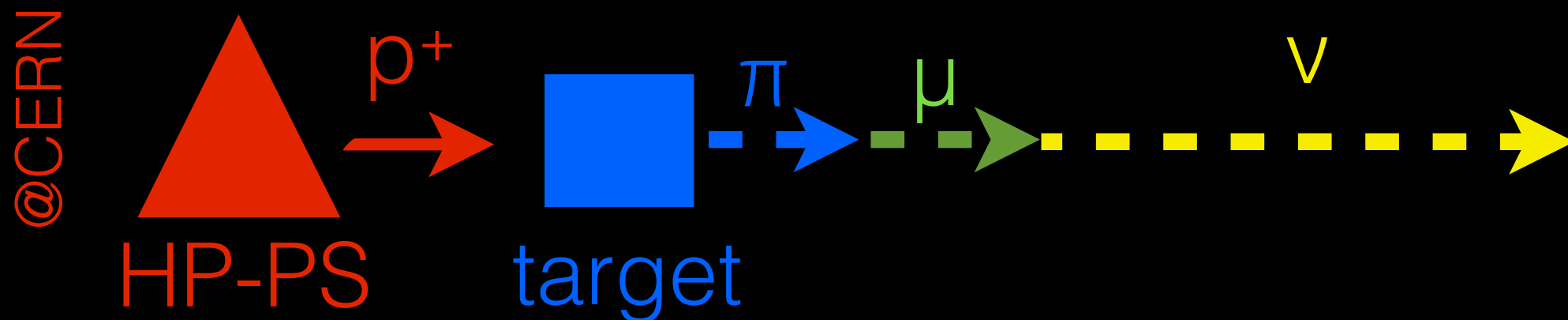
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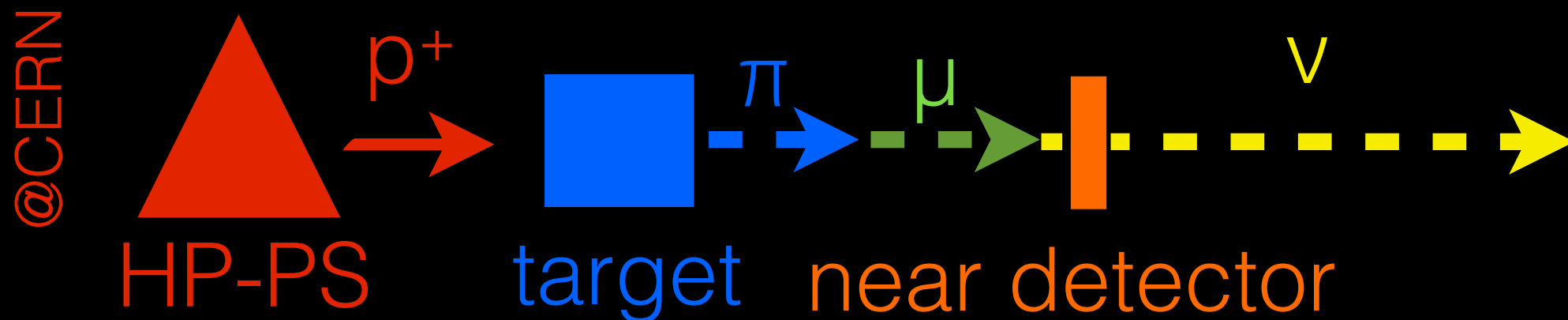




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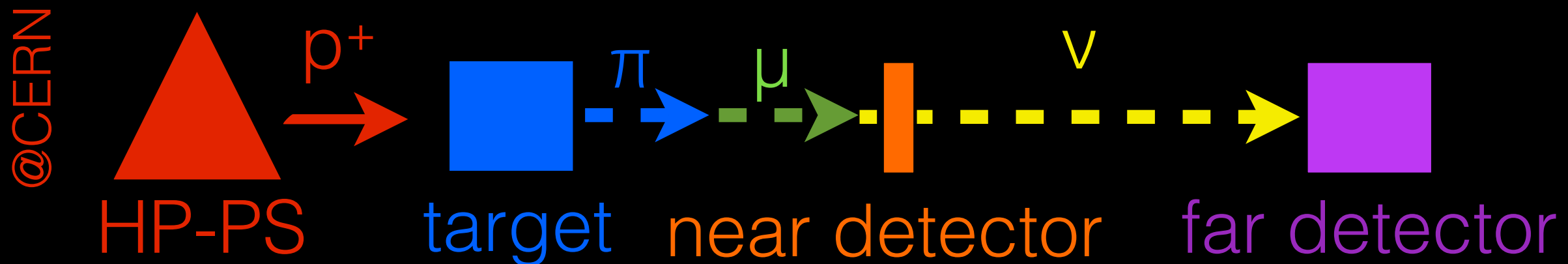


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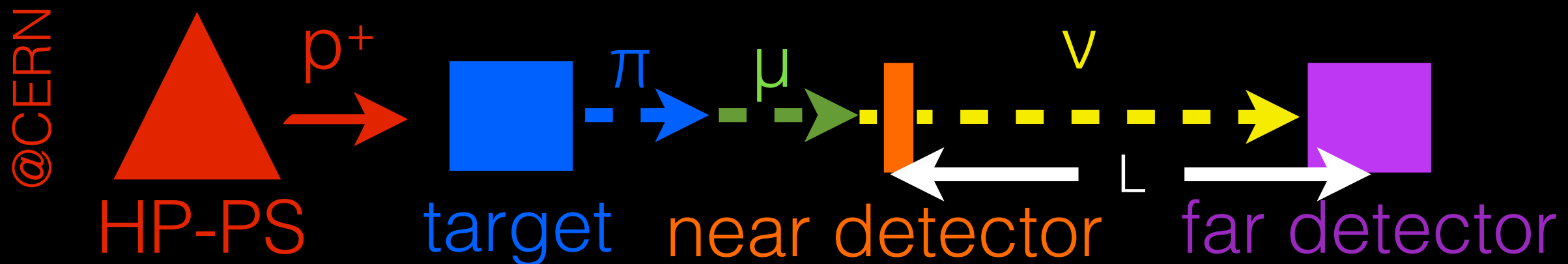
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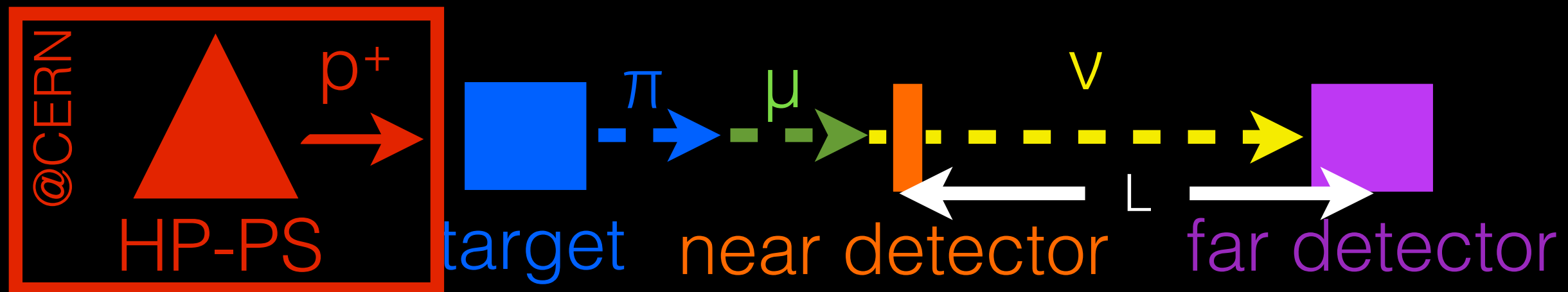


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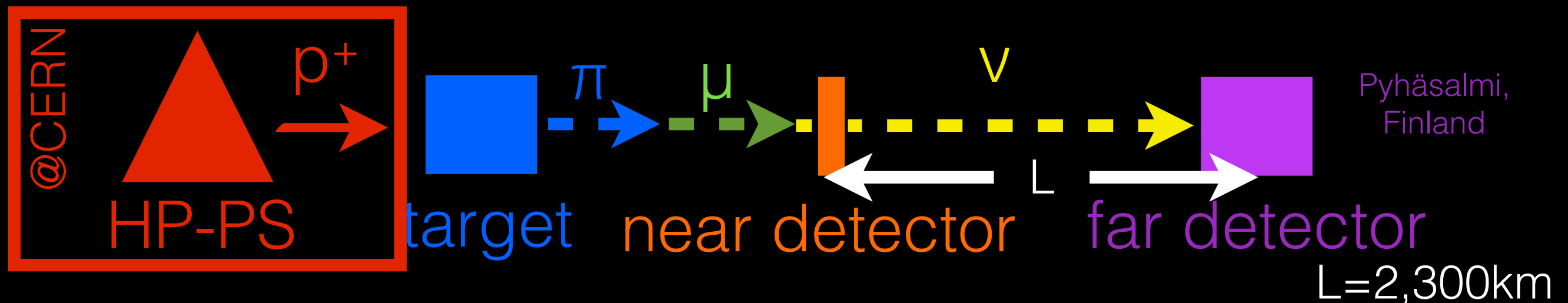
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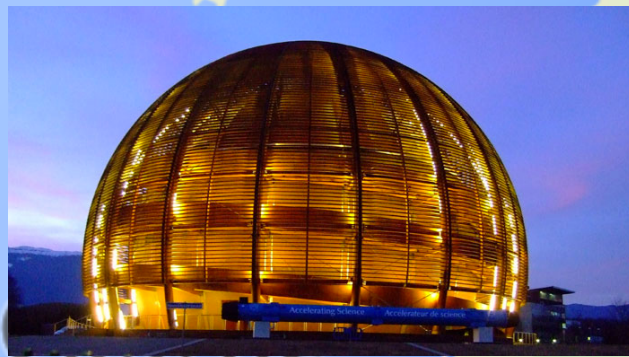




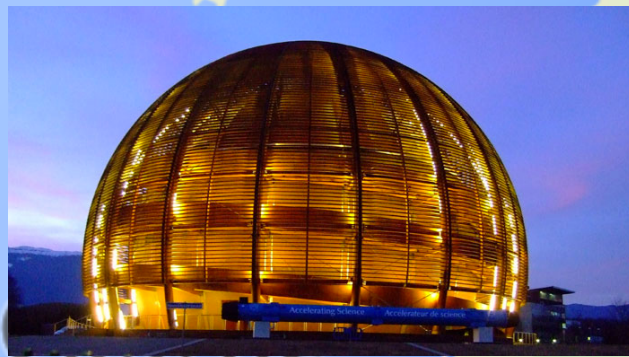
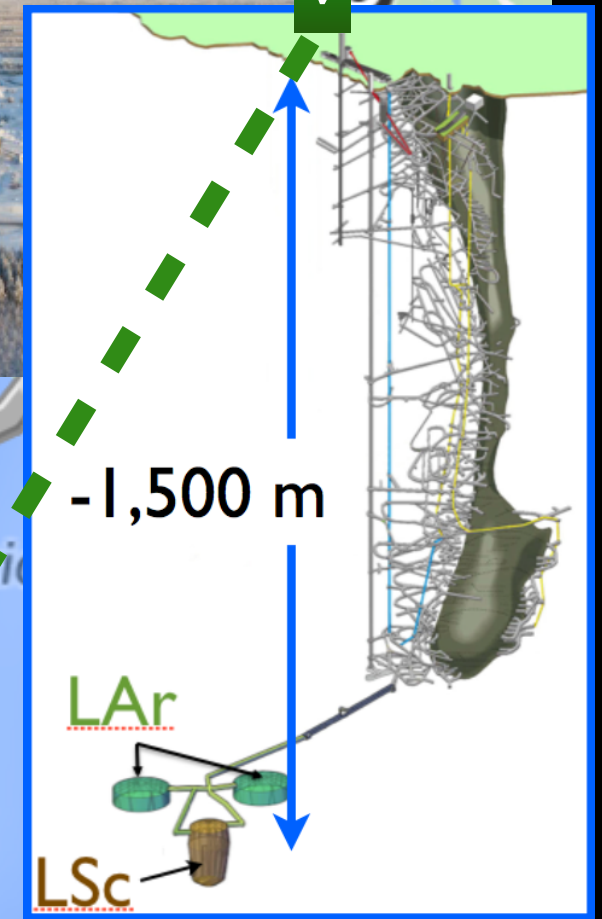




V

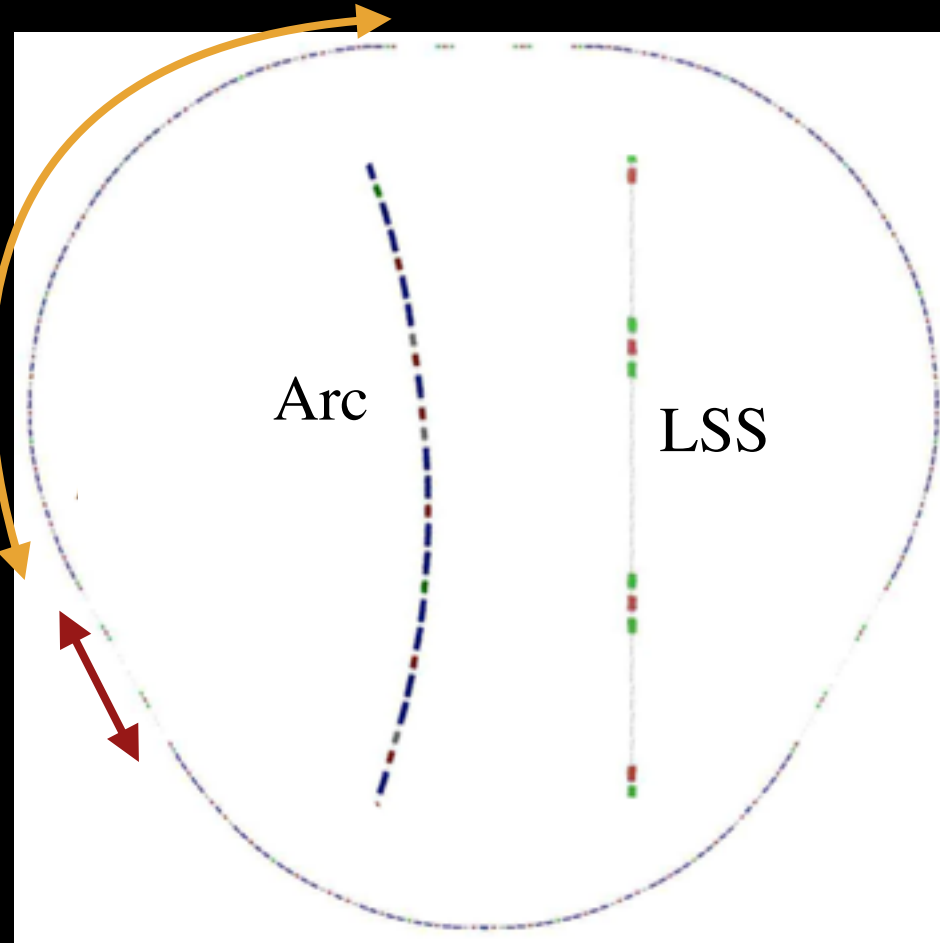






# HP-PS

- 3-fold symmetry
  - 3 Long Straight Sections (LSS):
    - a) injection/extraction
    - b) collimation
    - c) RF cavities
  - Protons final energy: 50 GeV
- Arc
- LSS



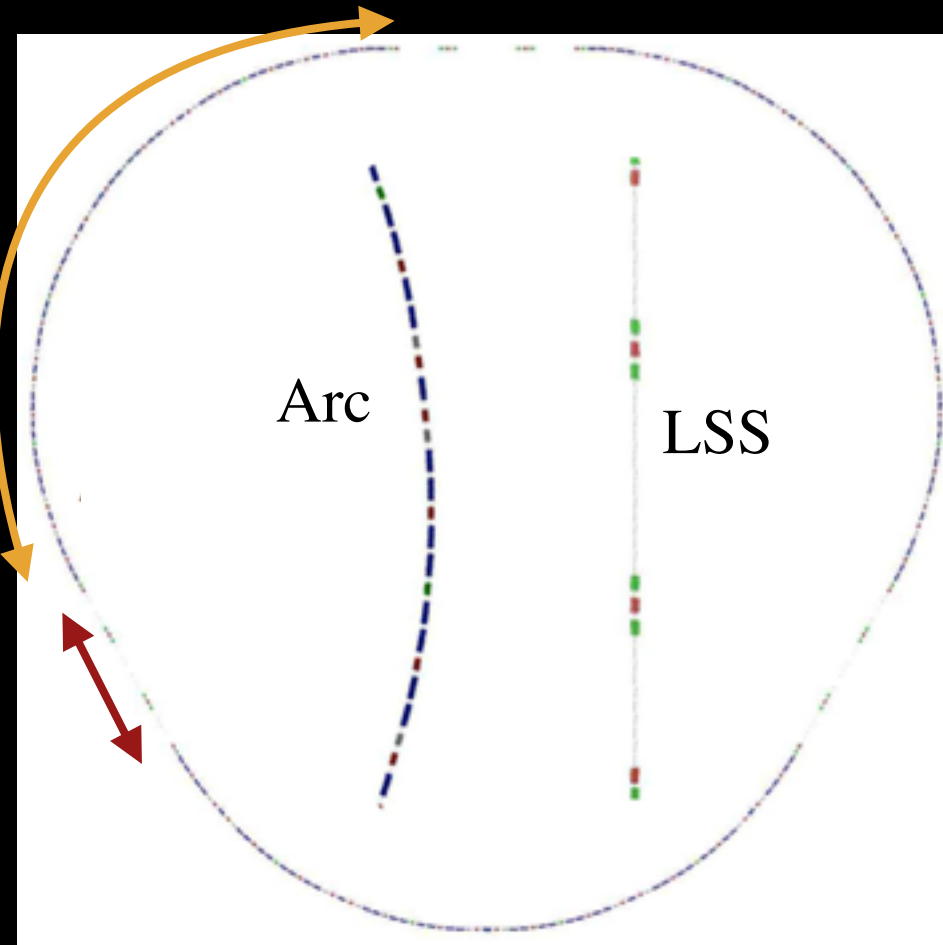
- Bending elements
- Focusing quads
- Defocusing quads

$L_{\text{Total}} = 1256 \text{ m}$   
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HP-PS: High Power Proton Synchrotron

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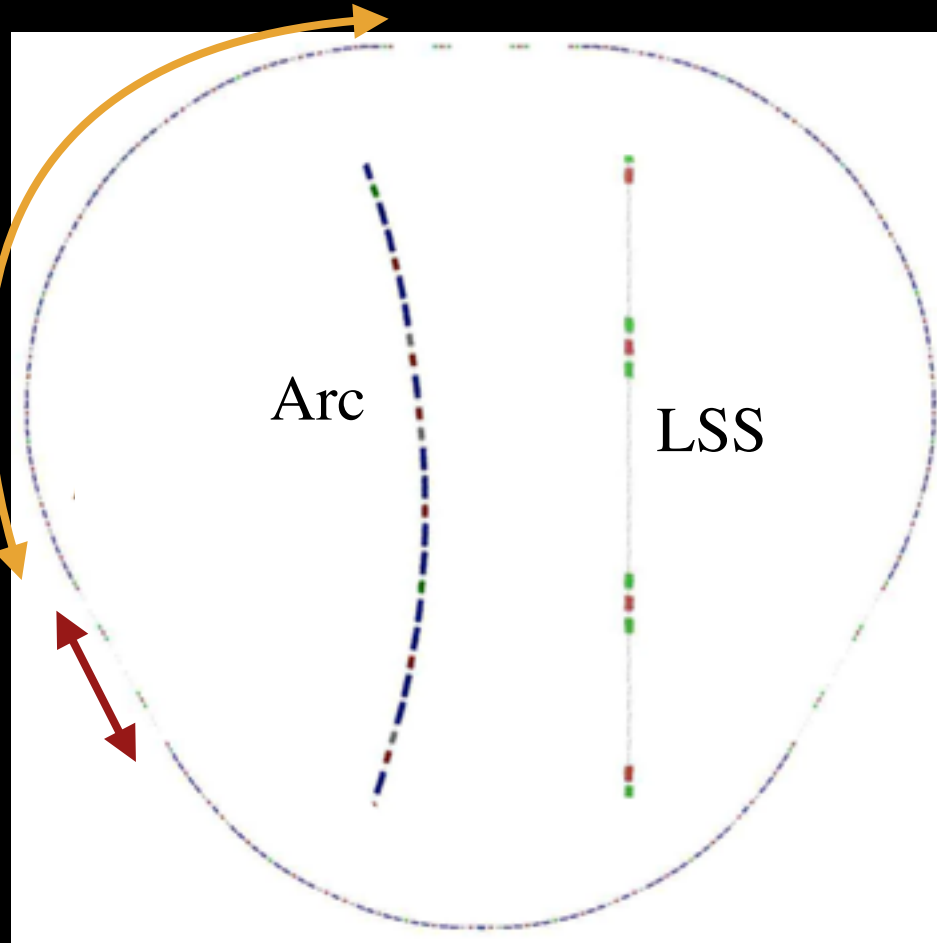
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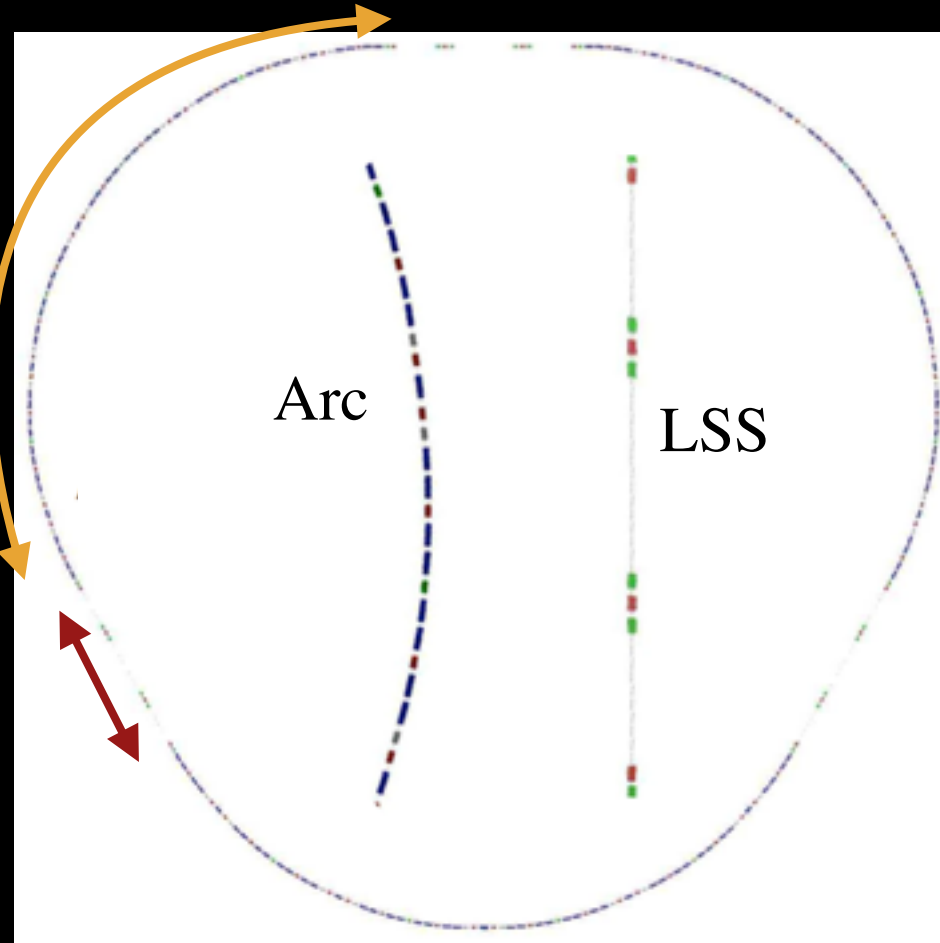
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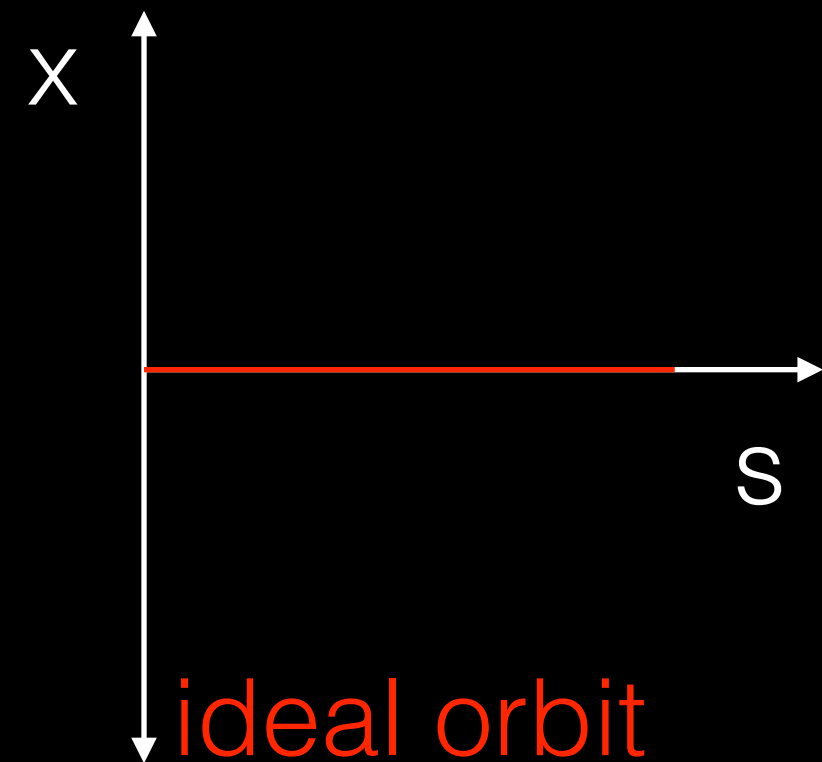
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- In ideal machine orbit is just a straight line



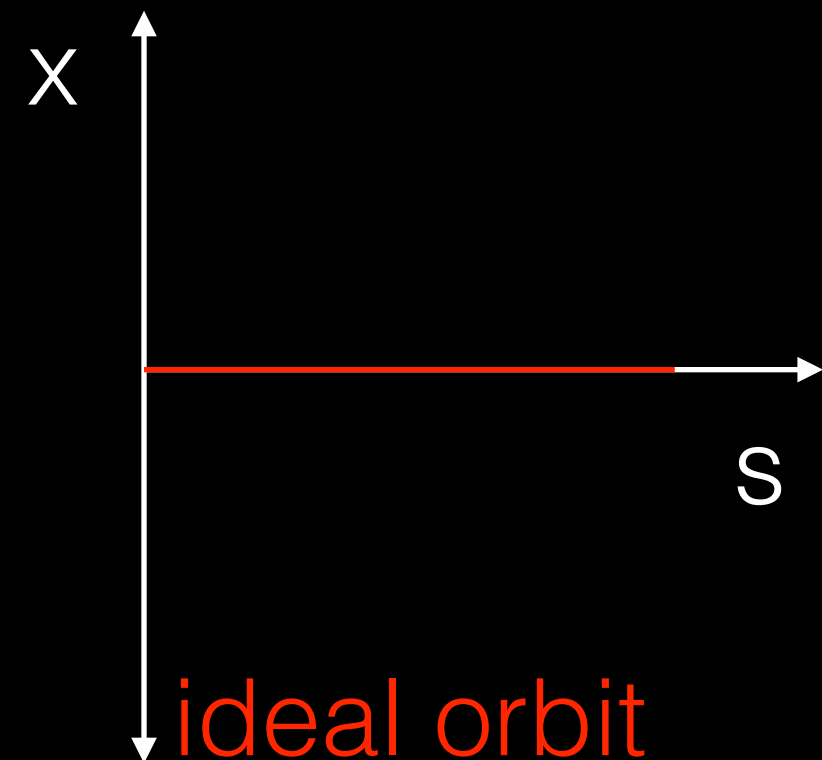
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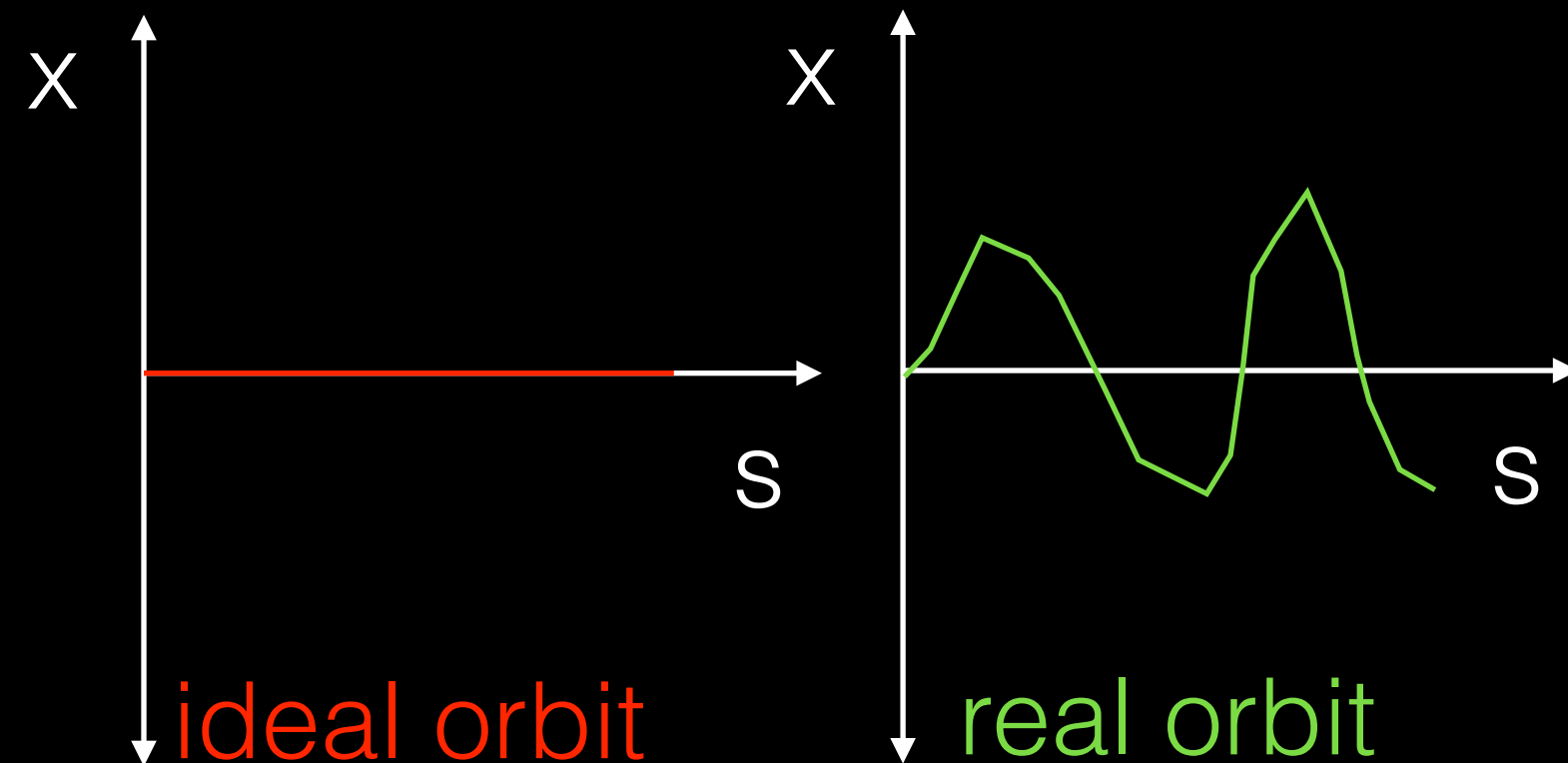
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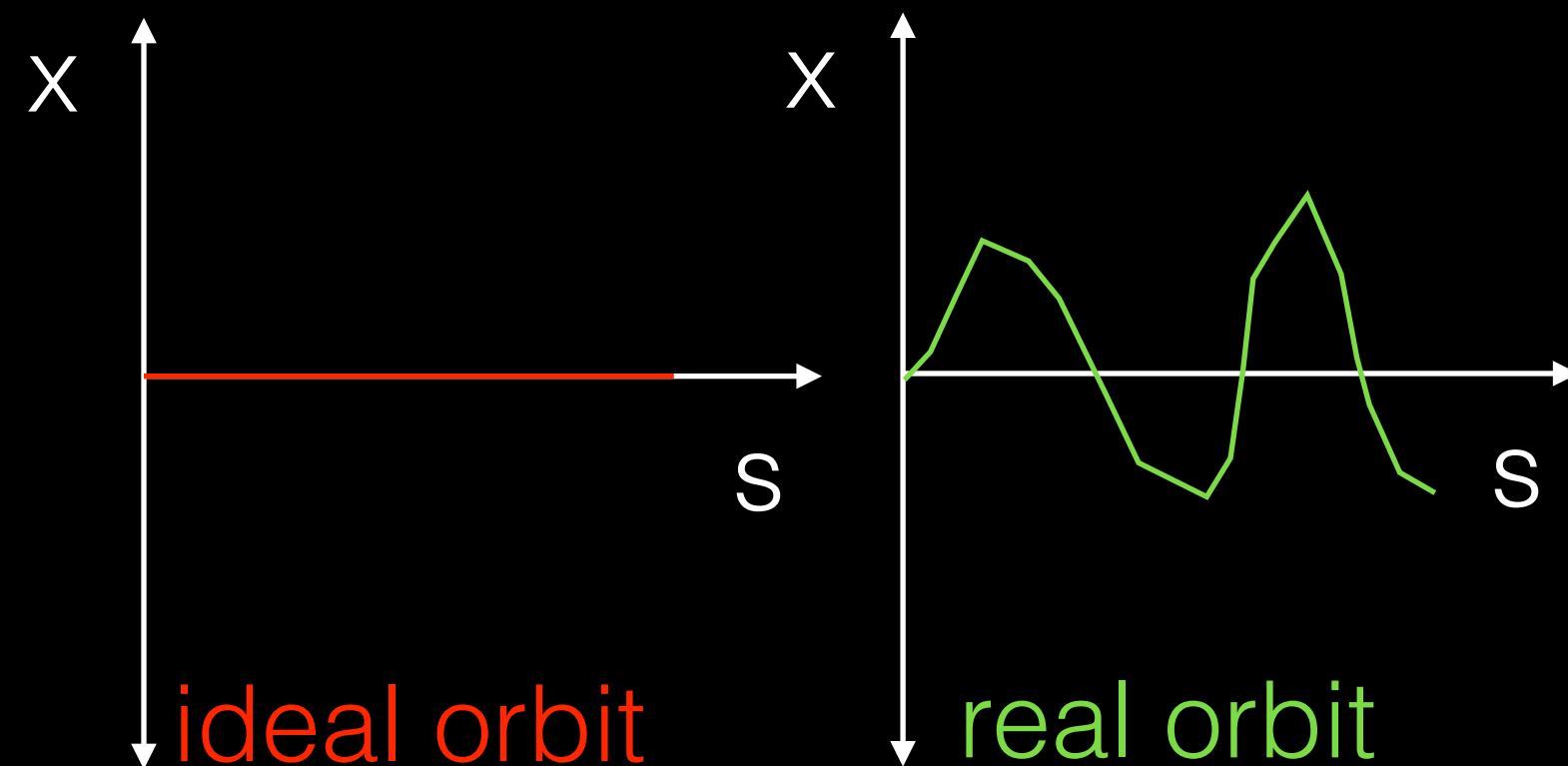
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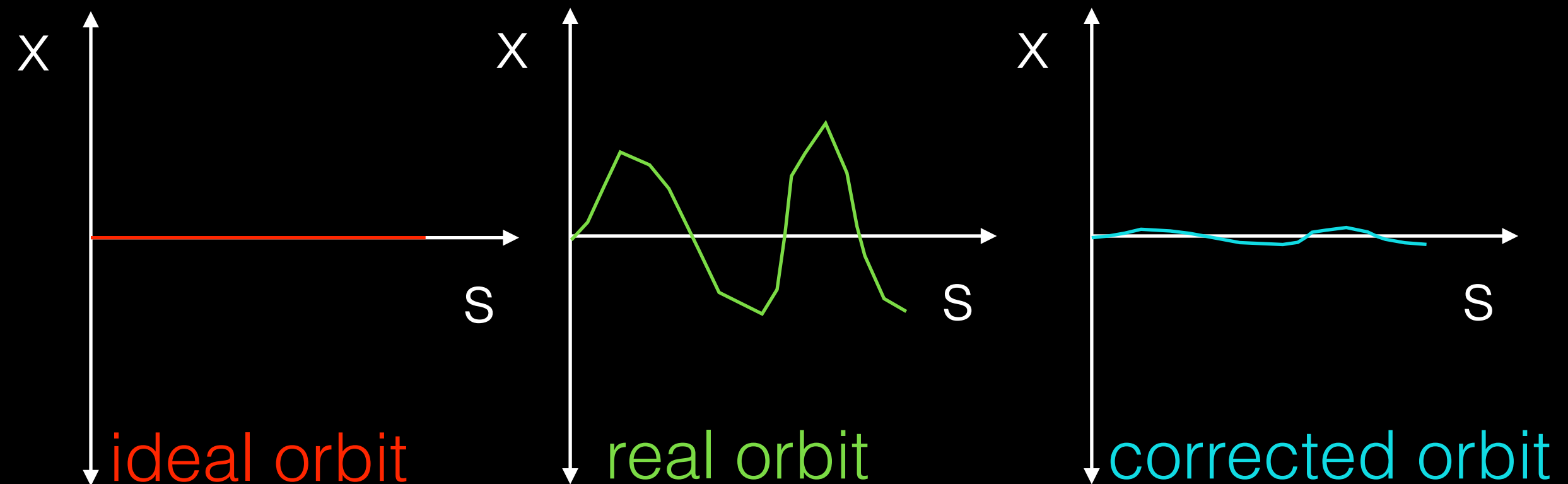
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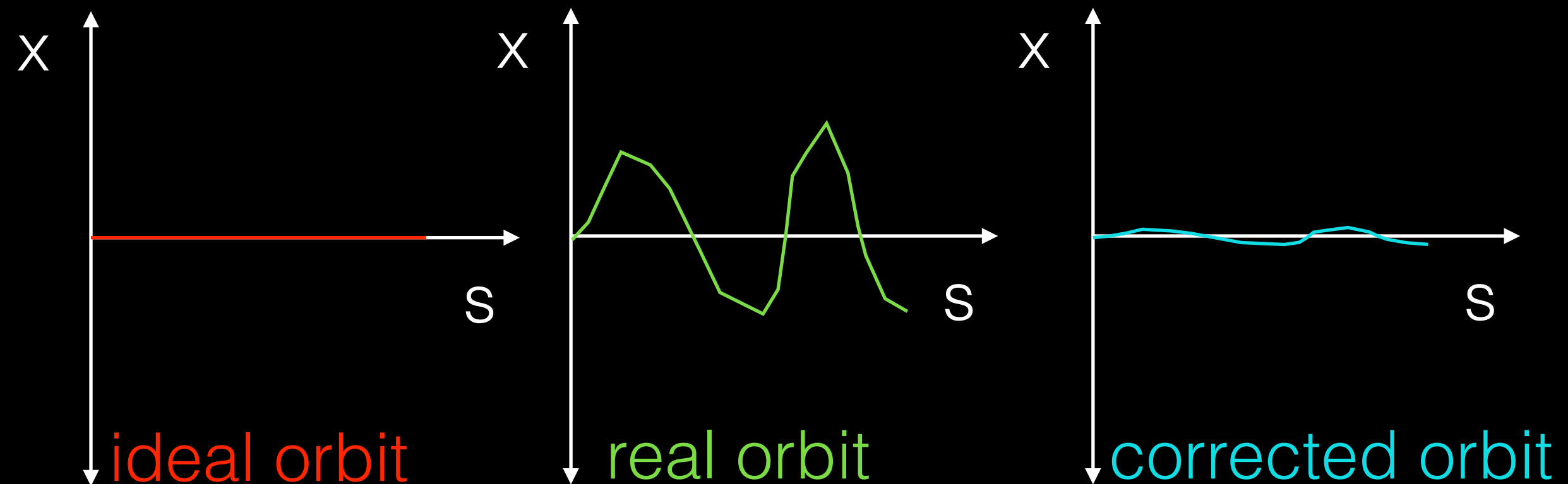
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Evaluate efficiency and performance of orbit correction system:

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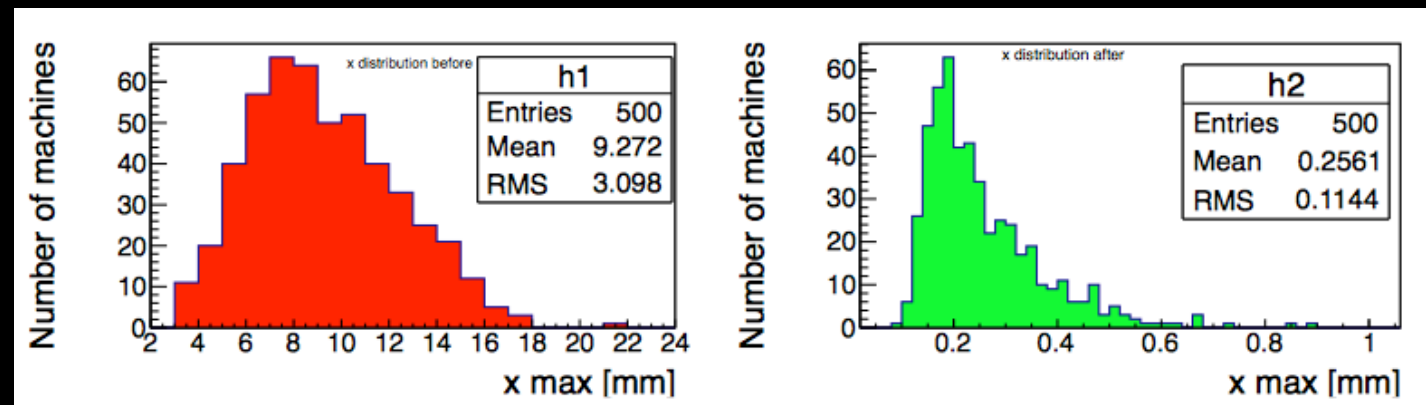
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Max H and V orbit deviation **before** and **after** correction



Before Correction

After Correction

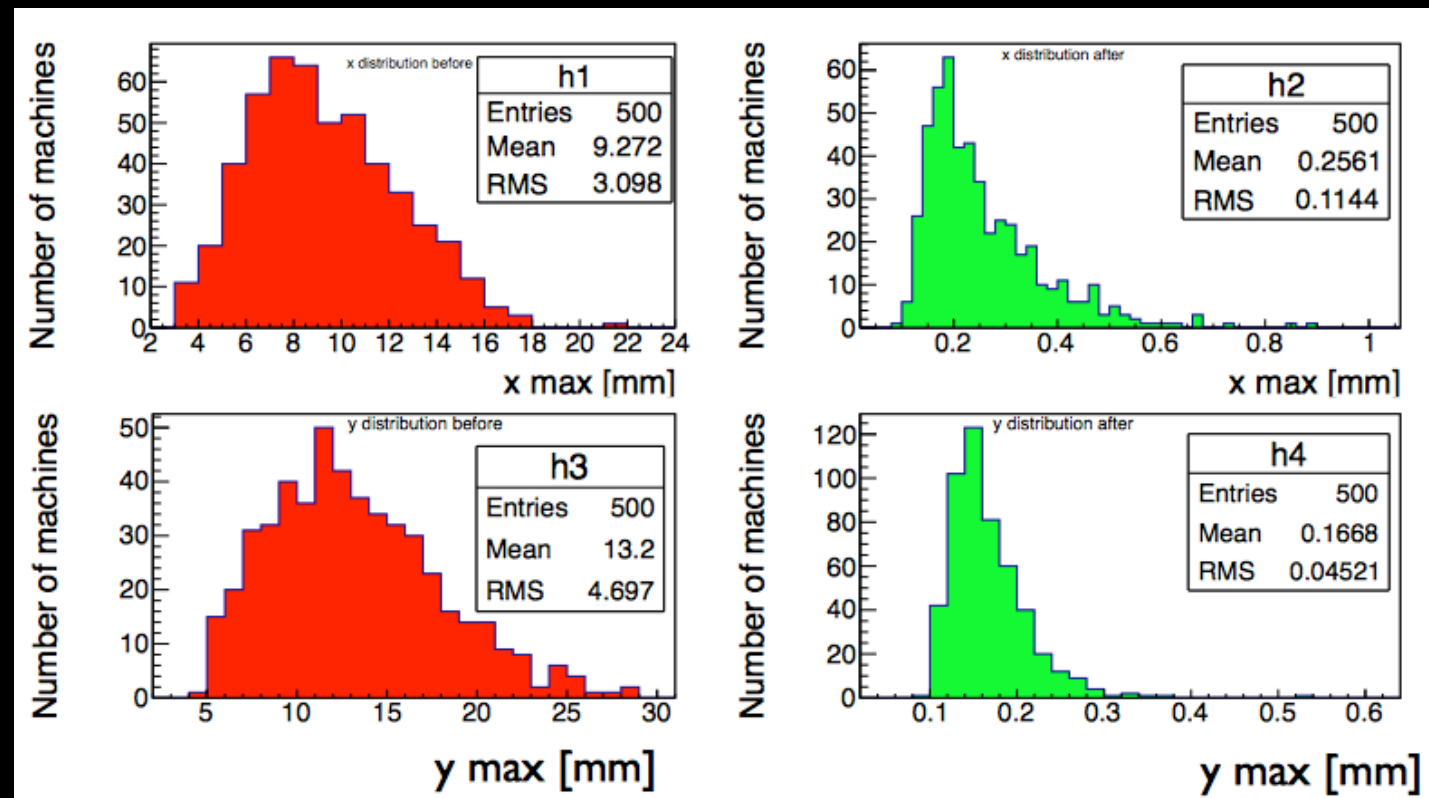
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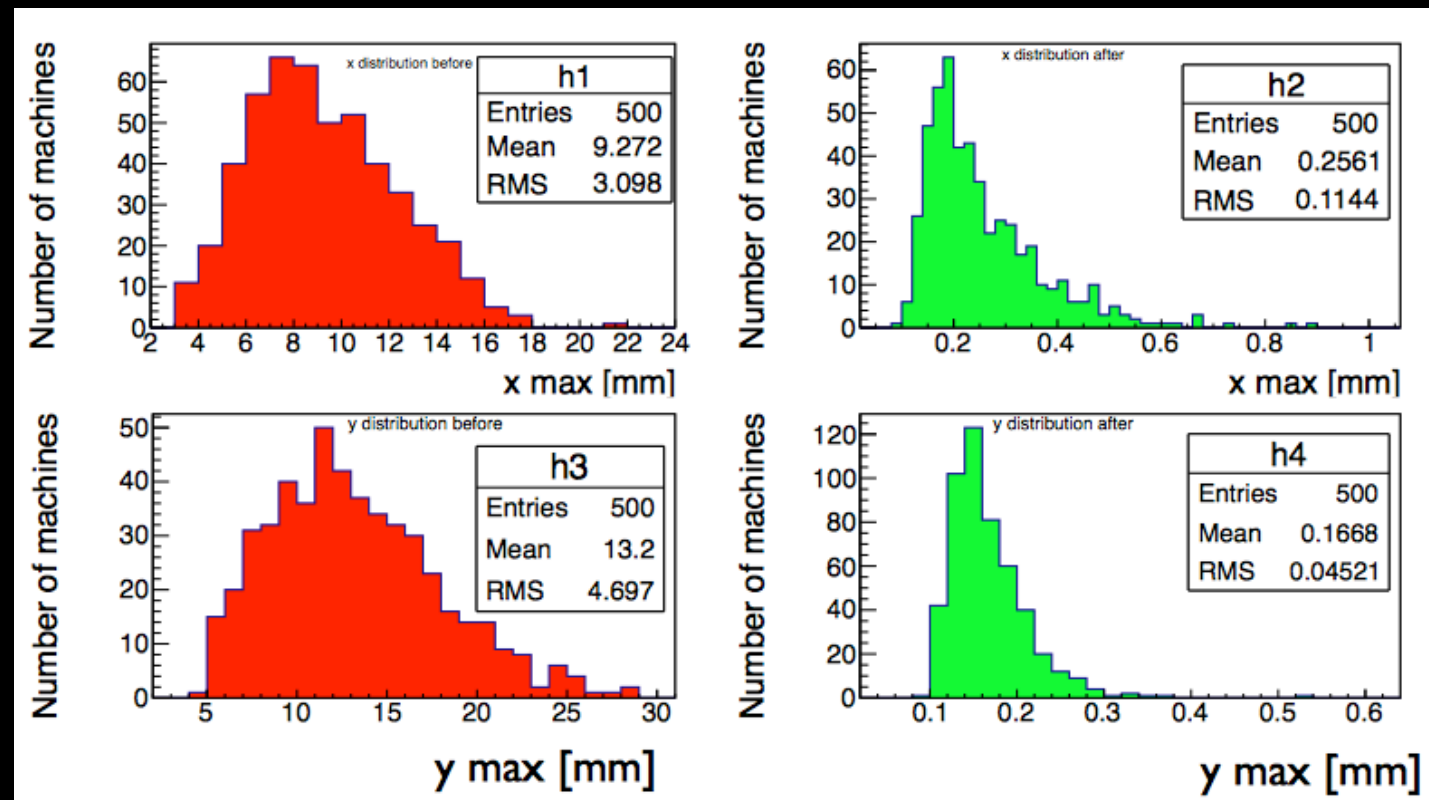
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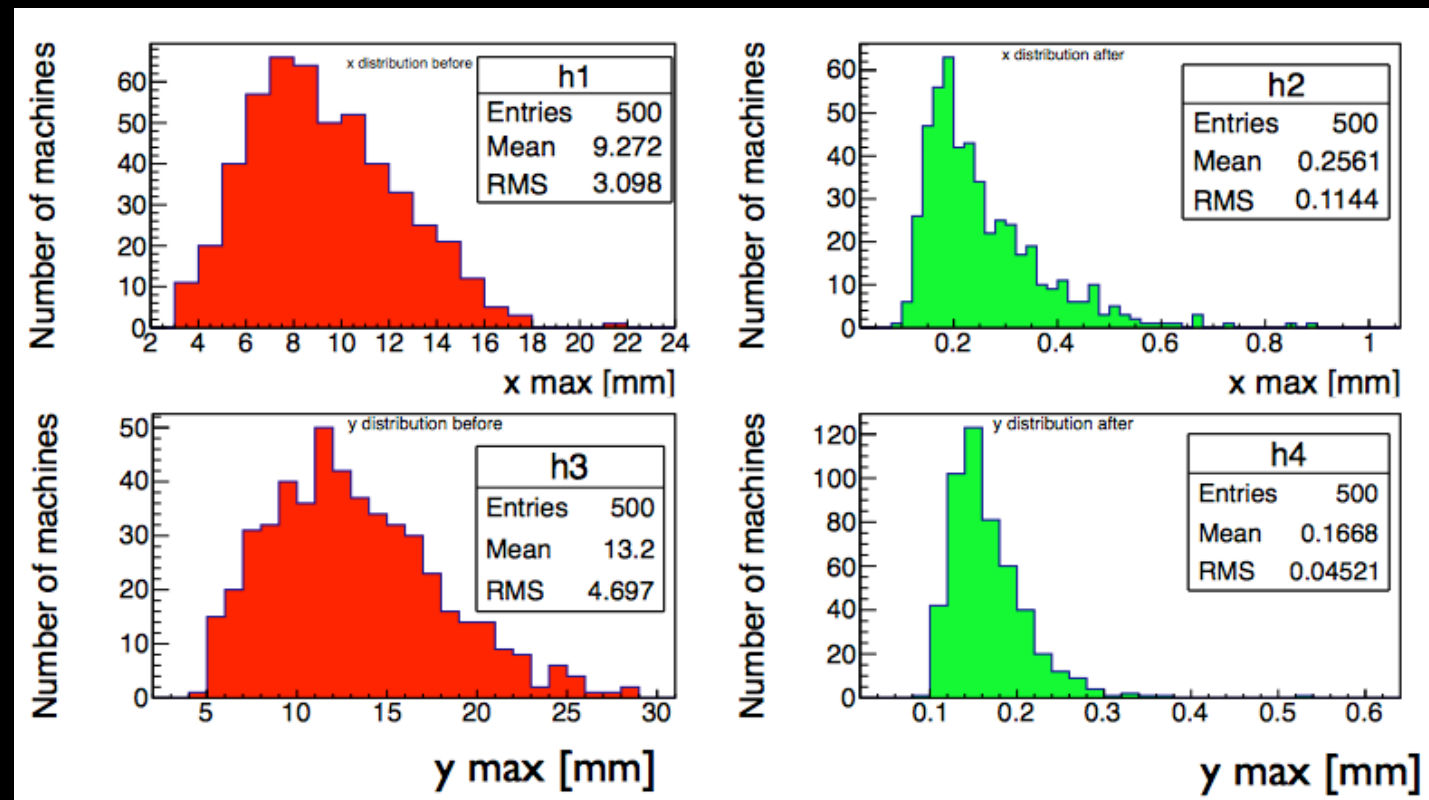
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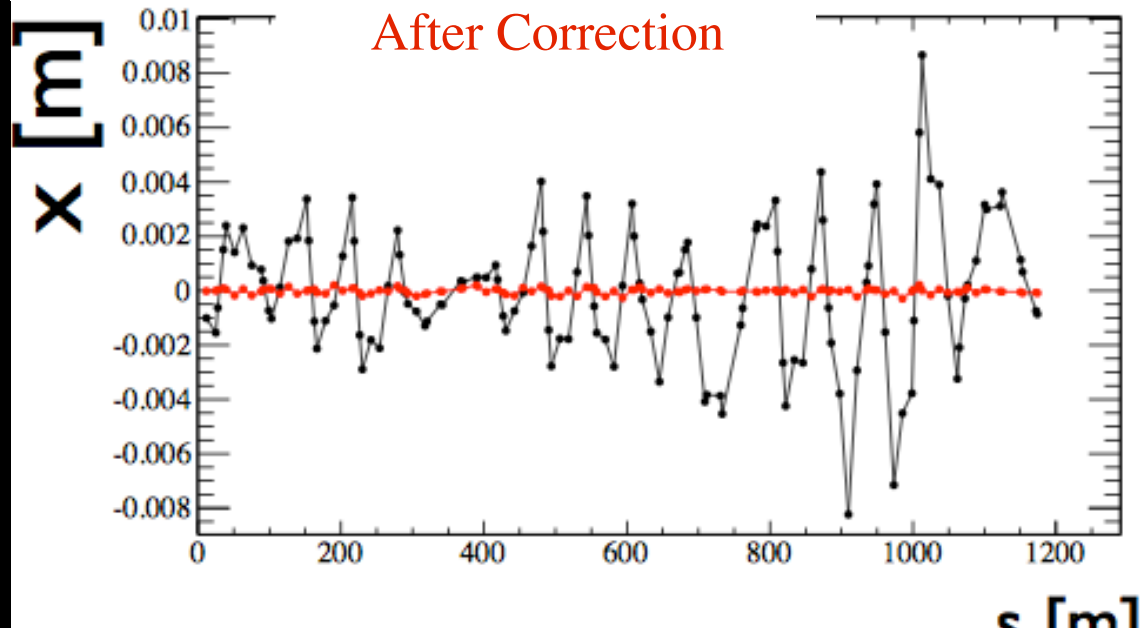


**Before Correction**

**After Correction**

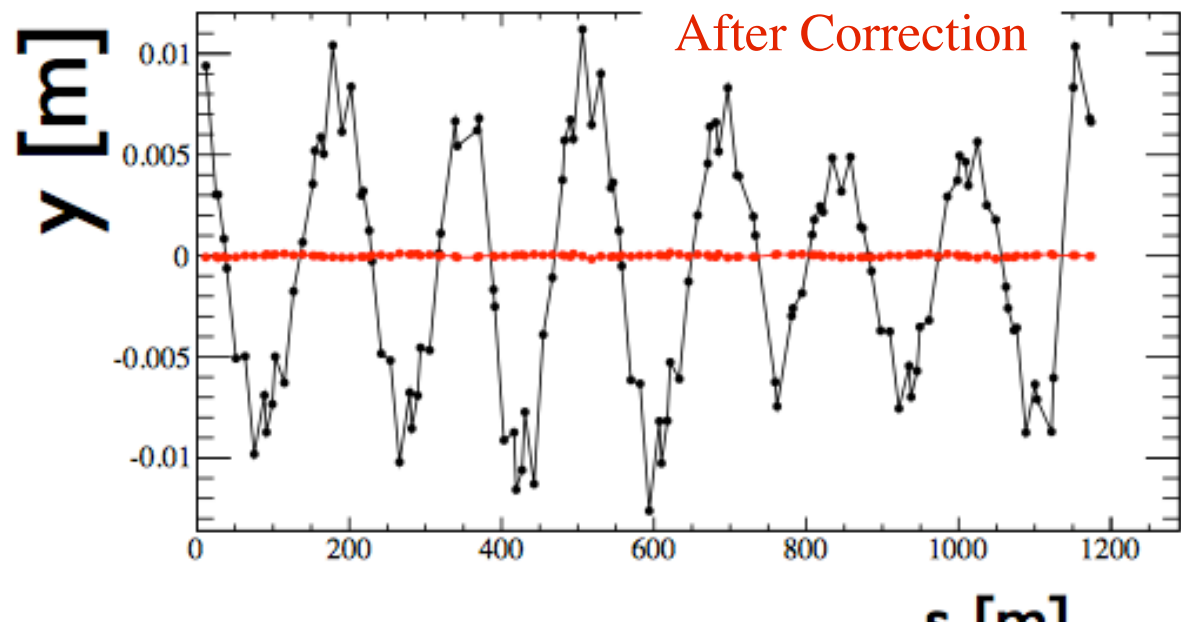
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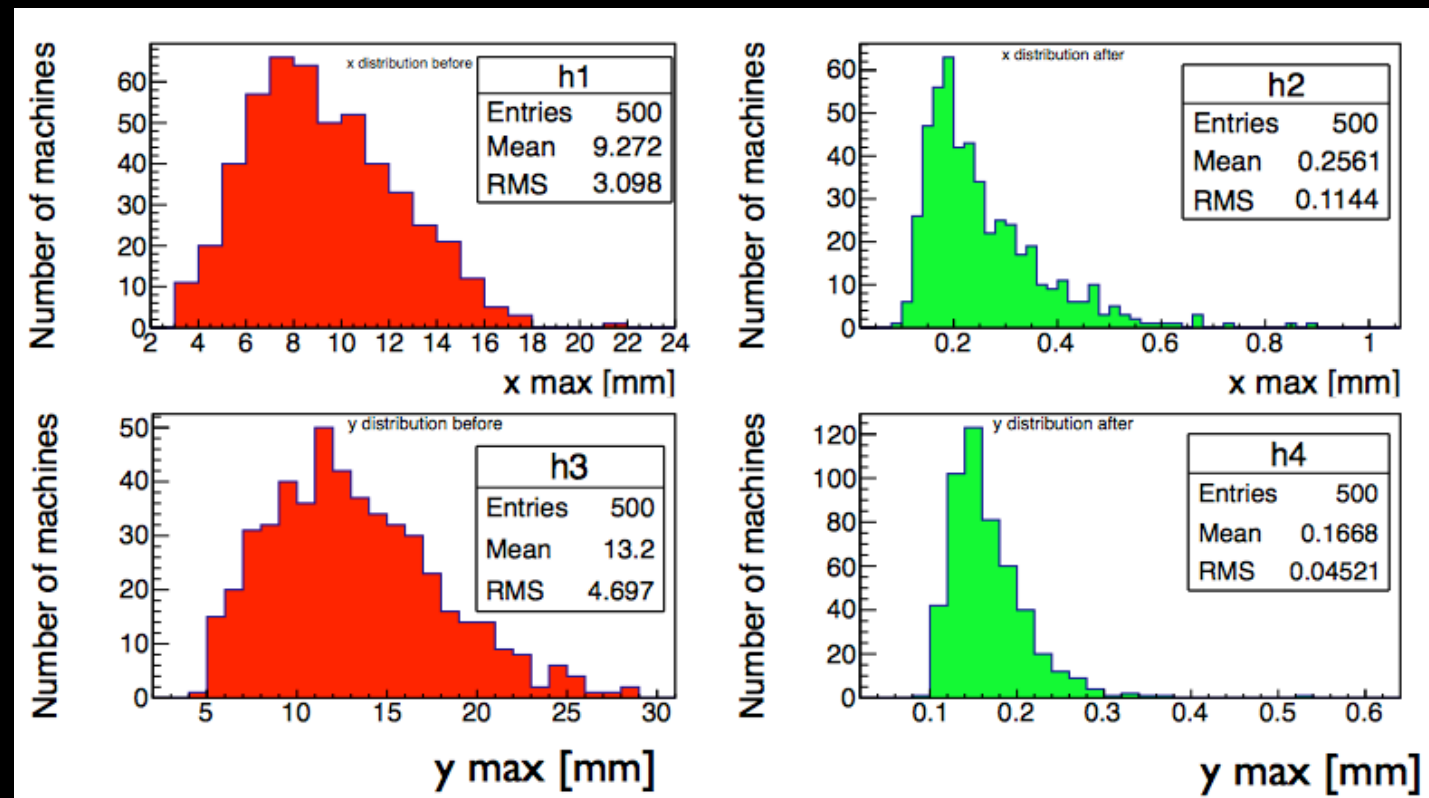
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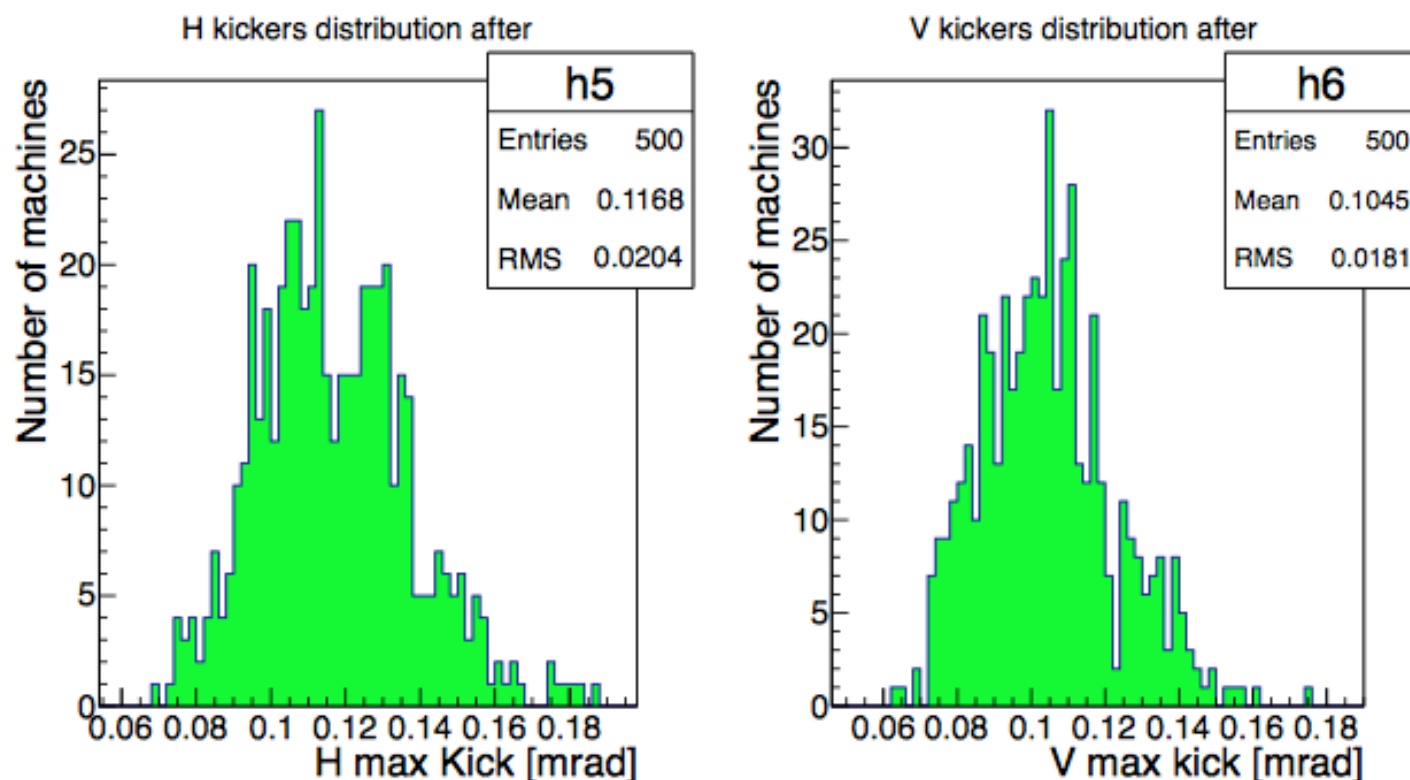
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**Before Correction**      **After Correction**



Correctors' strength needed <0.2 mrad ( $\sim 0.05$  T for  $E=50$  GeV), i.e. well within the limits [7]

# Collimators

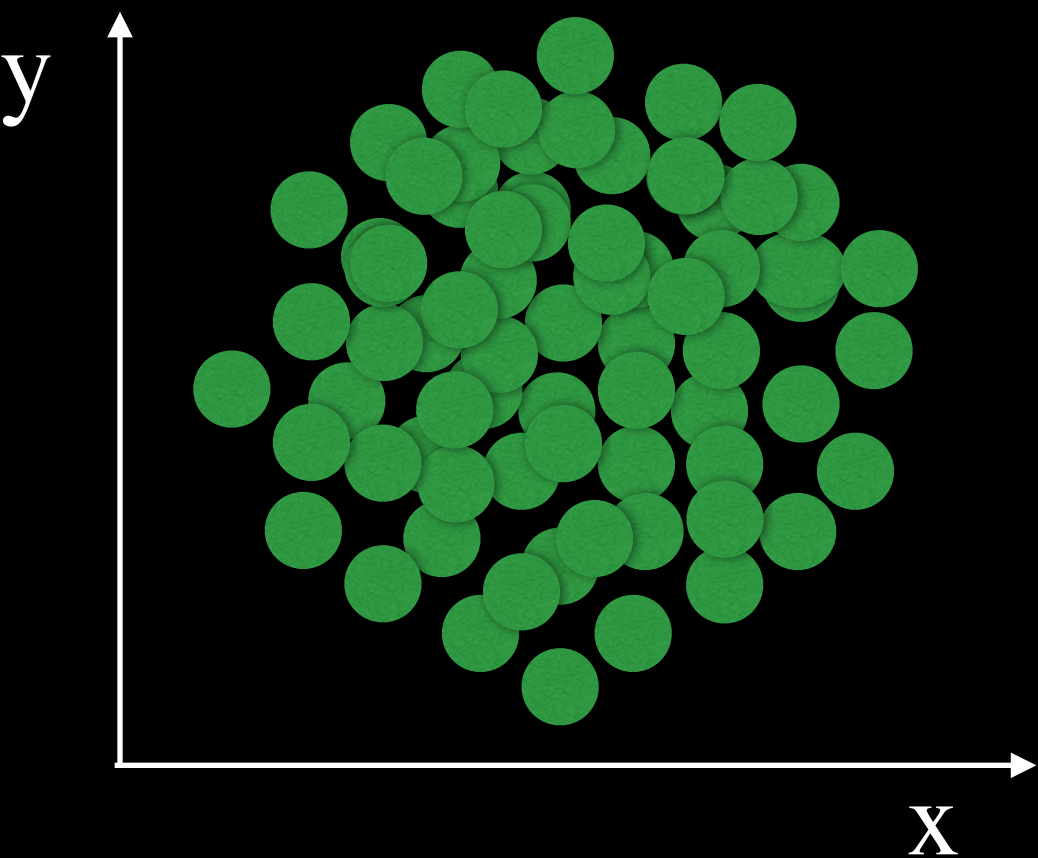
# Collimators

- Why do we need collimators?



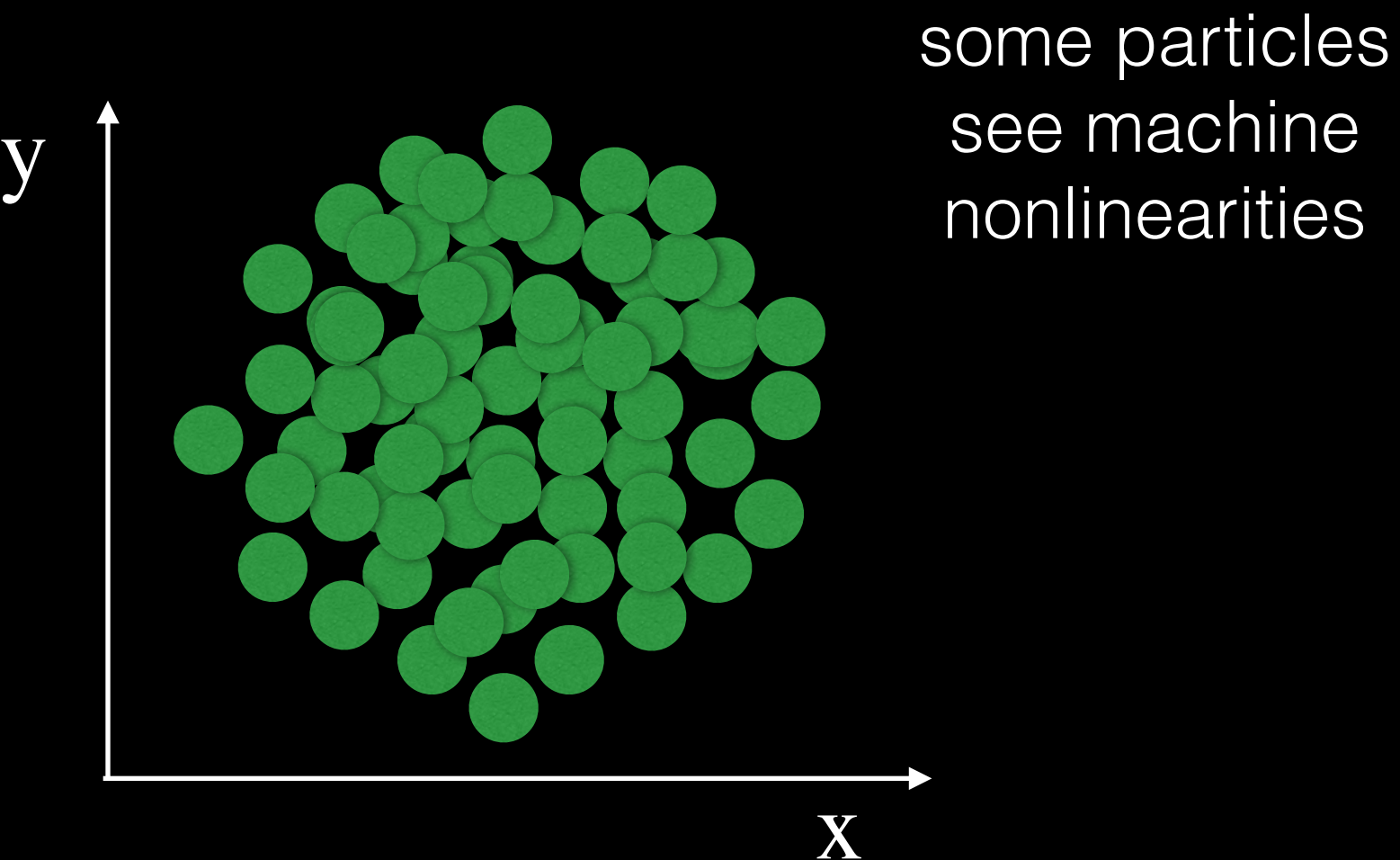
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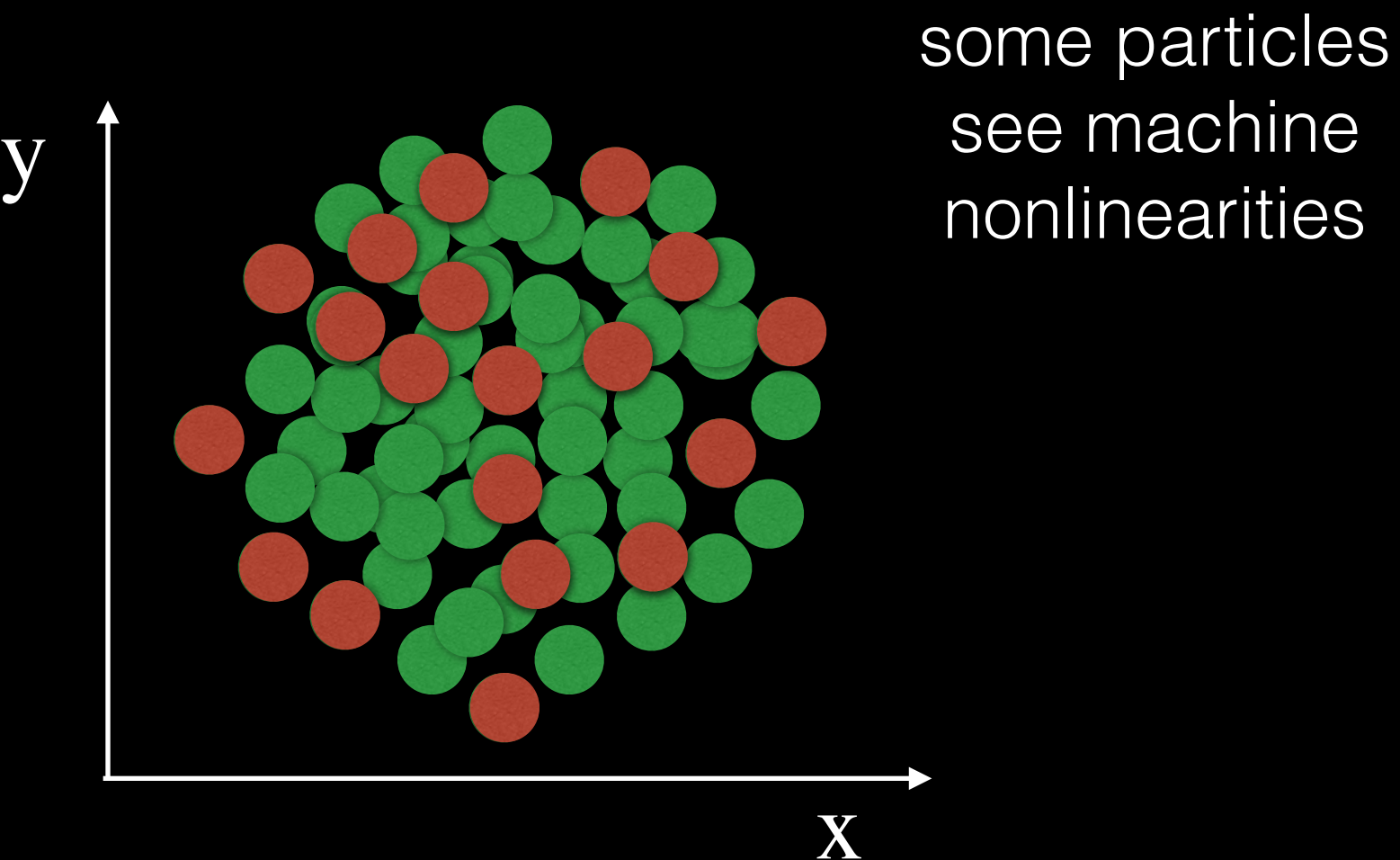
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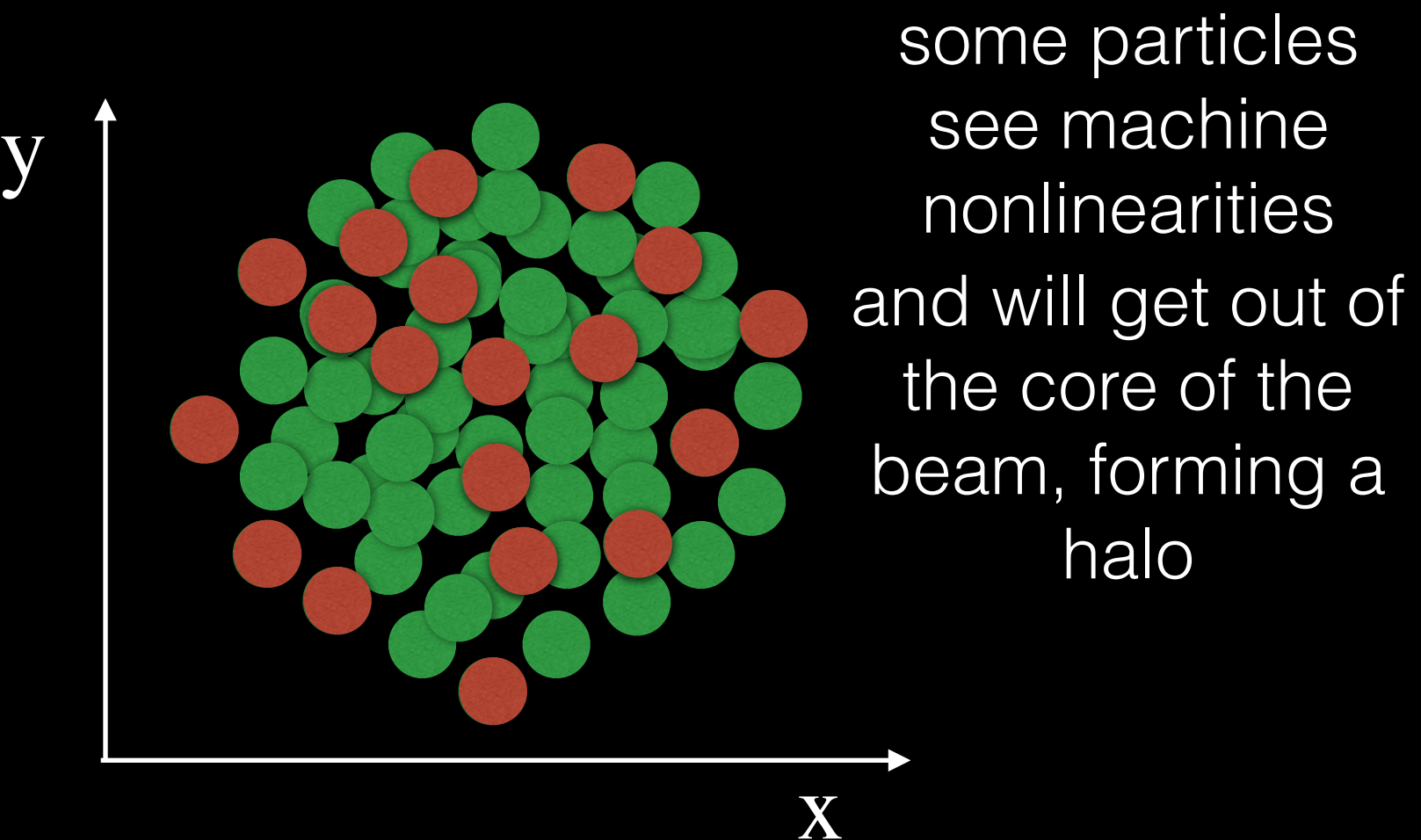
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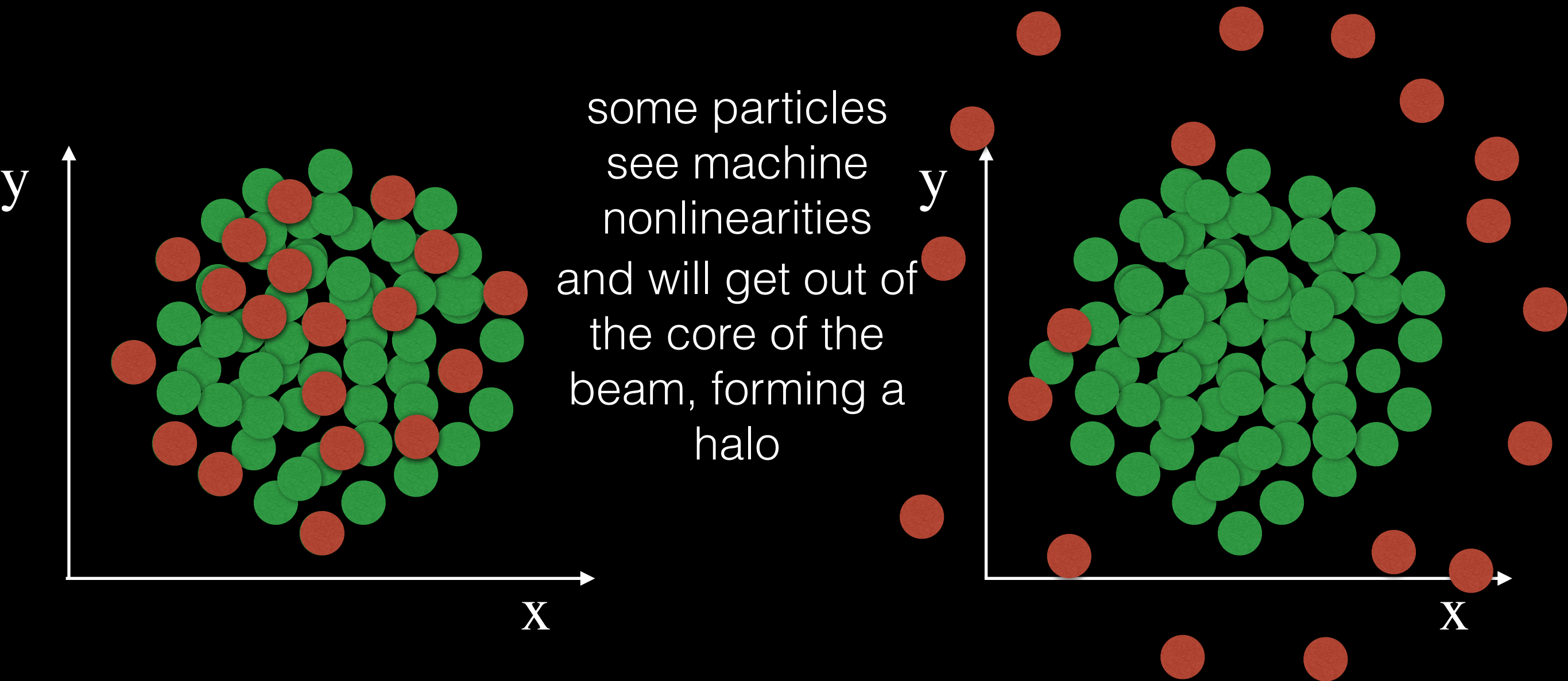
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  - to prevent halo particles from hitting the superconducting magnets of the HP-PS ring (avoid magnets quenching)
  - to limit equipment irradiation close to the beam
  - to localise slow losses in controlled way in properly equipped locations: dedicated LSS (Long Straight Section) for transverse collimation



## ■ What type of collimators?

- Primaries (HP): increase chance that halo particles will be absorbed later on by secondary collimators
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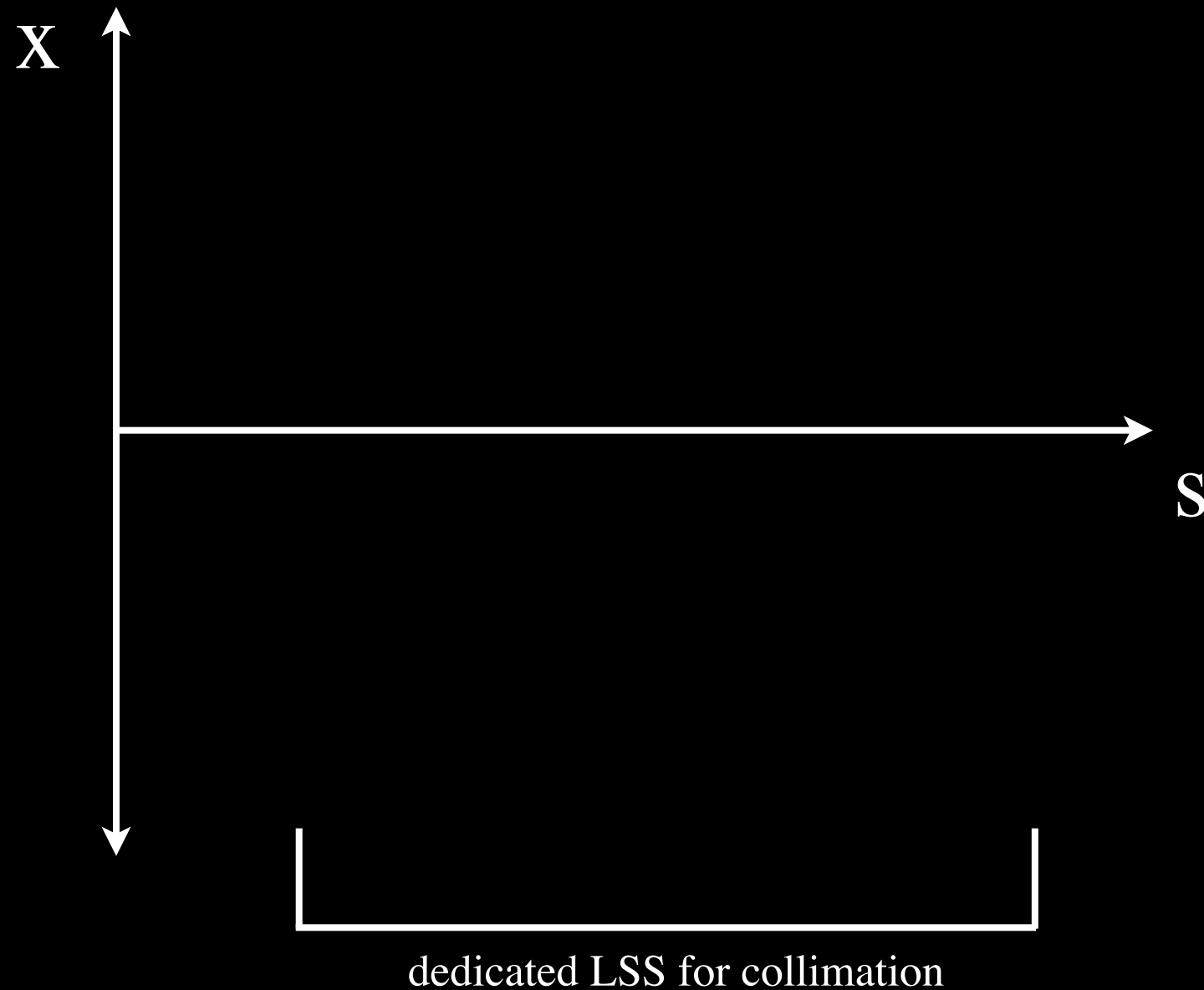
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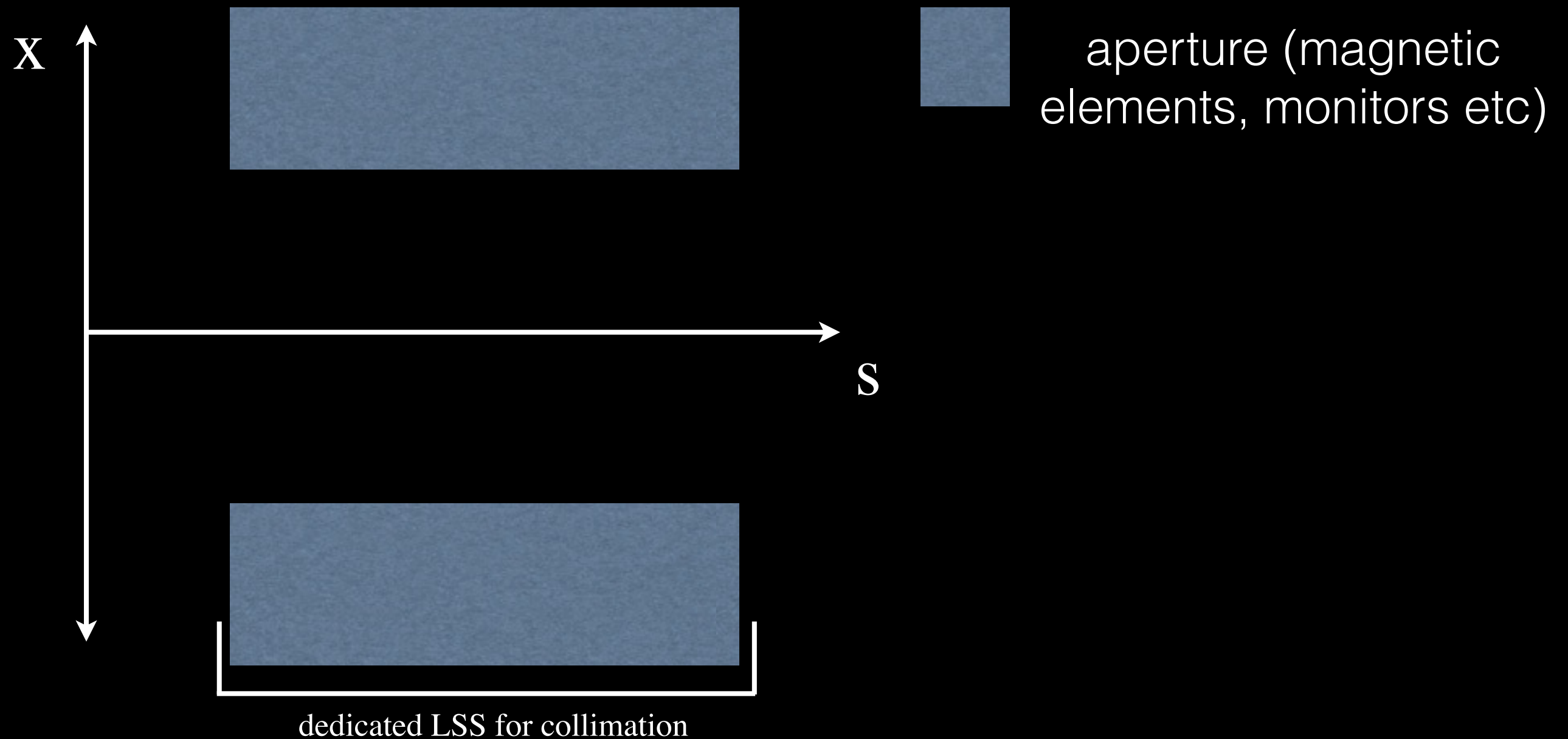
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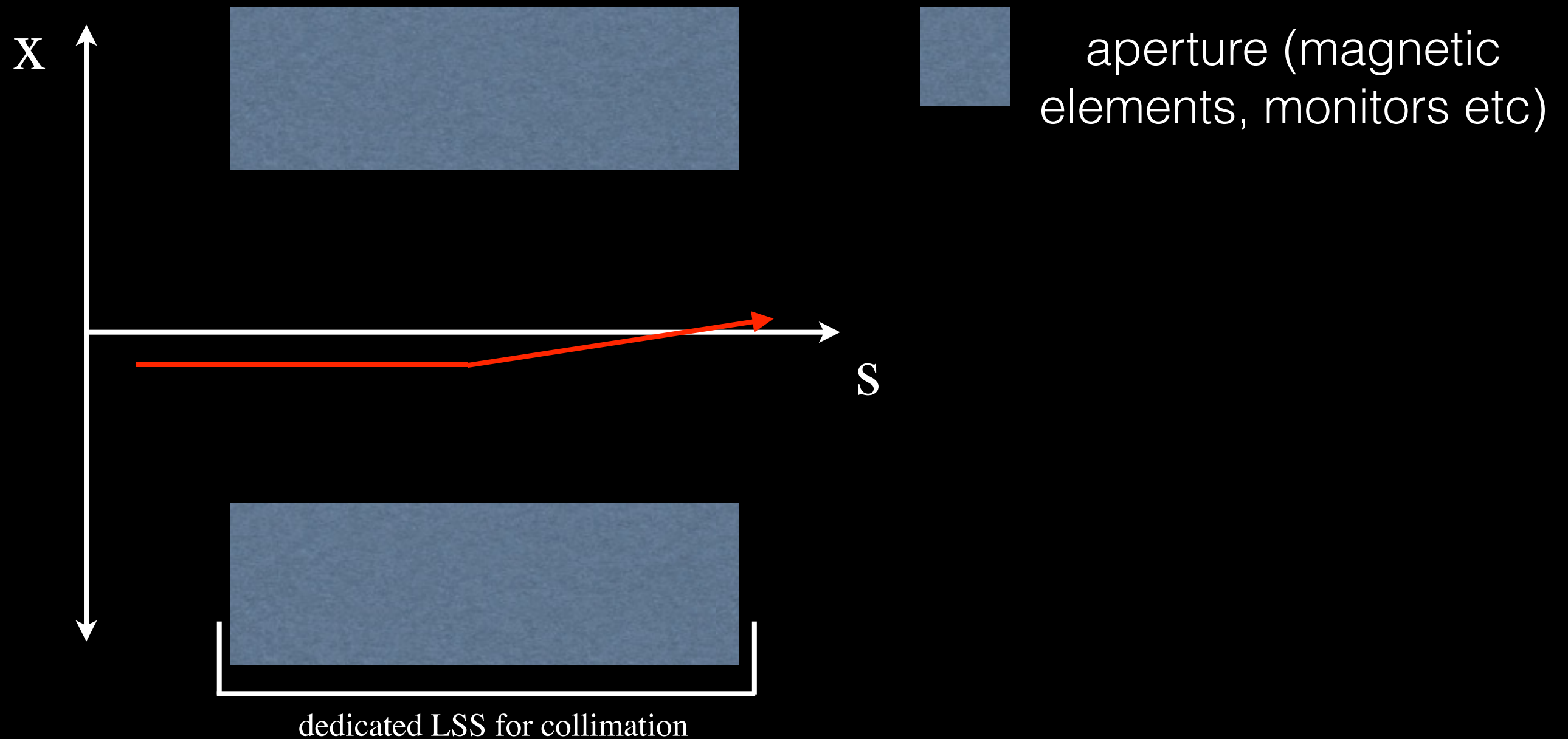
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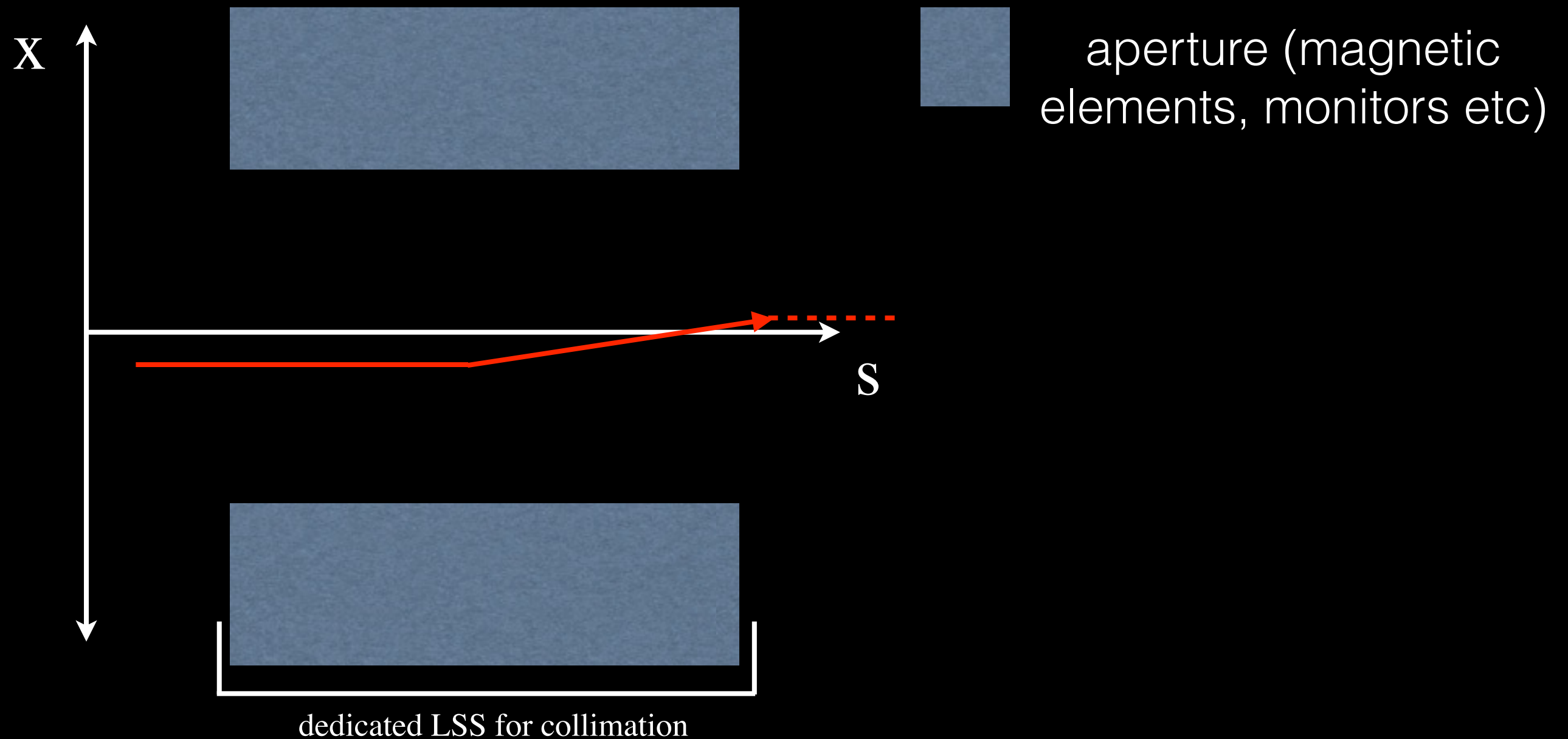
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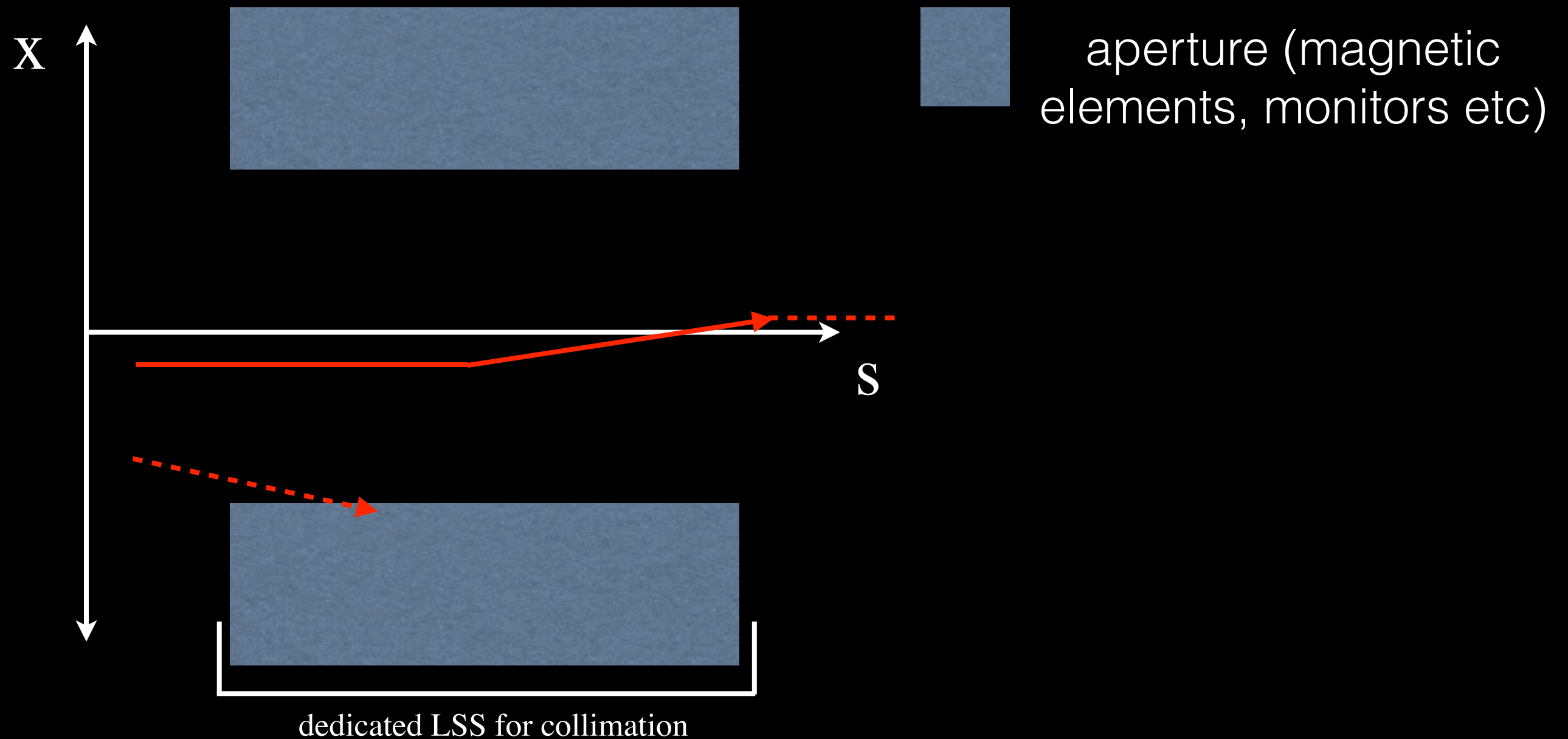


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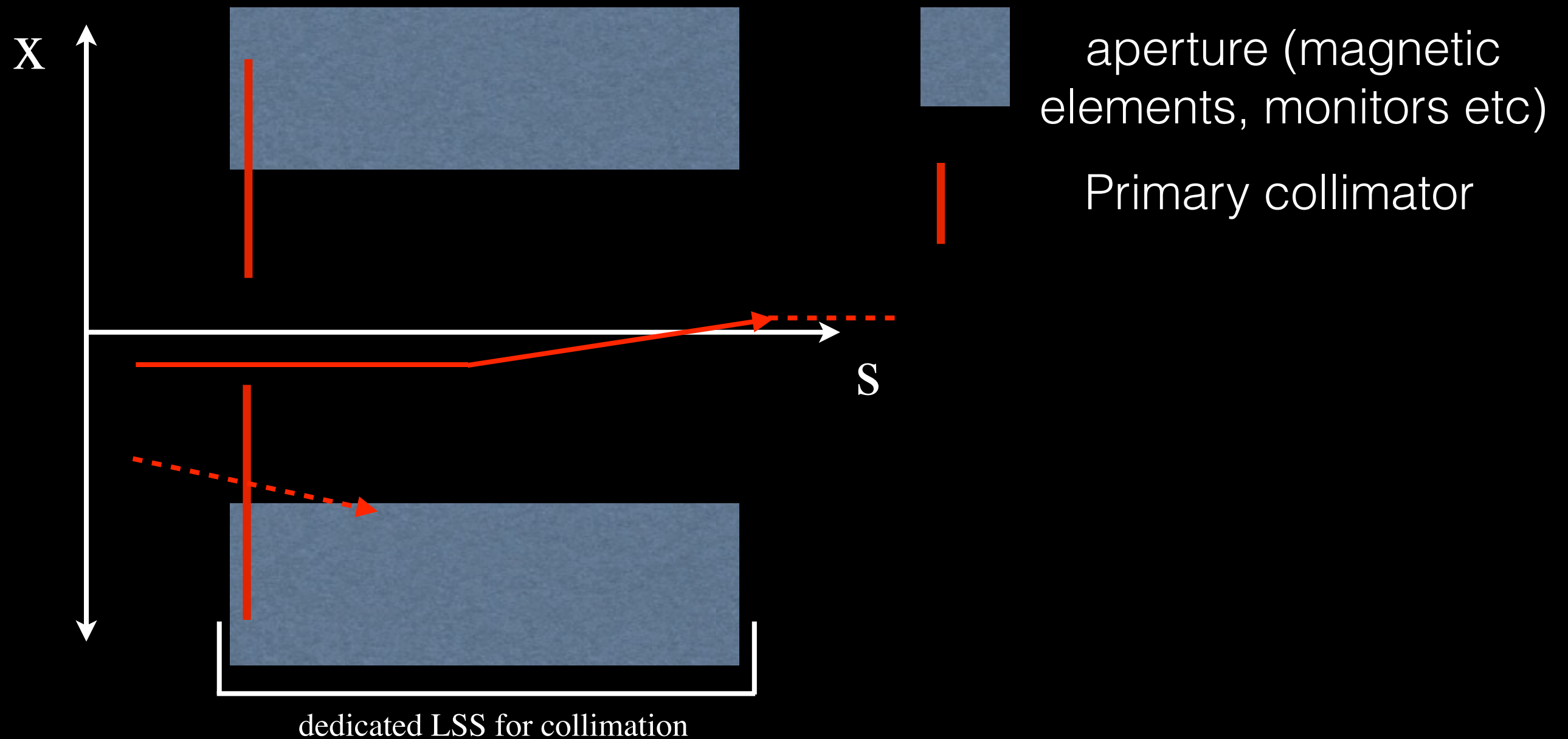
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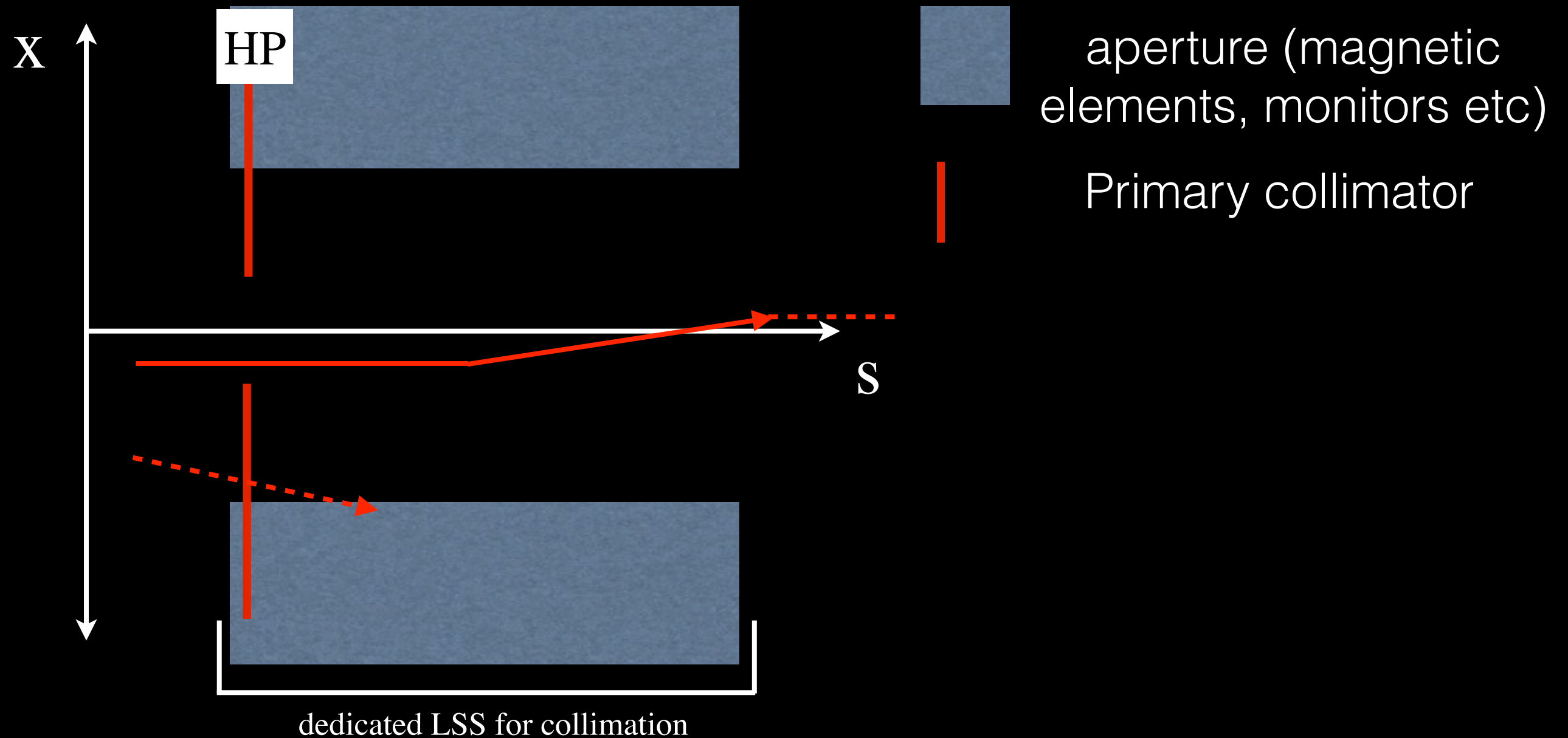
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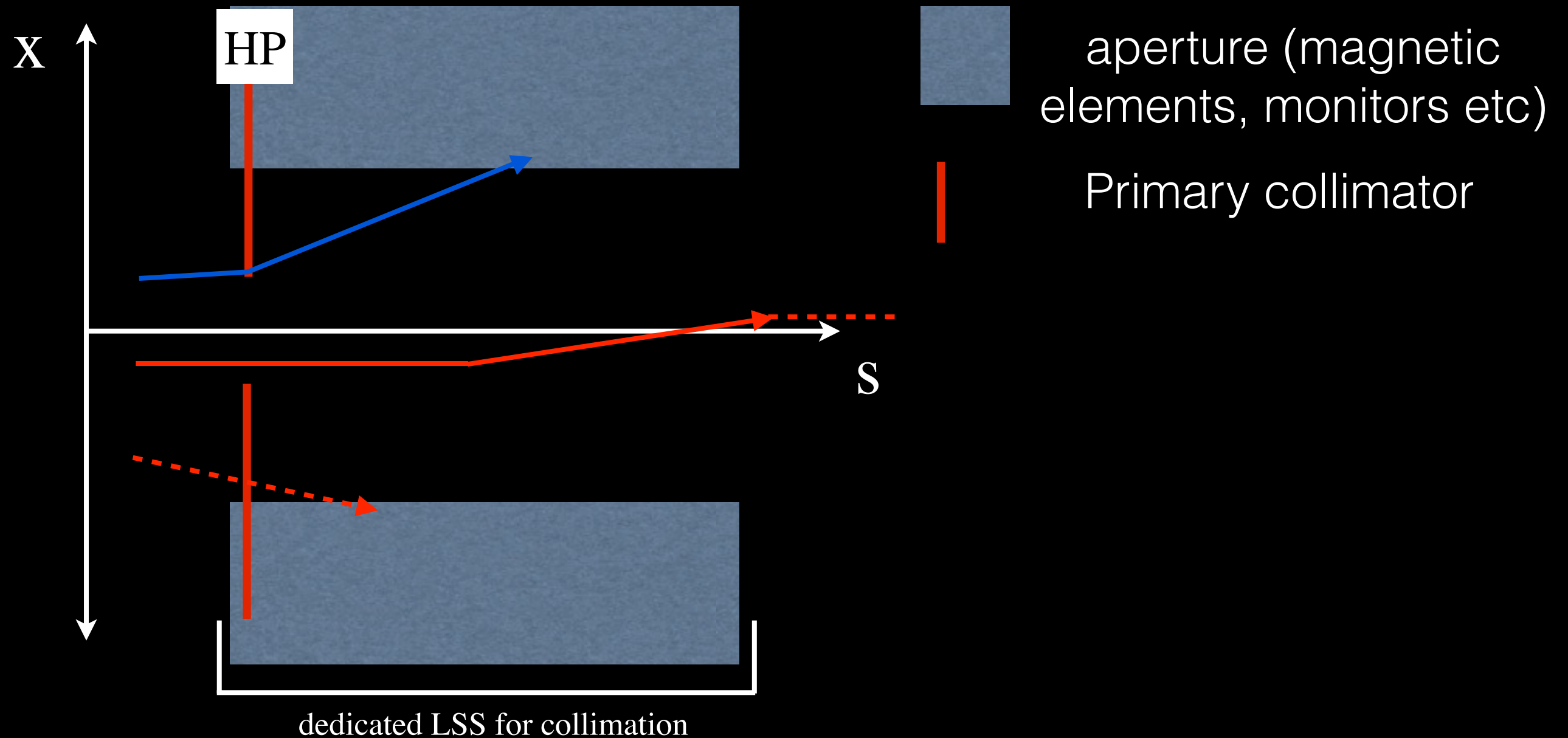
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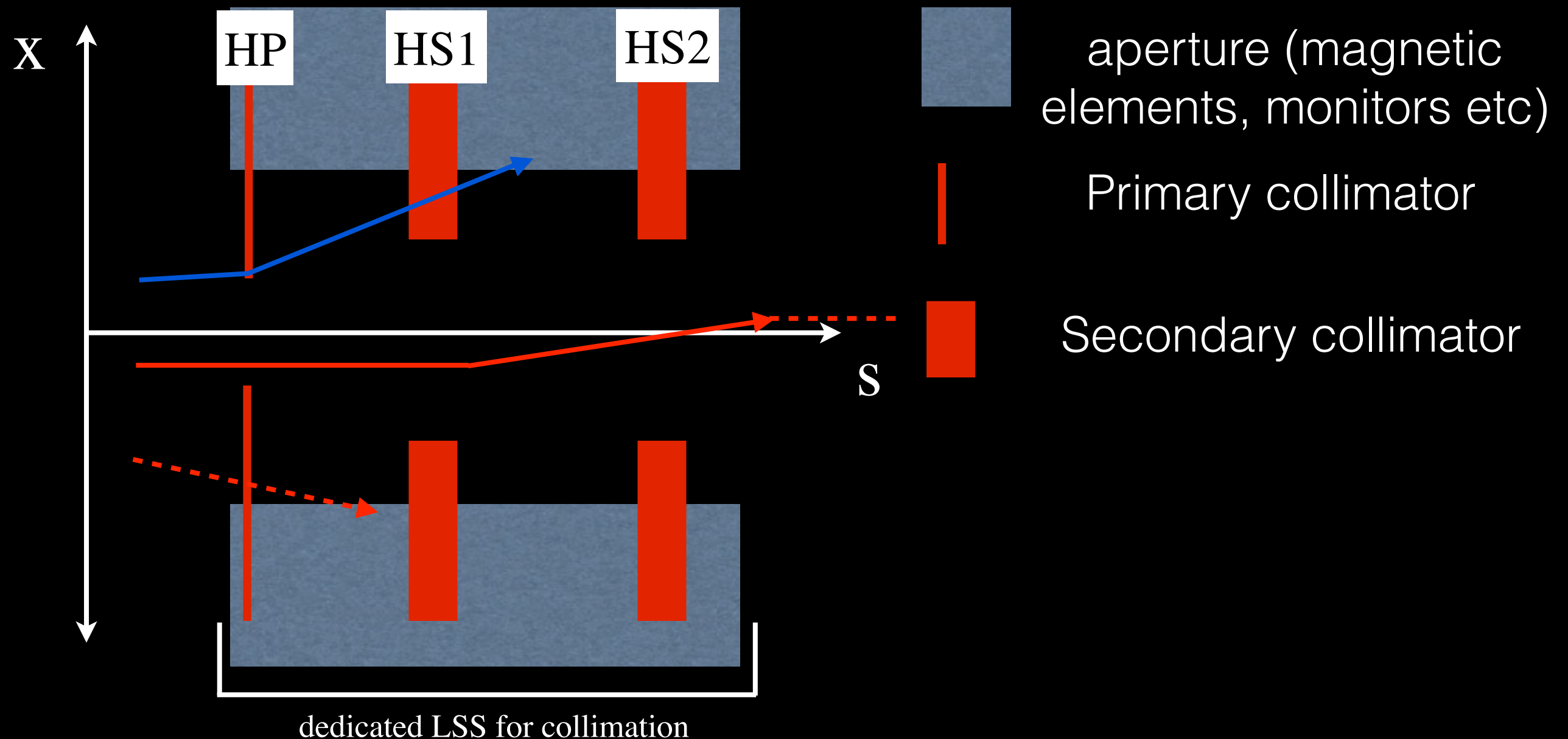


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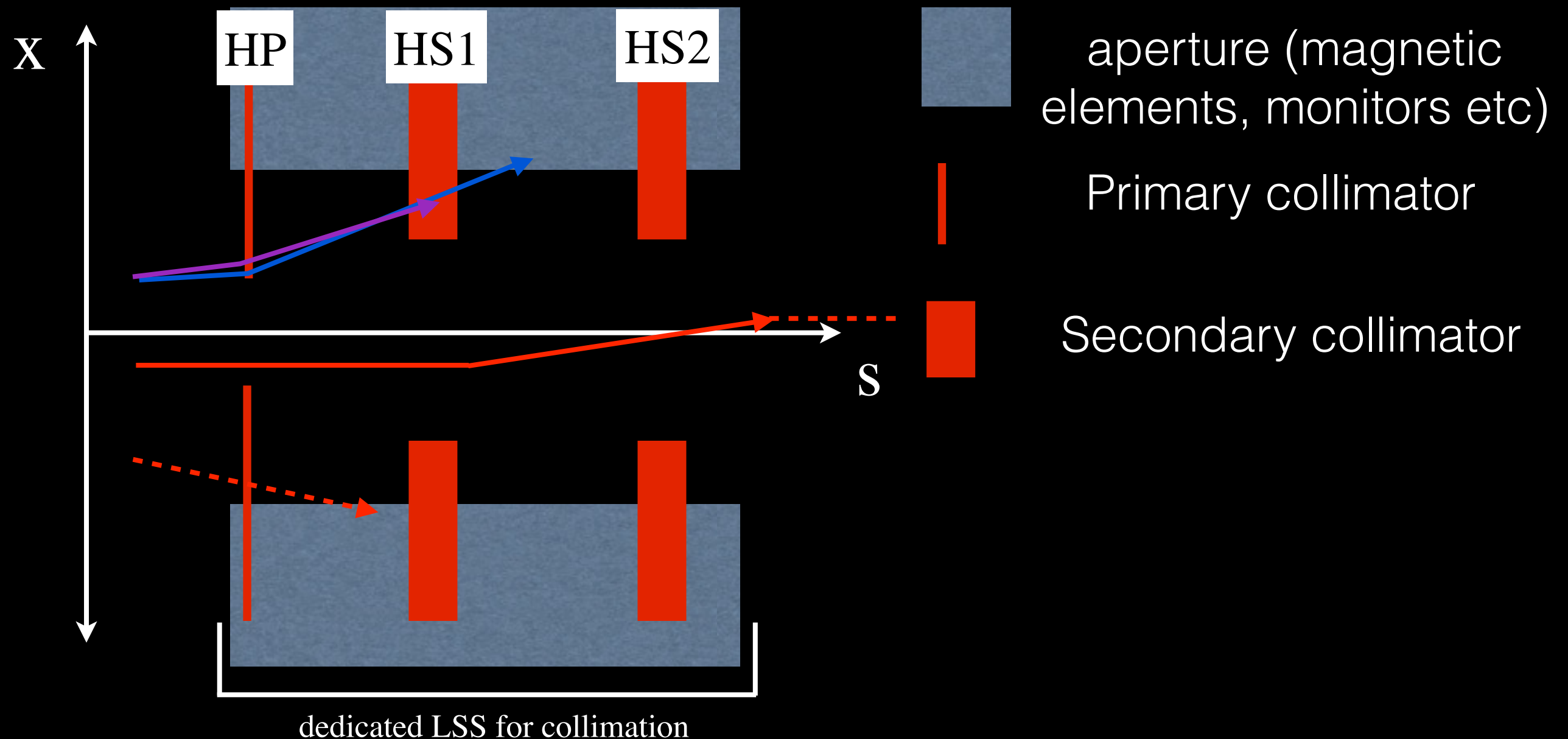
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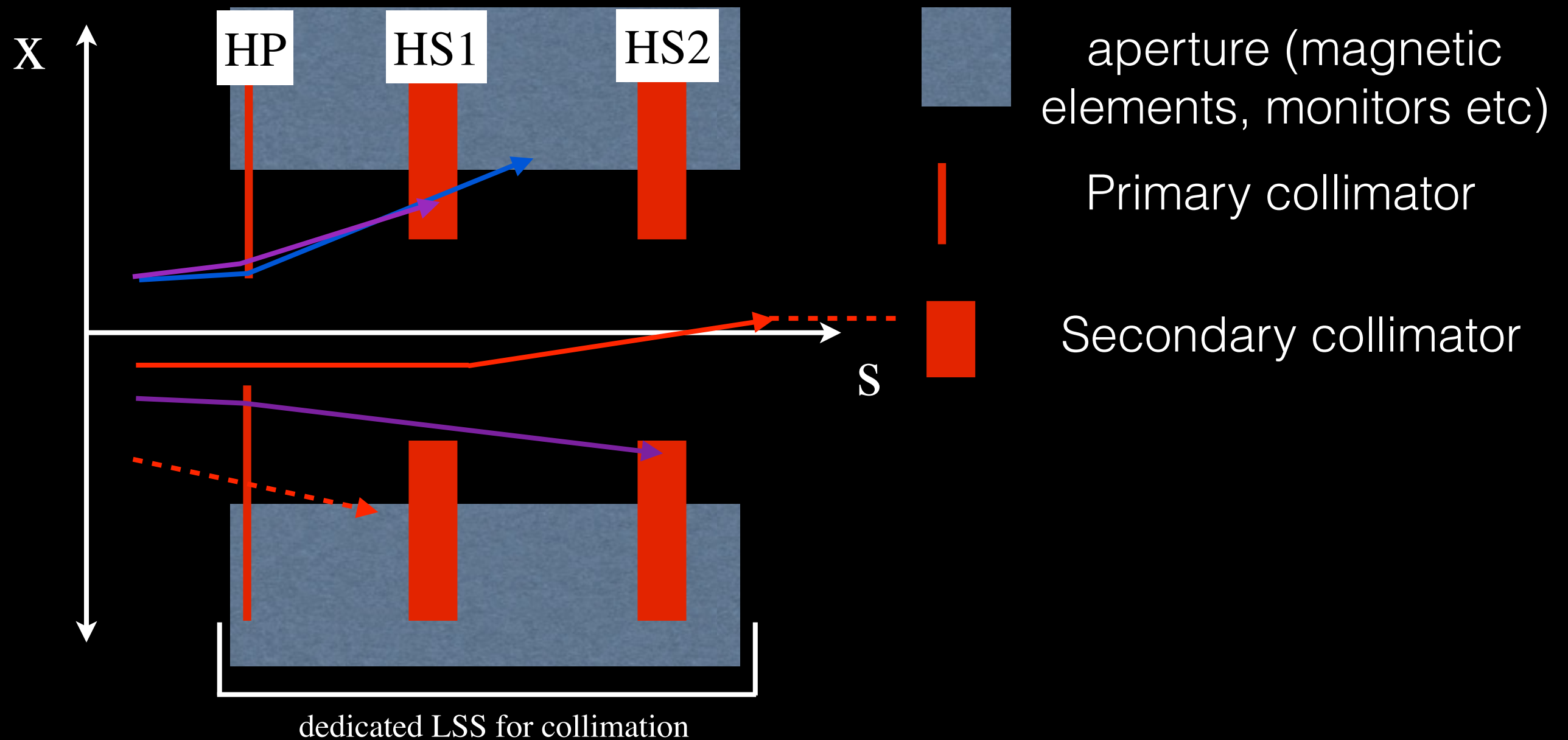


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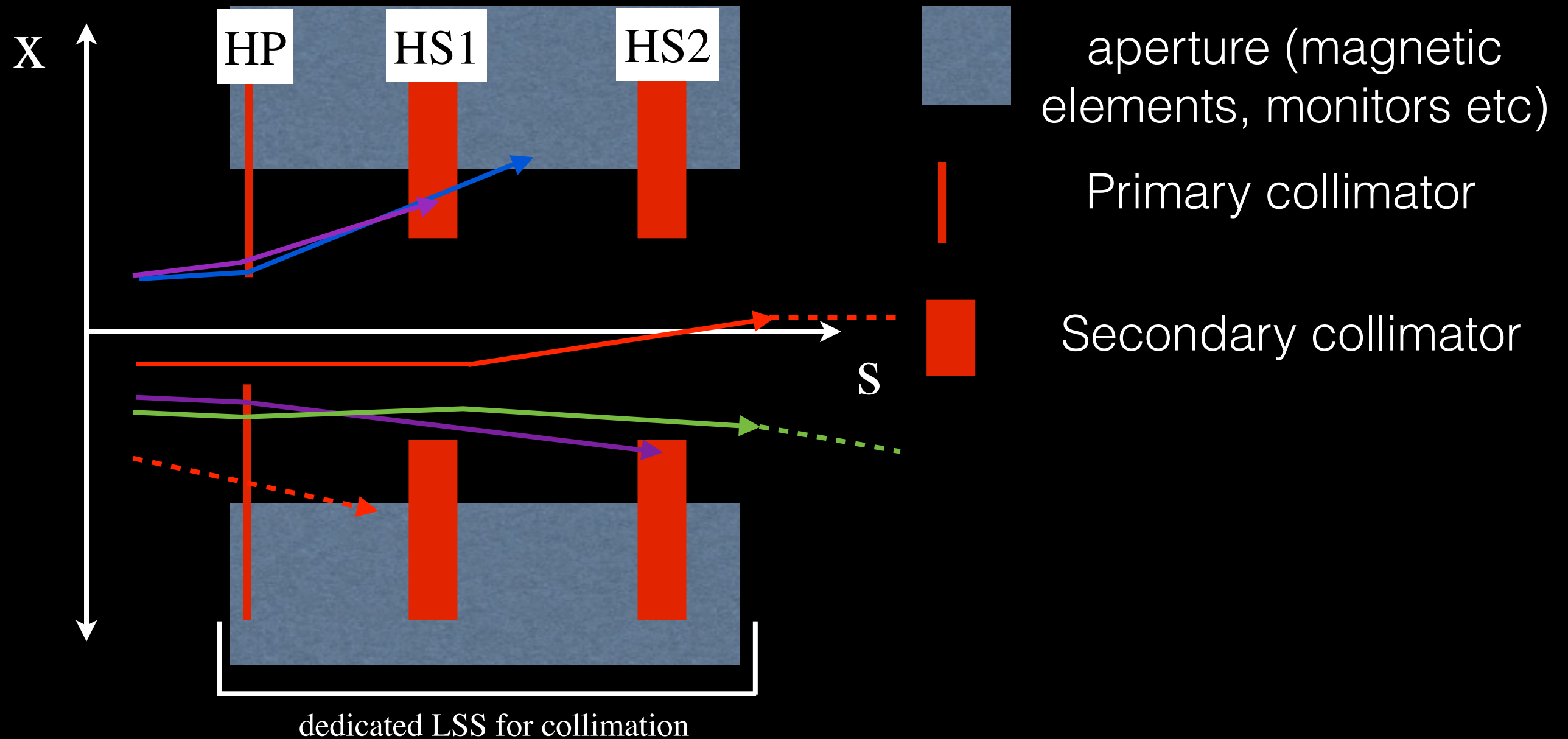


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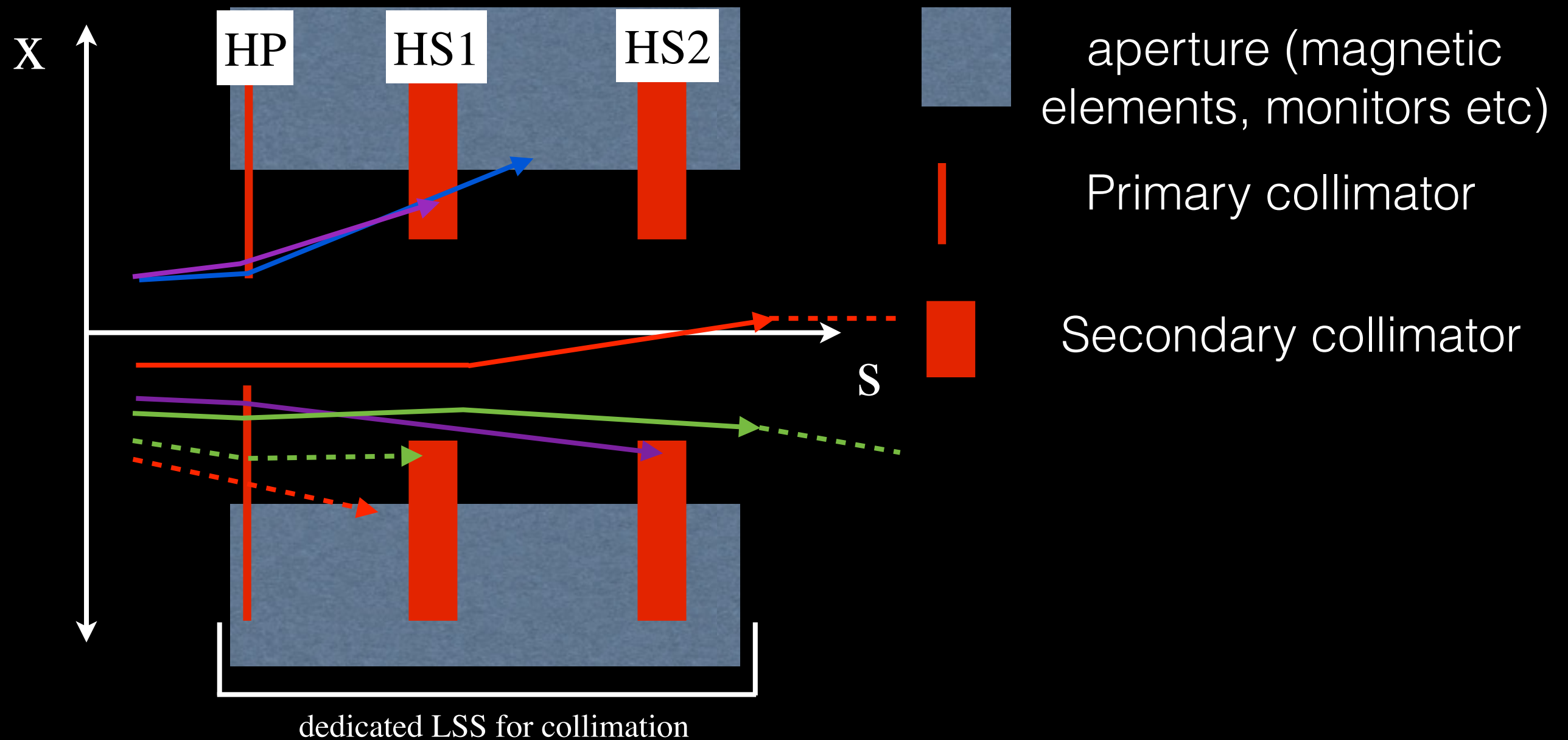
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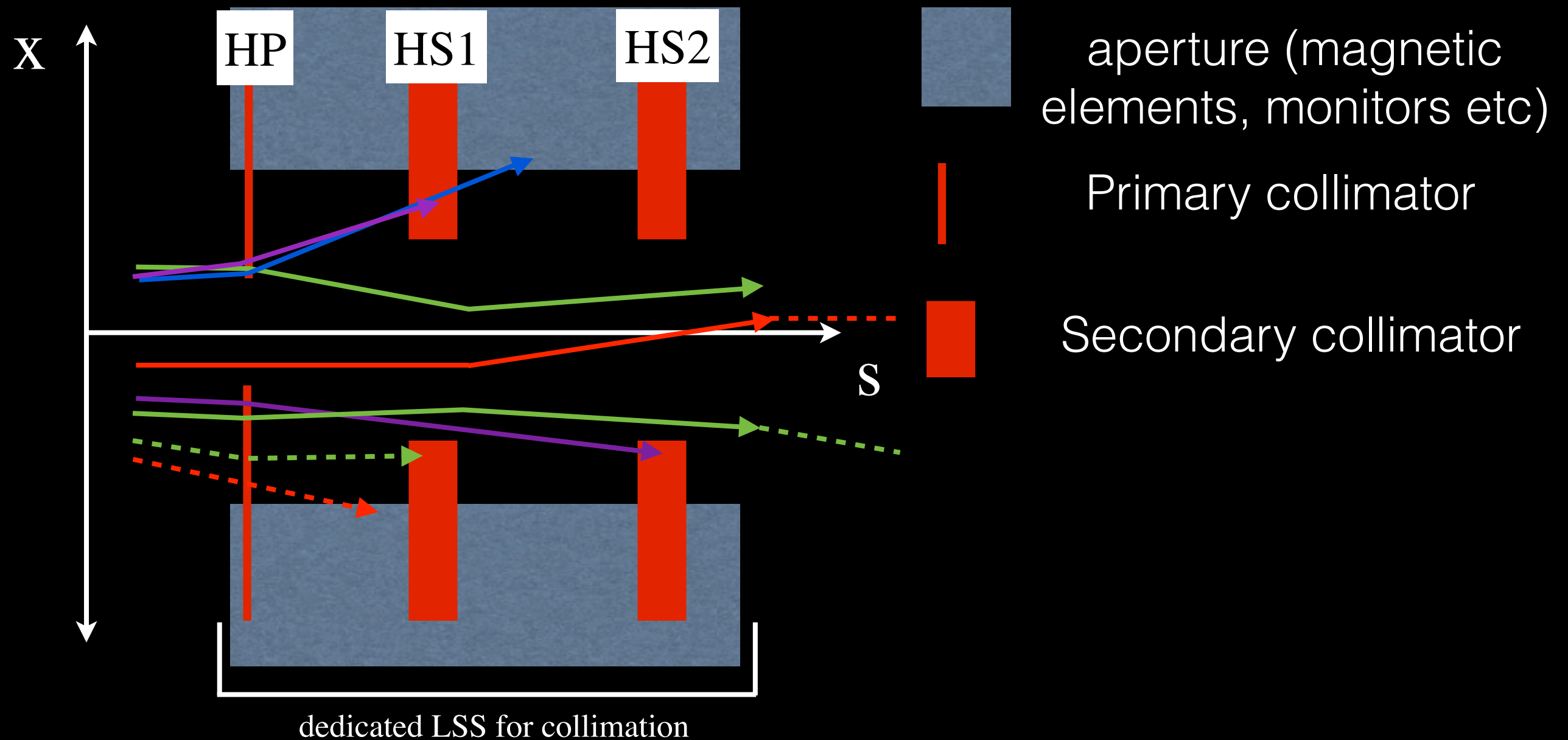
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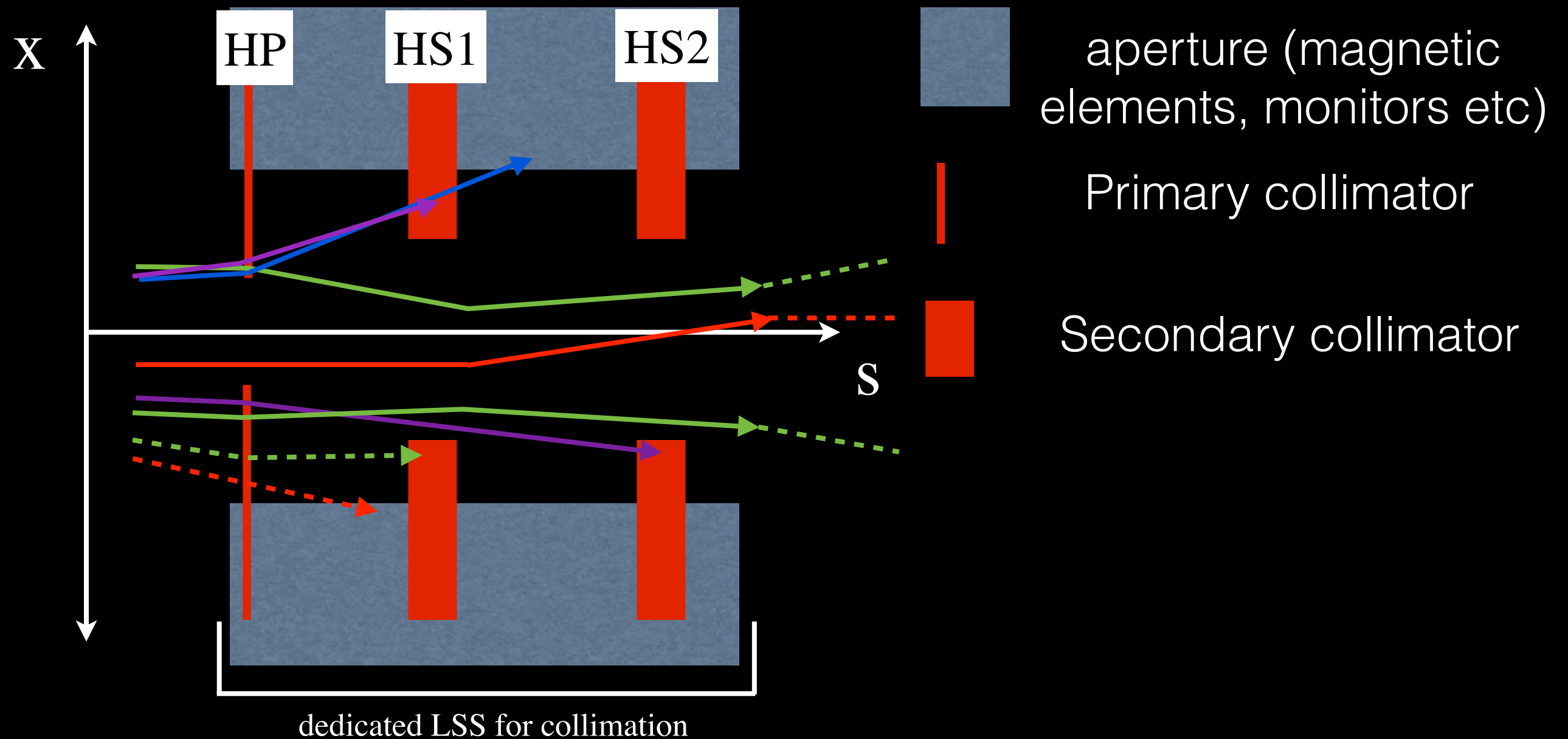
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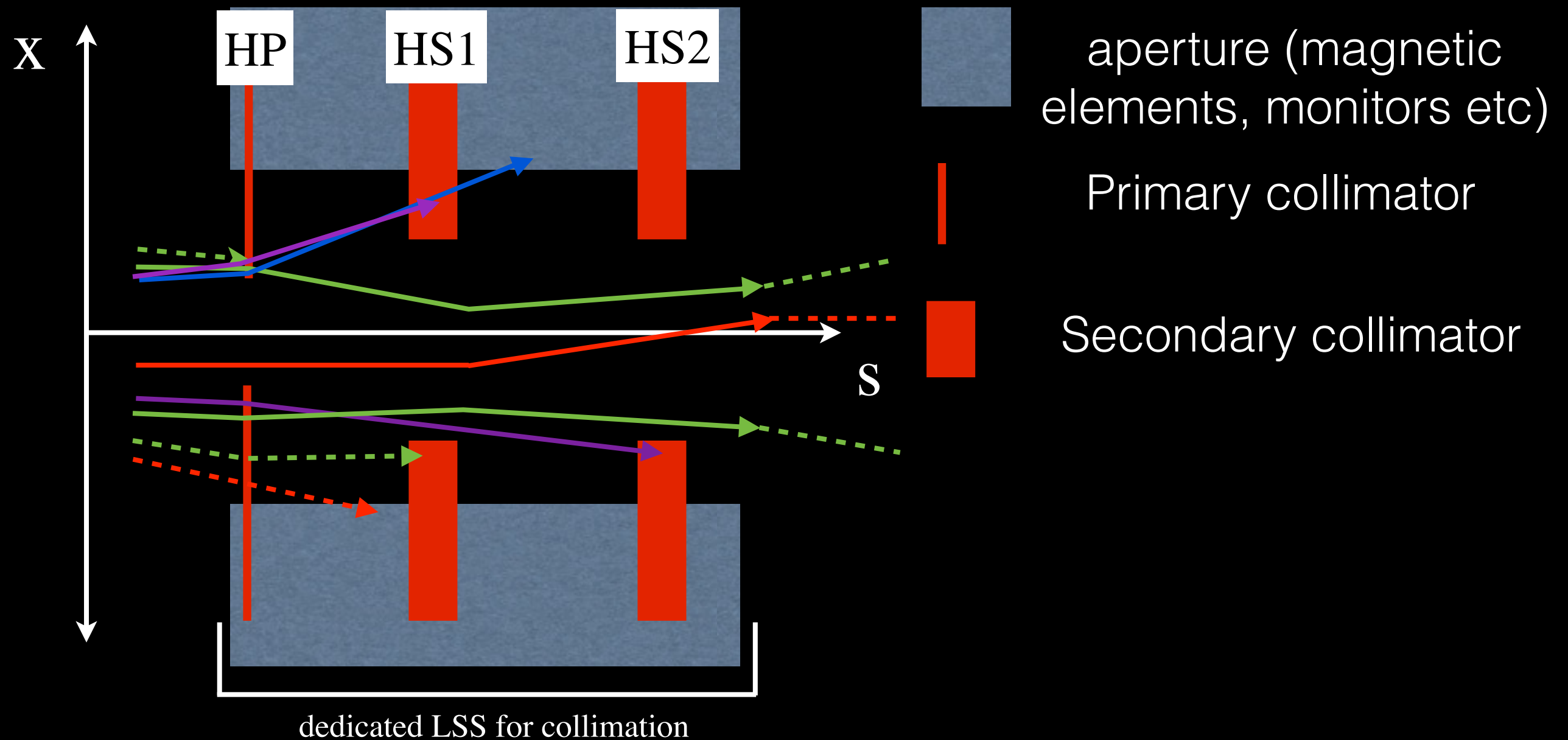
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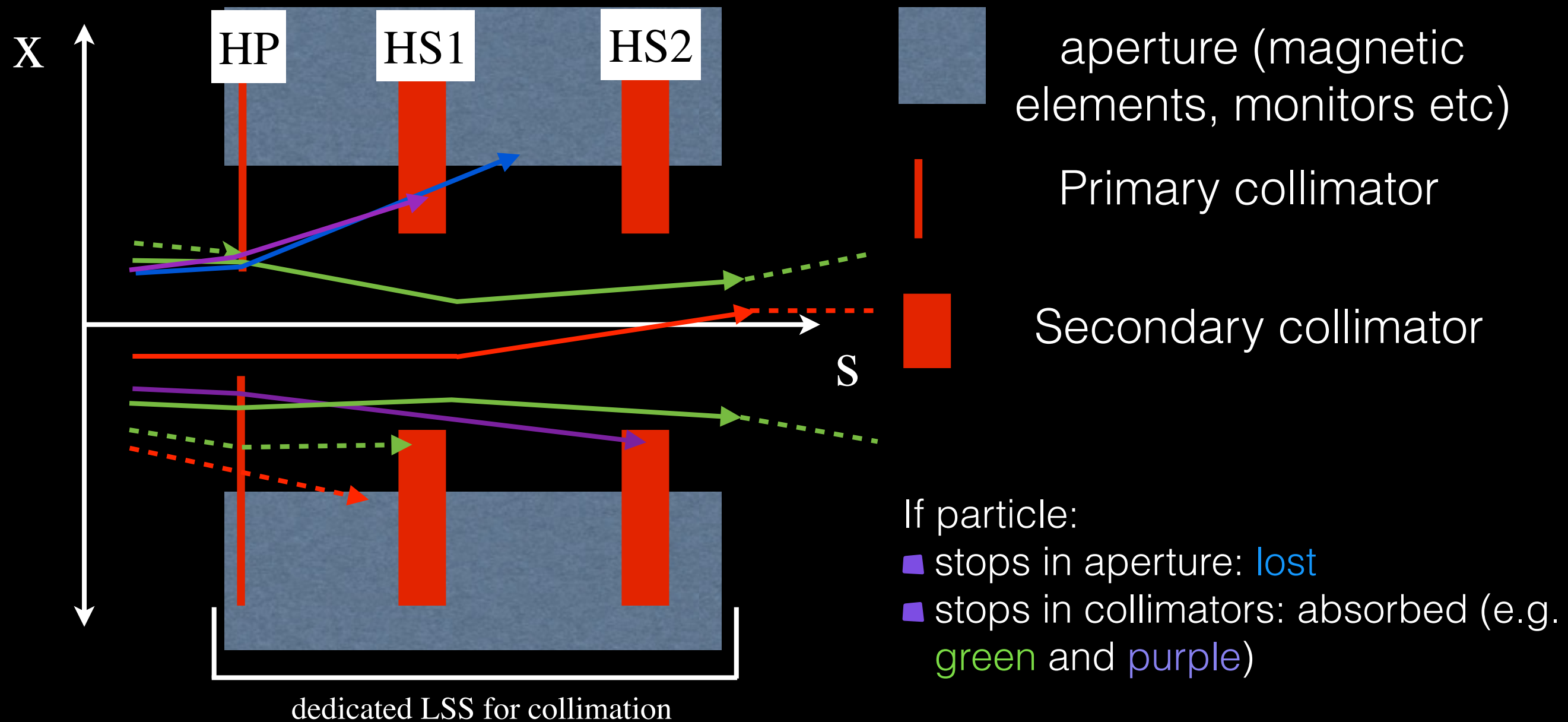


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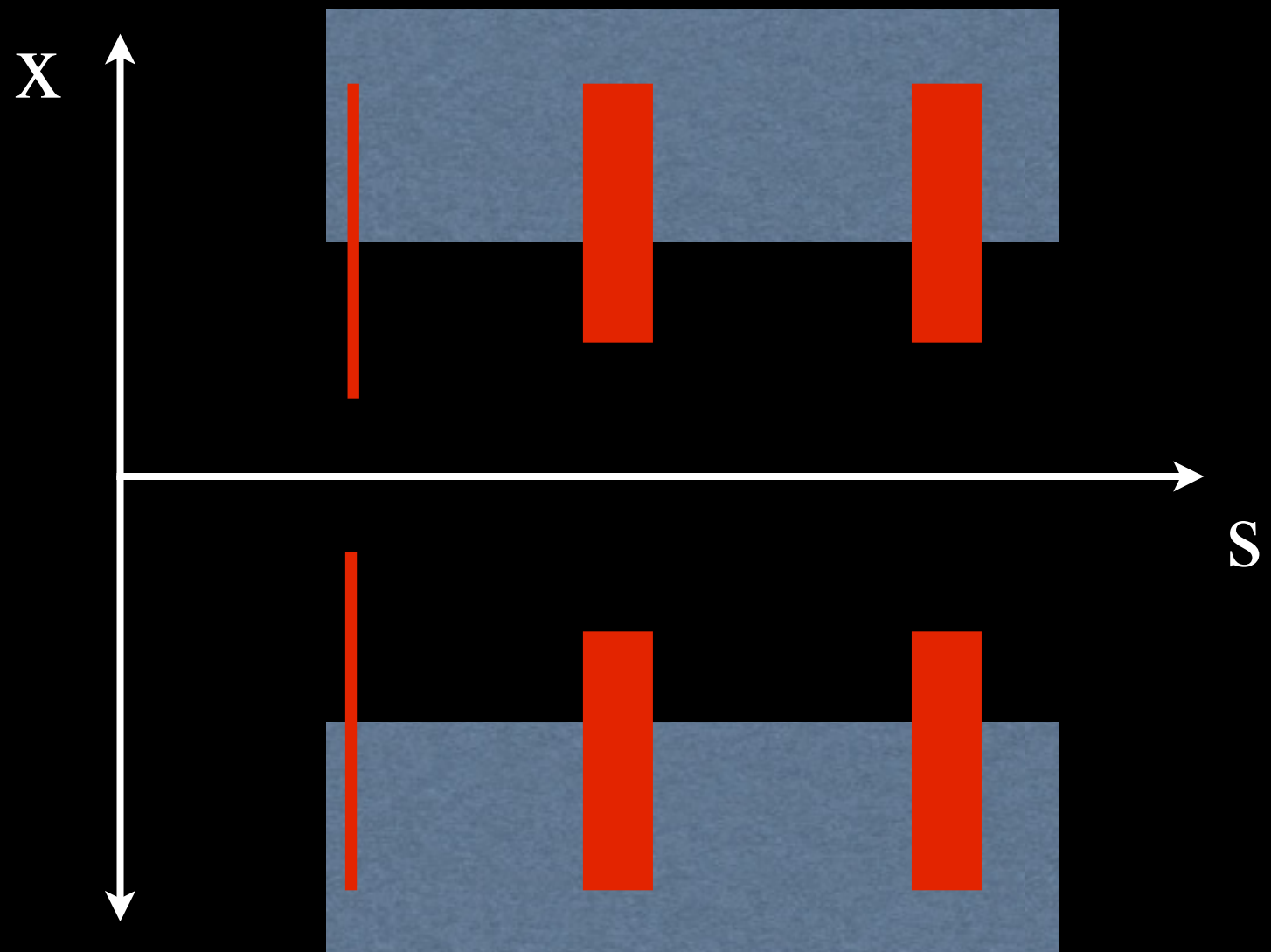
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# Optimising Collimation Efficiency

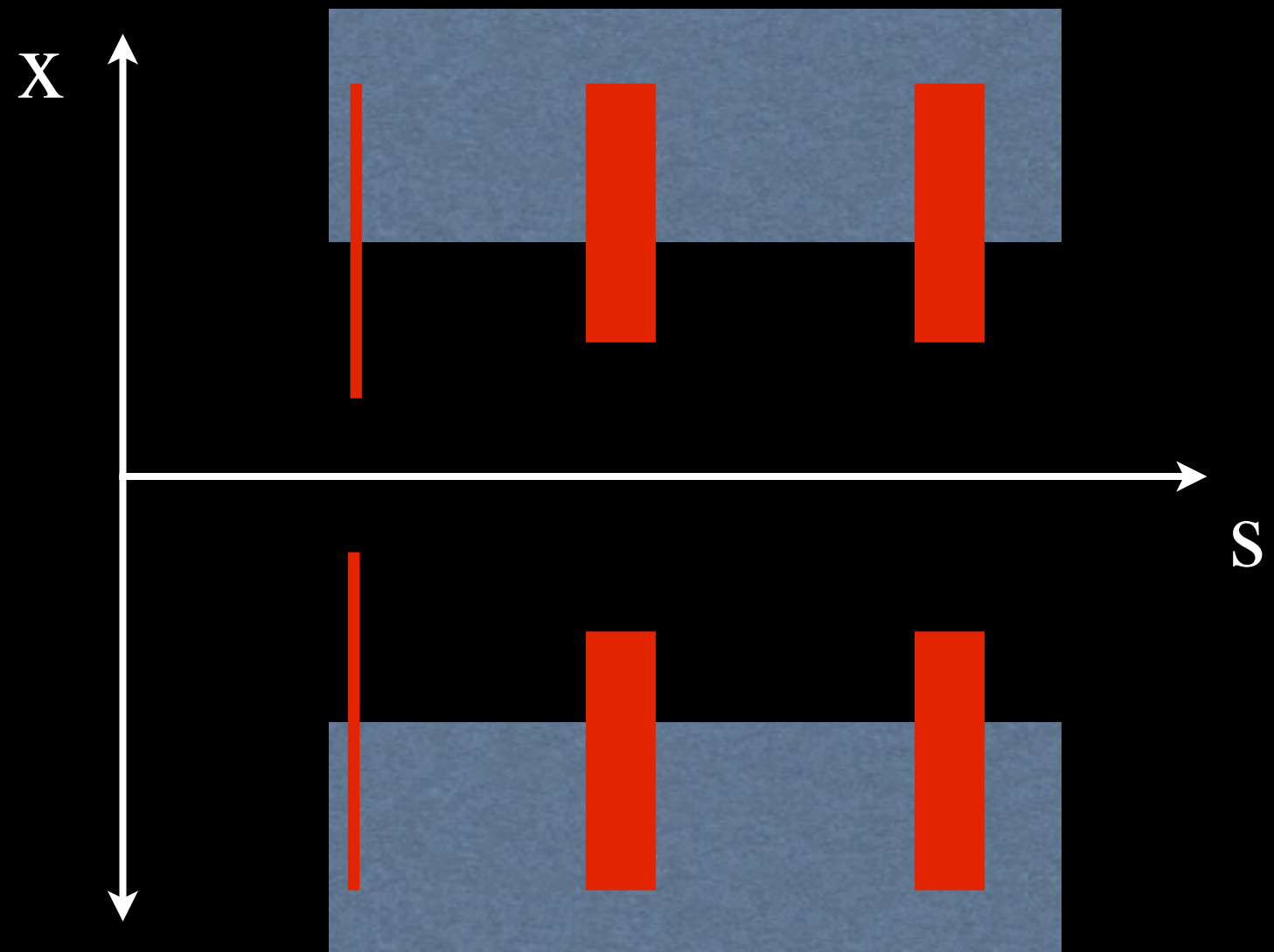
Parameters:



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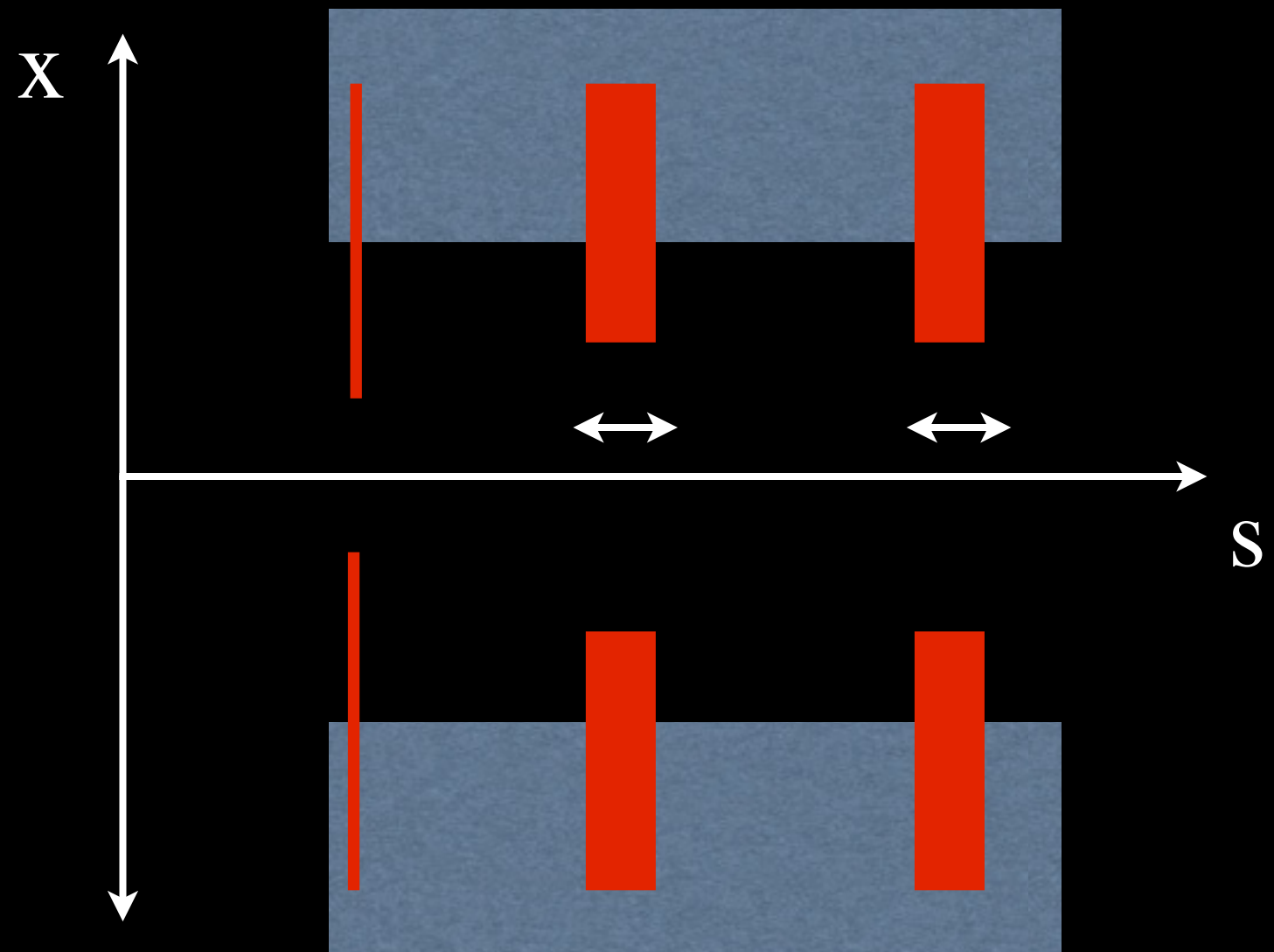
■ collimators thickness



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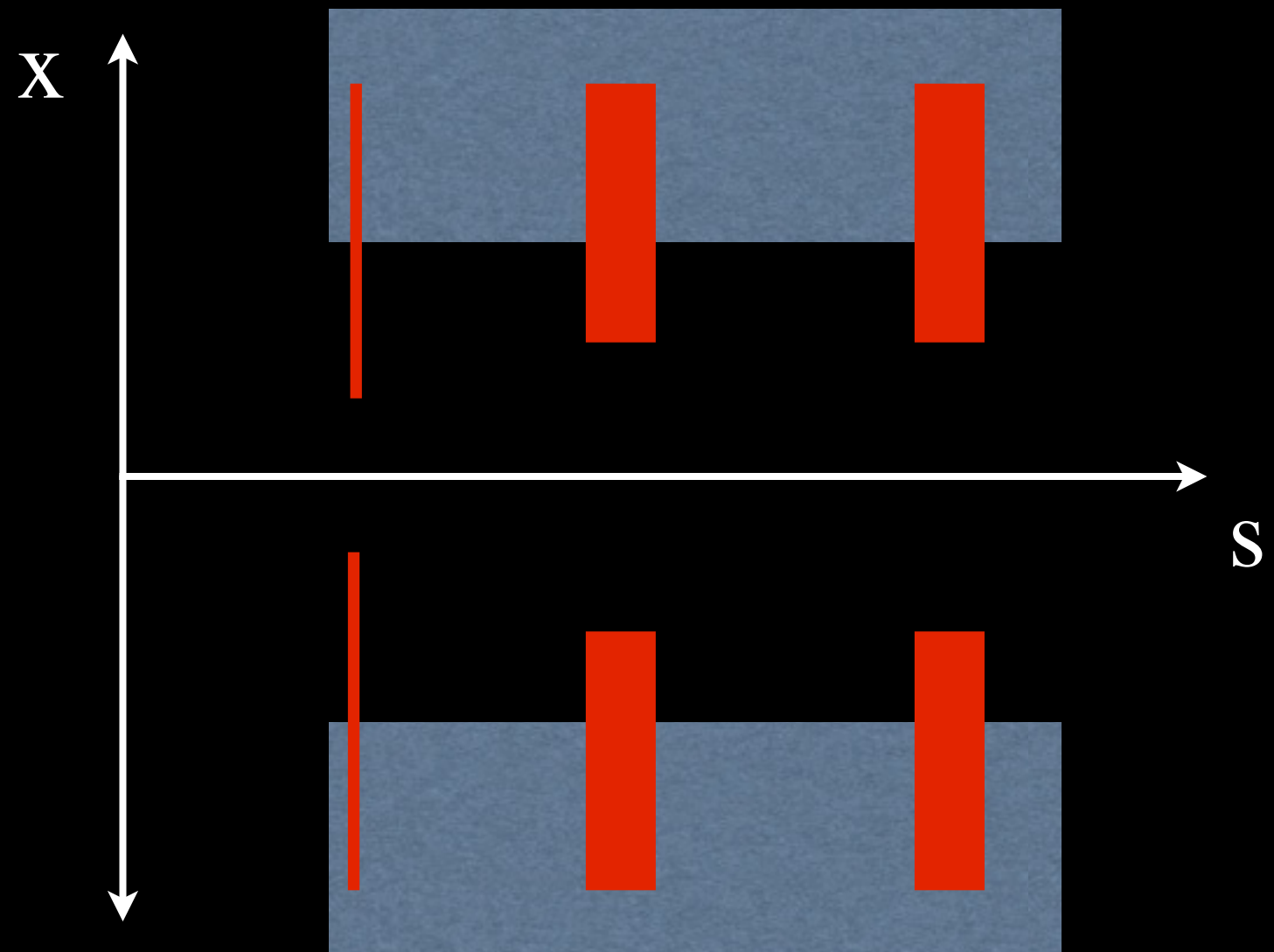
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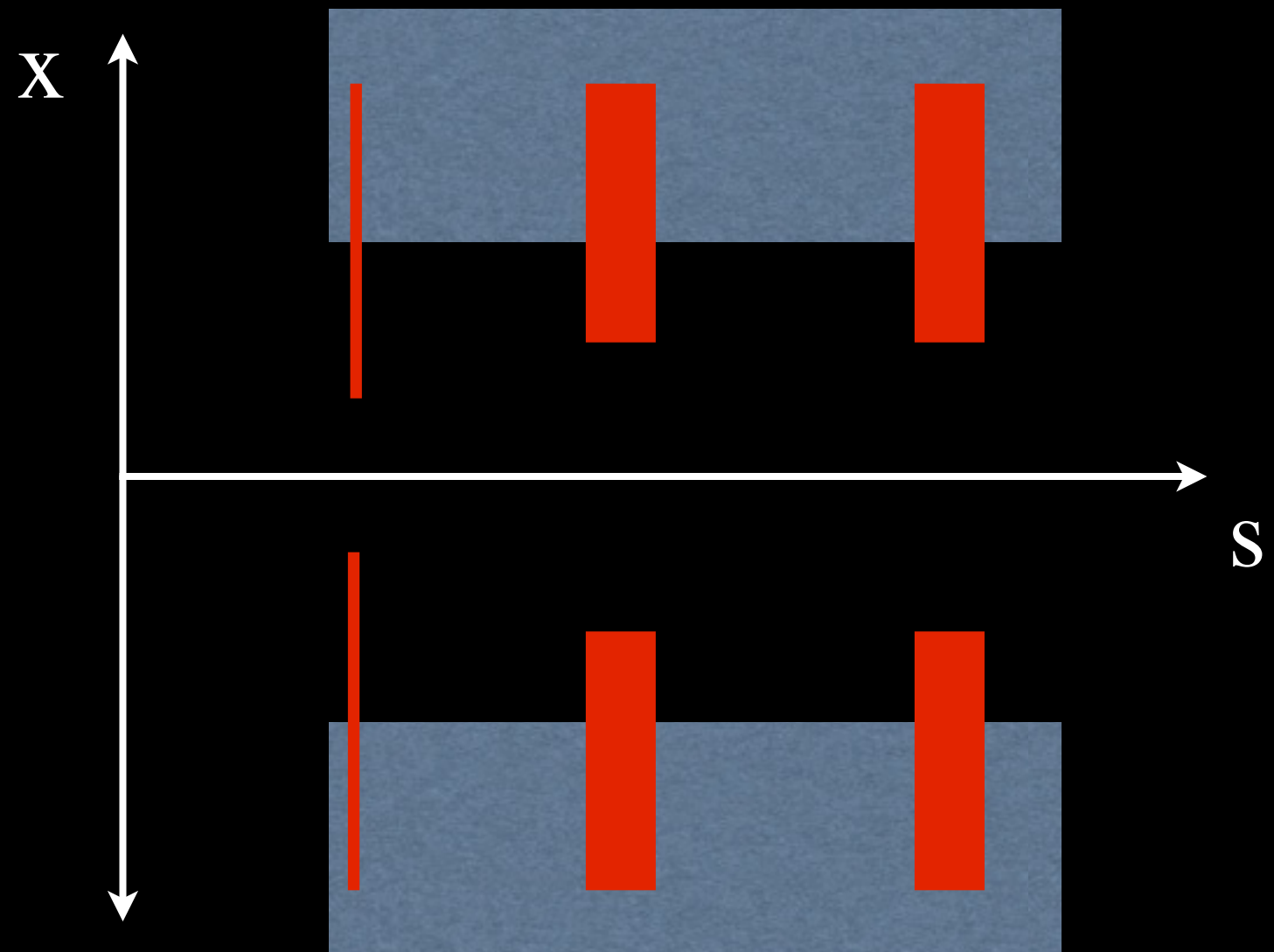
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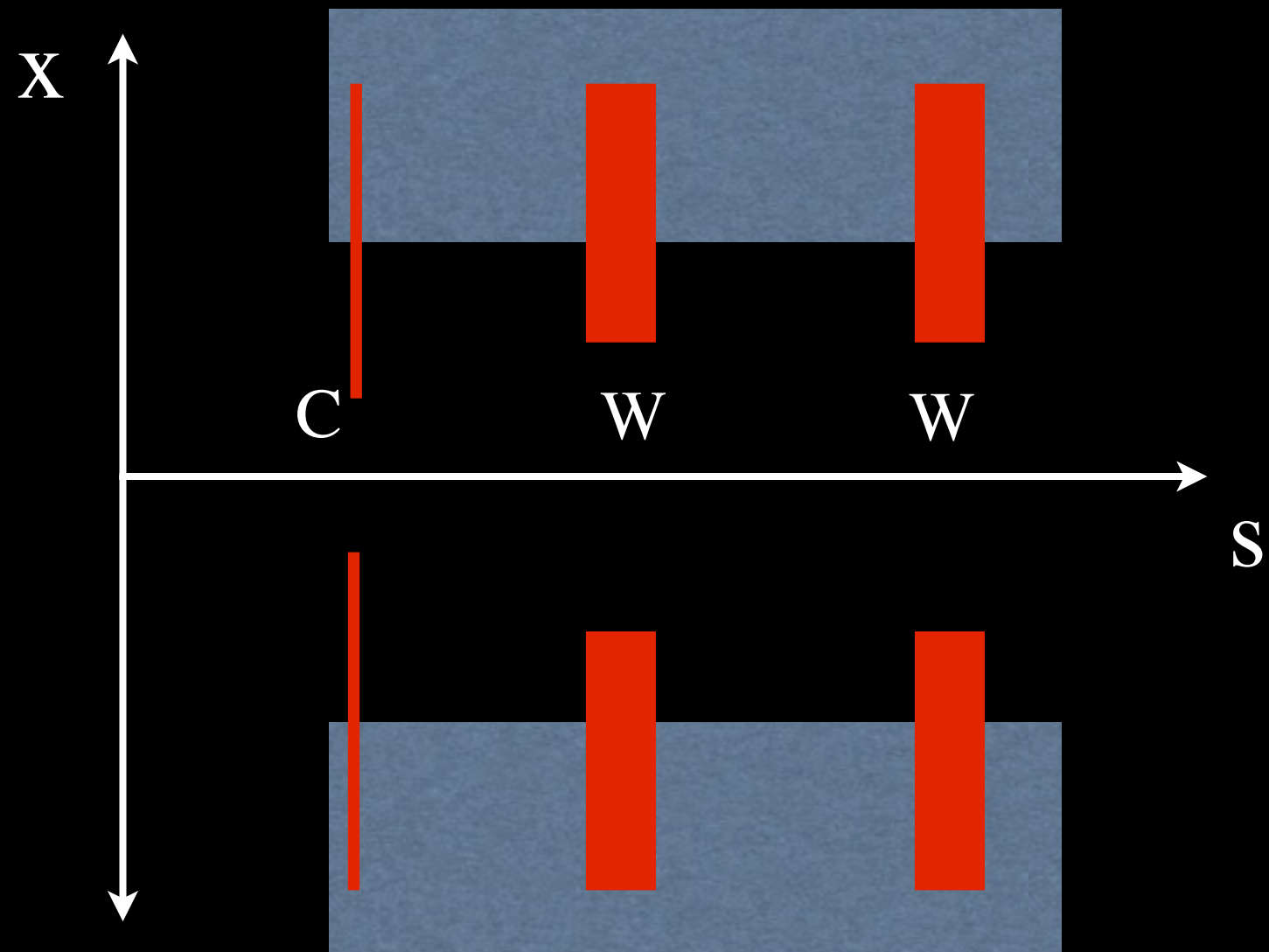
- collimators thickness
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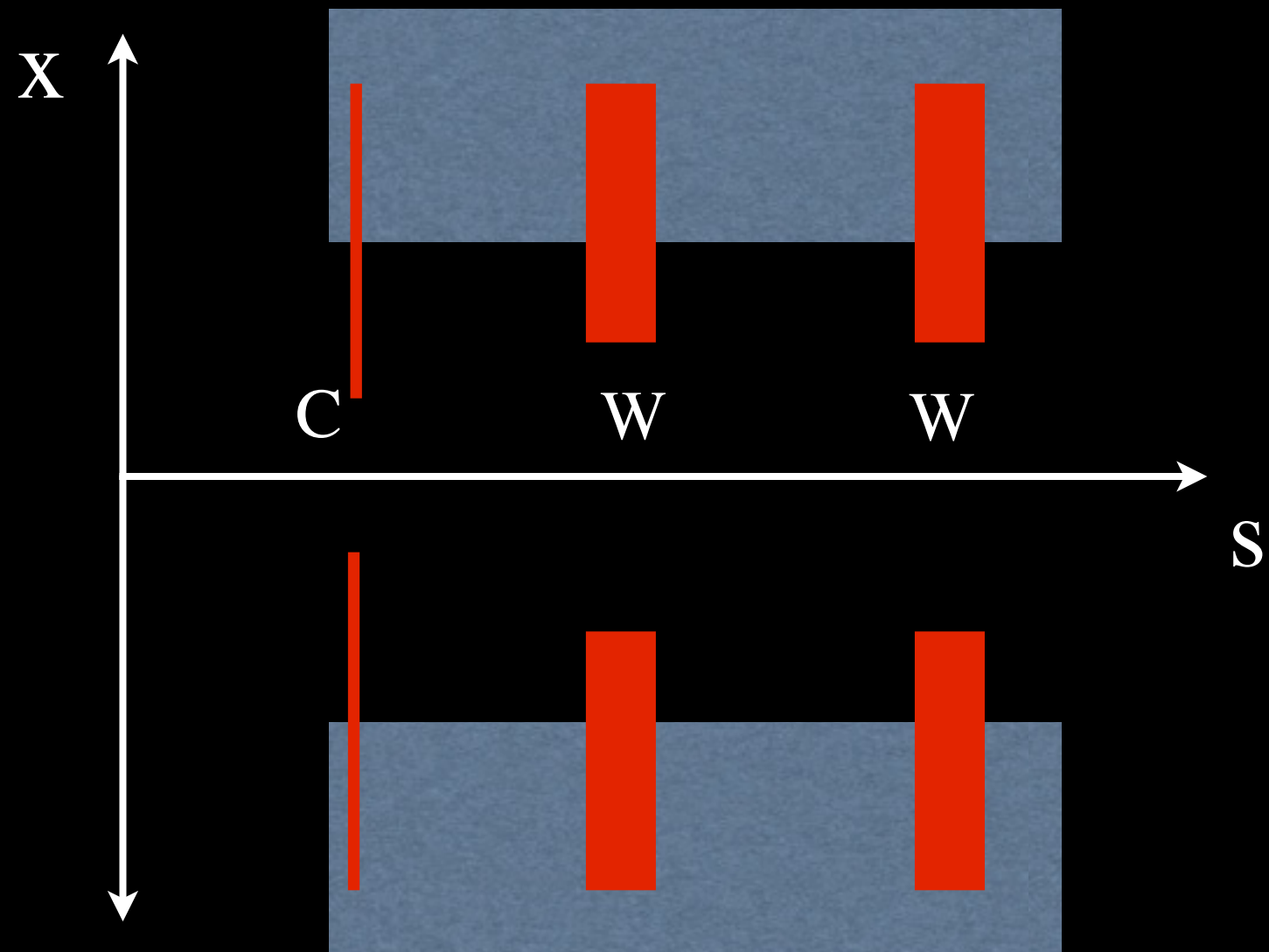




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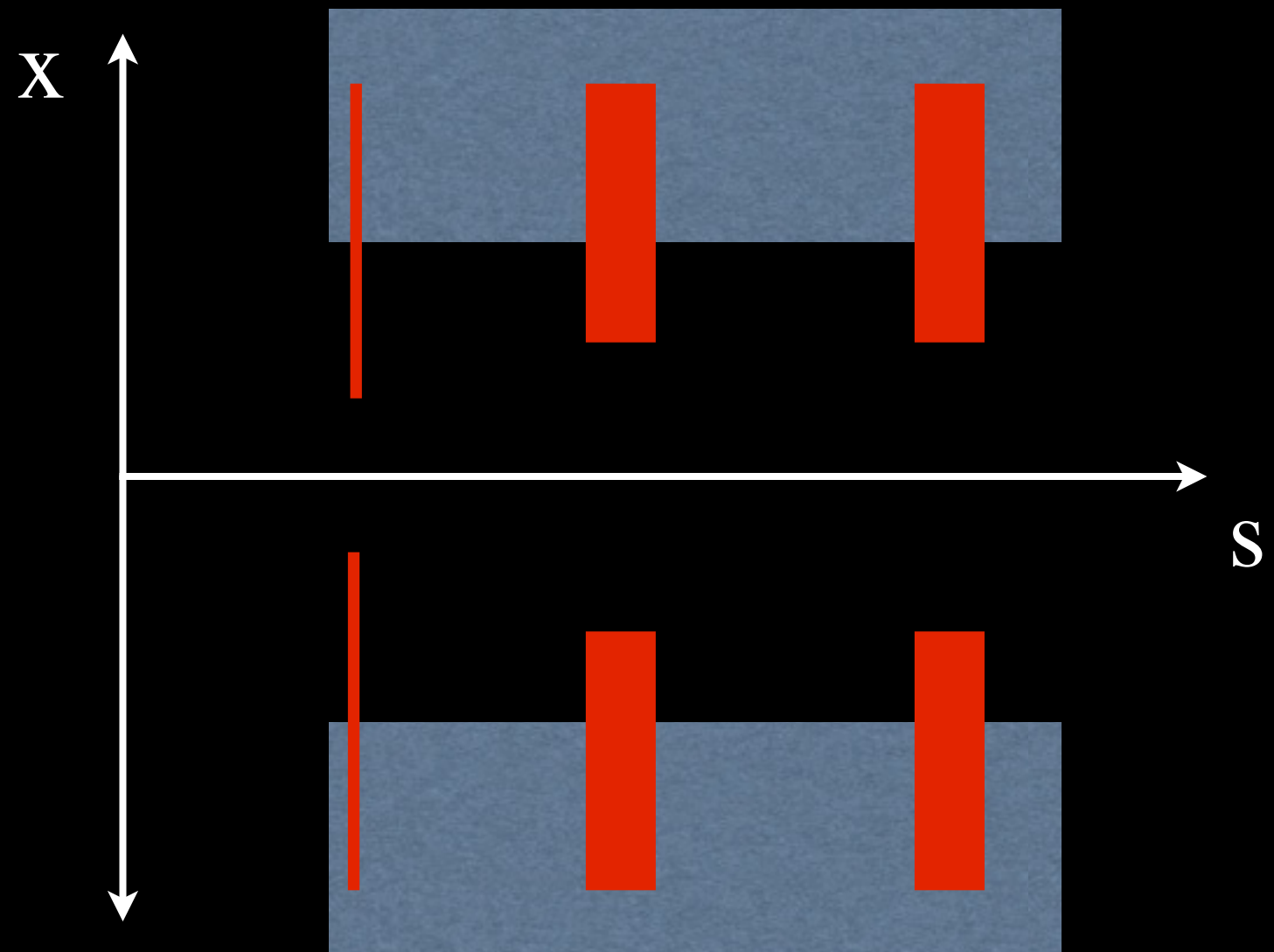
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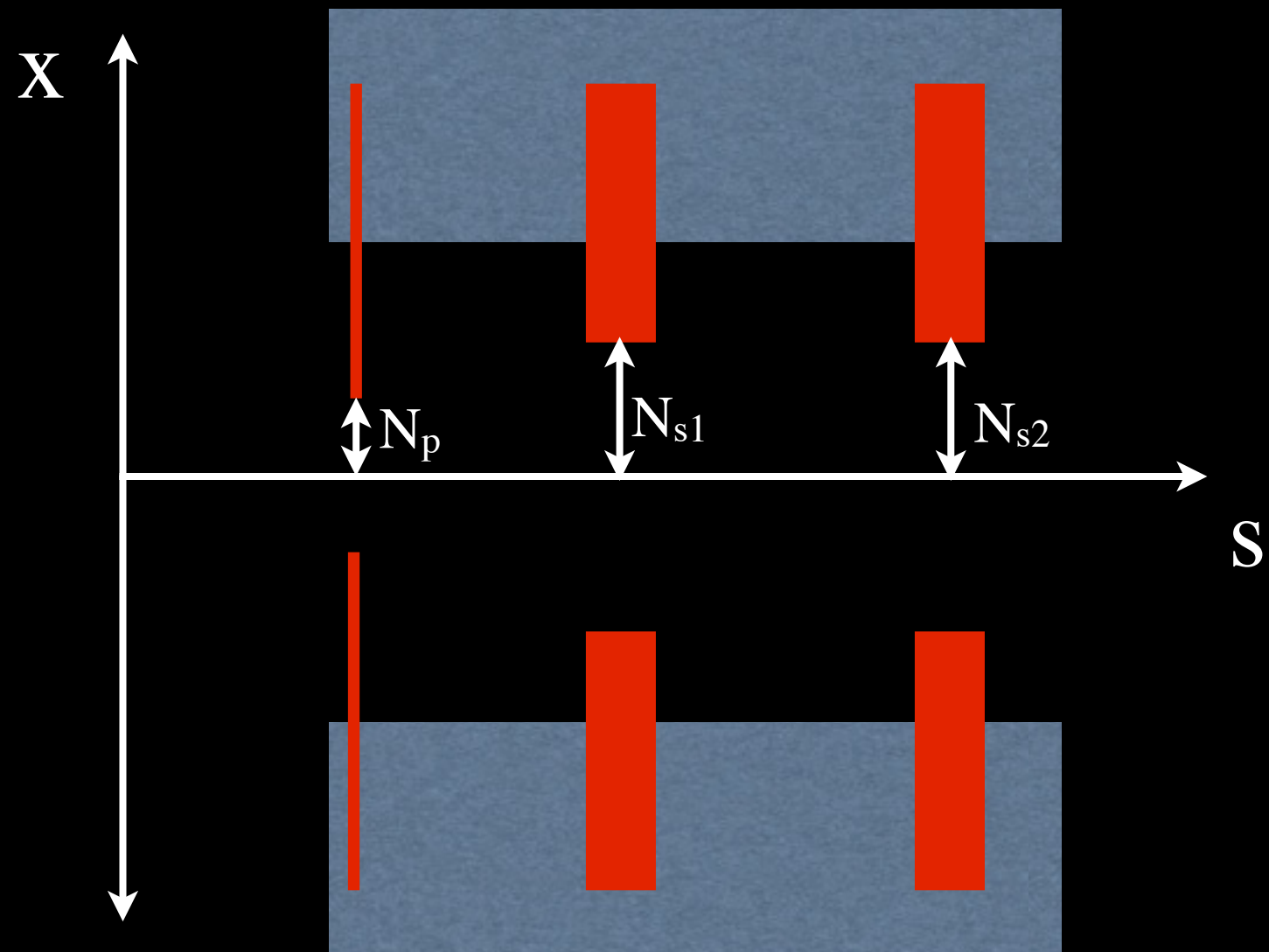
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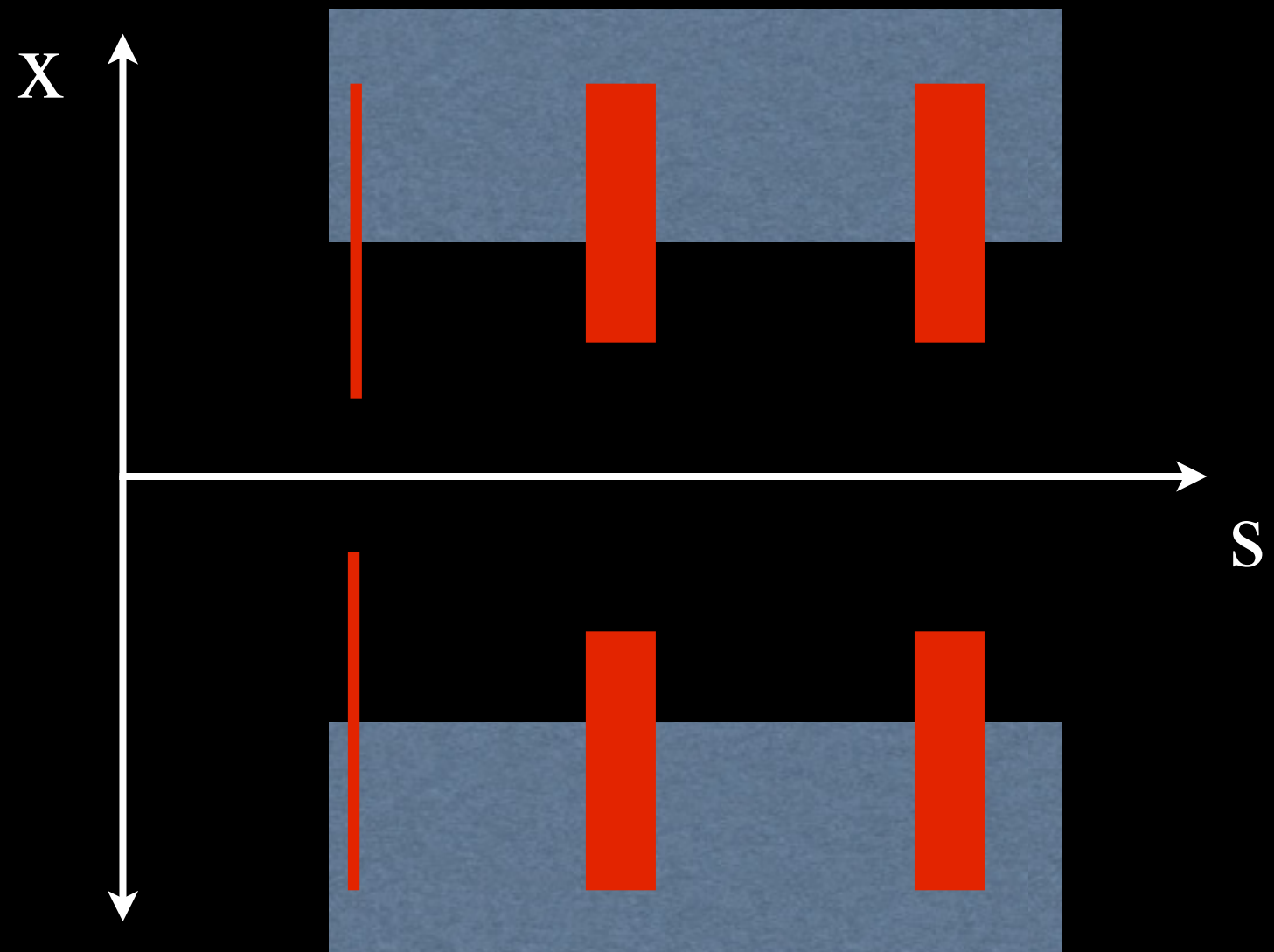
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For different:

# Optimising Collimation Efficiency

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For different:

- beam halo type (H or V)

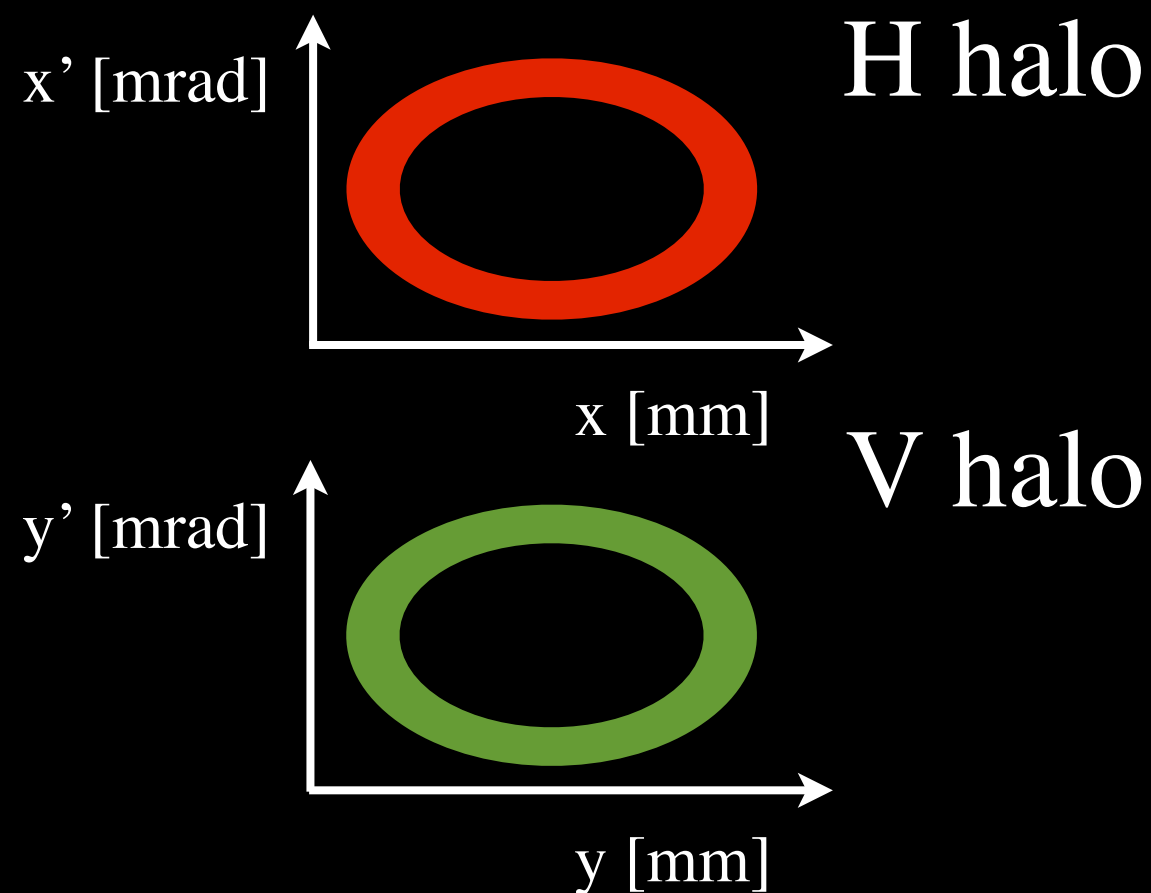
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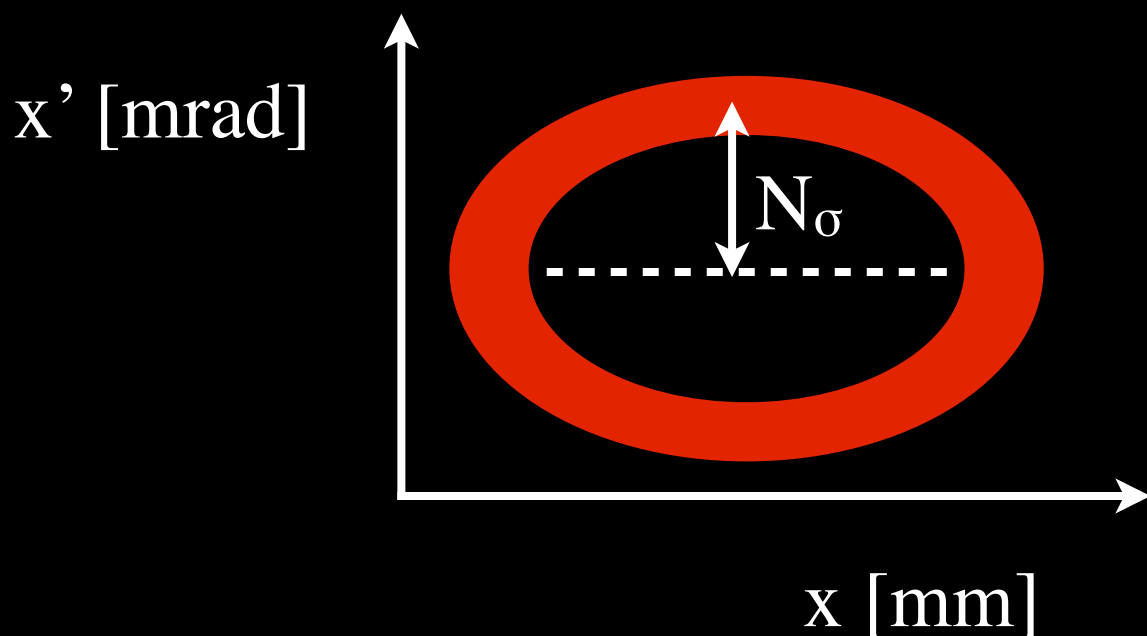
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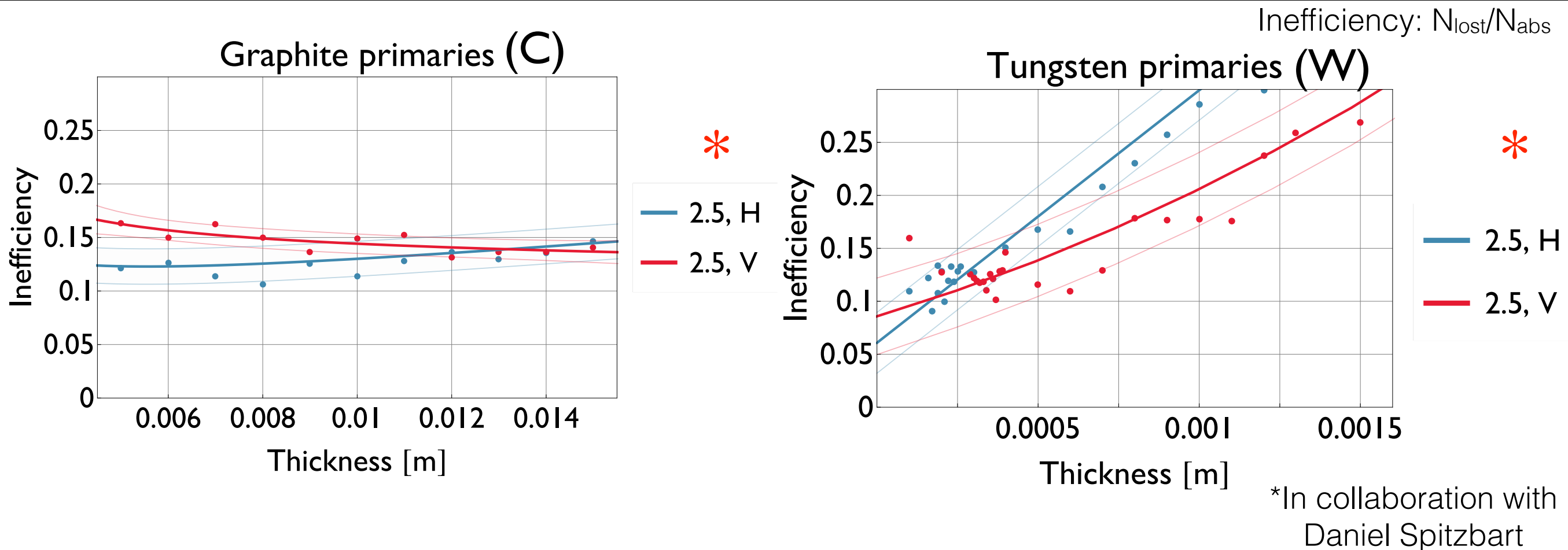
For different:

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H halo



# Thickness and material of primary collimators

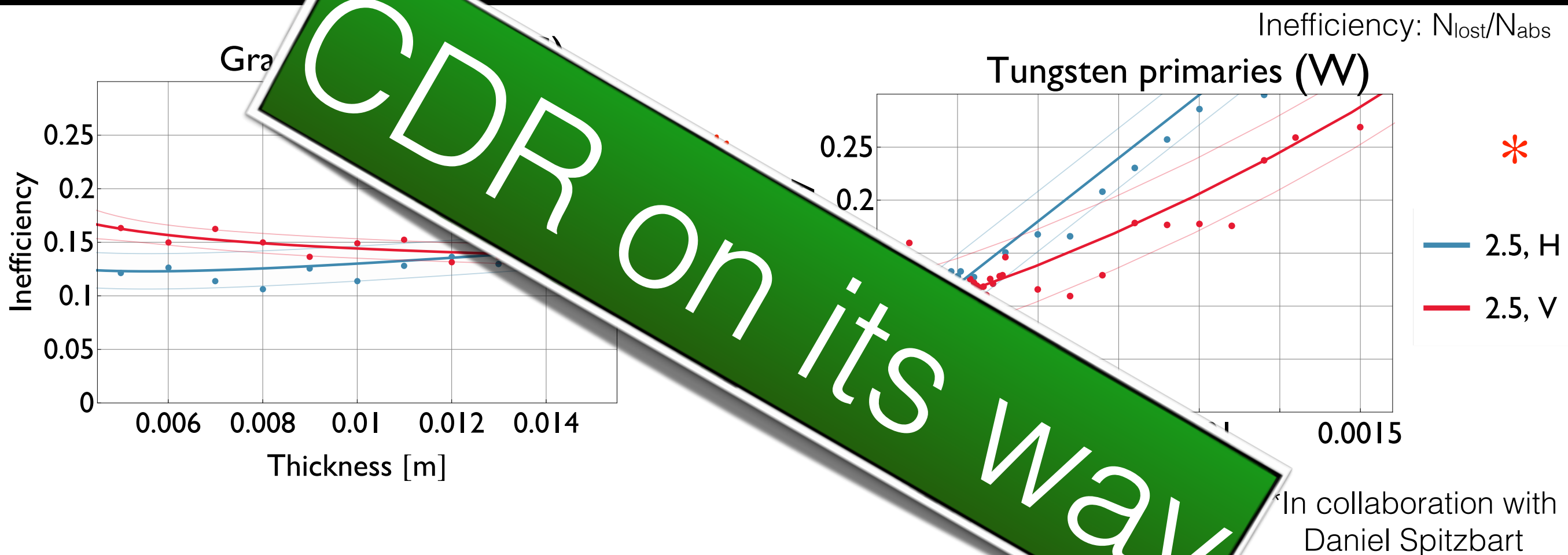


- W: Small inefficiency for very small thickness (feasibility)
- C: Small inefficiency for larger range of thickness

C chosen for primary collimators

W chosen for secondary collimators

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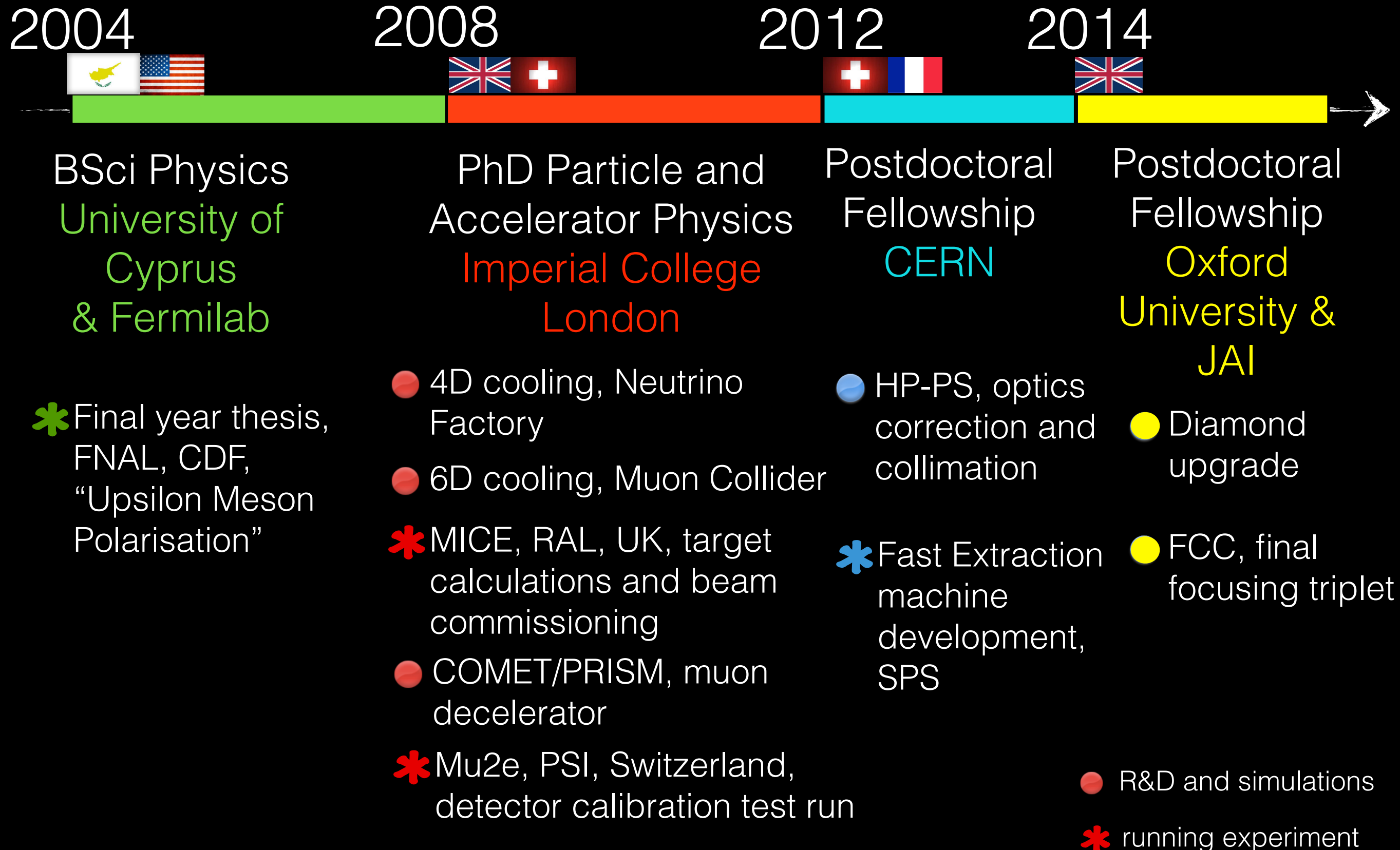


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# Timeline



# Diamond Light Source

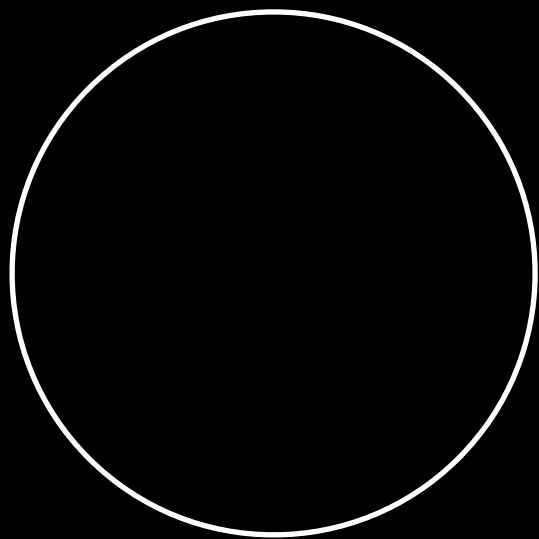
# Diamond Light Source

- High energy electrons can emit extremely bright and coherent beams of high energy photons via synchrotron radiation; powerful microscopes: the higher the energy the better the resolution ( $E=h/\lambda$ )



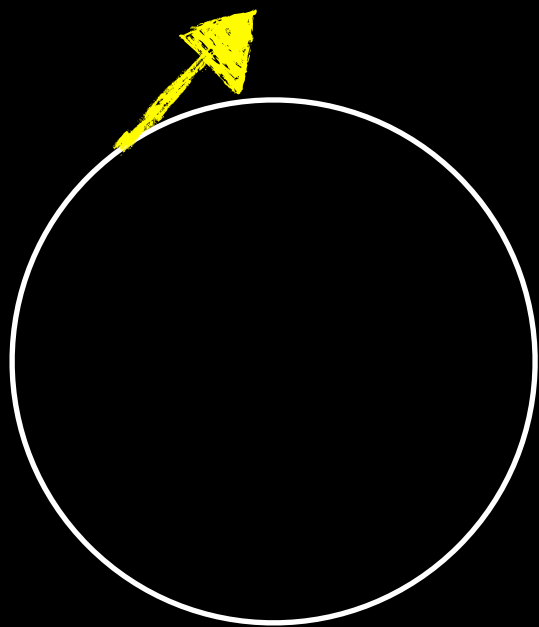
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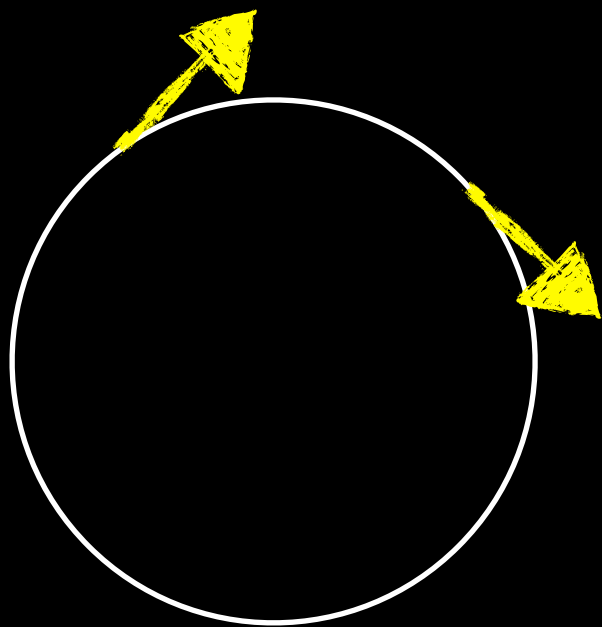
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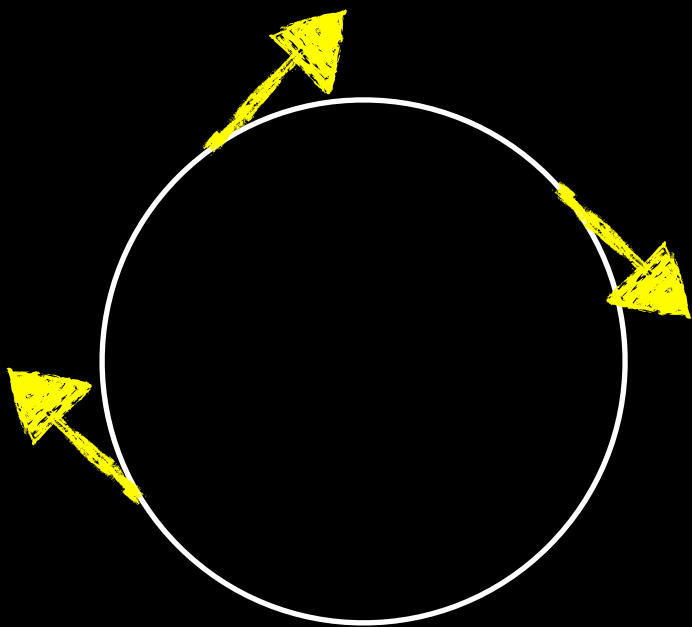
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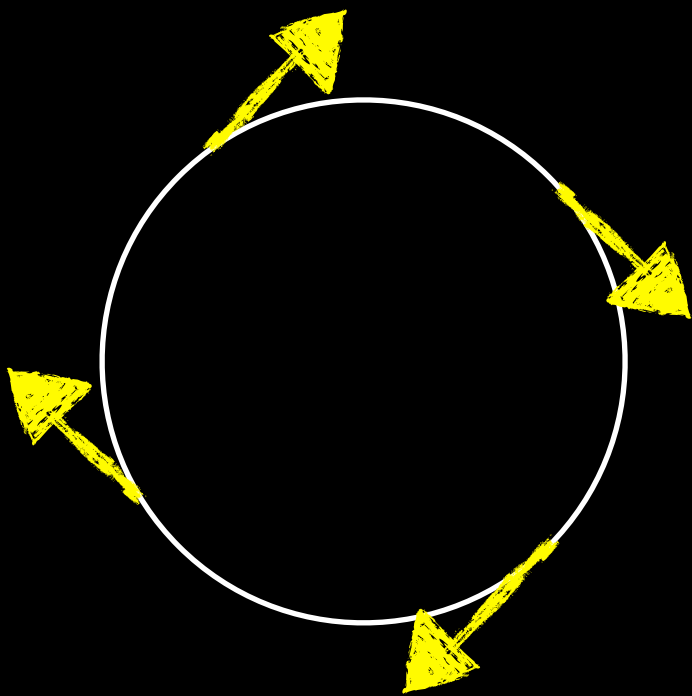
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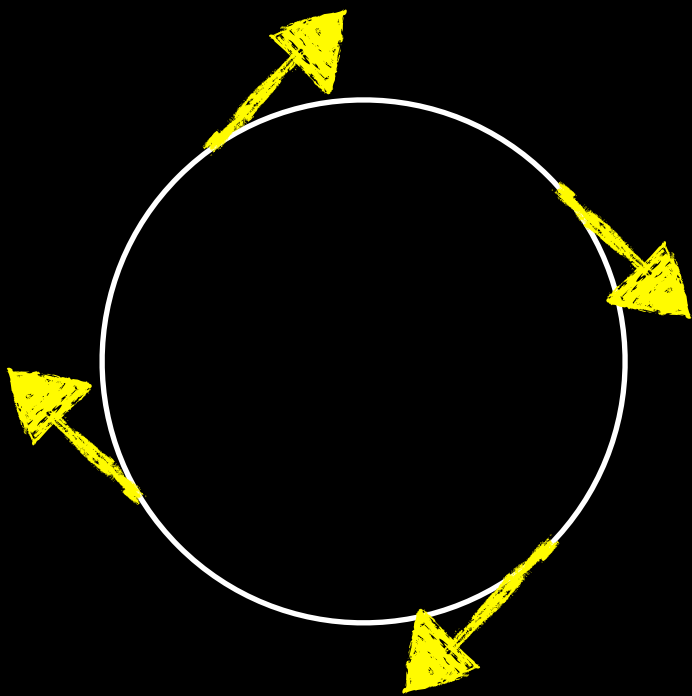
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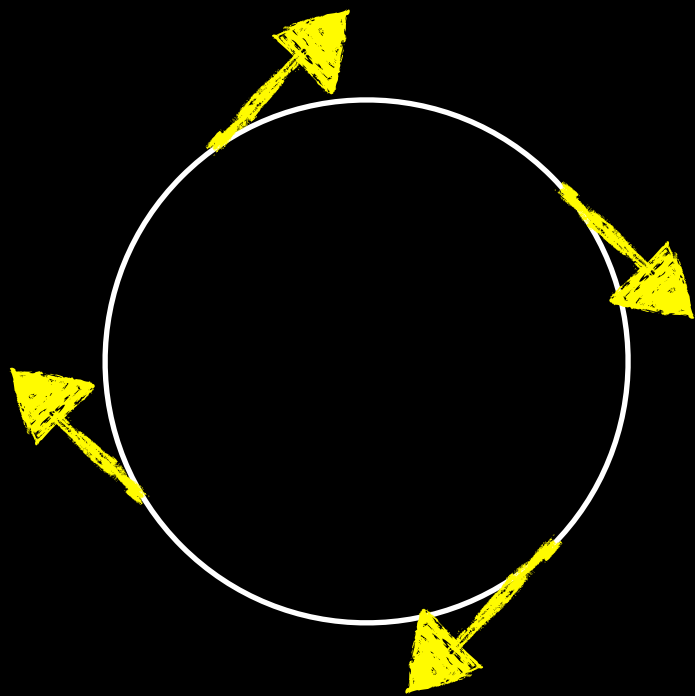
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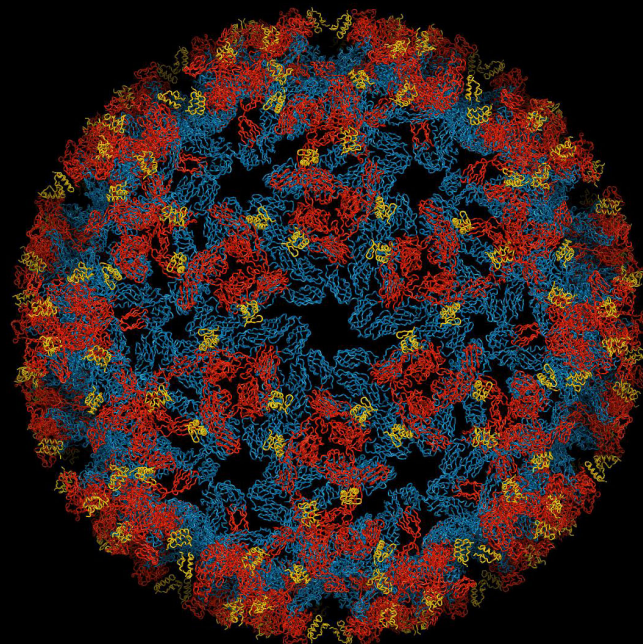
- Numerous uses in the study of atomic structure, chemistry, condensed matter physics, biology, and technology

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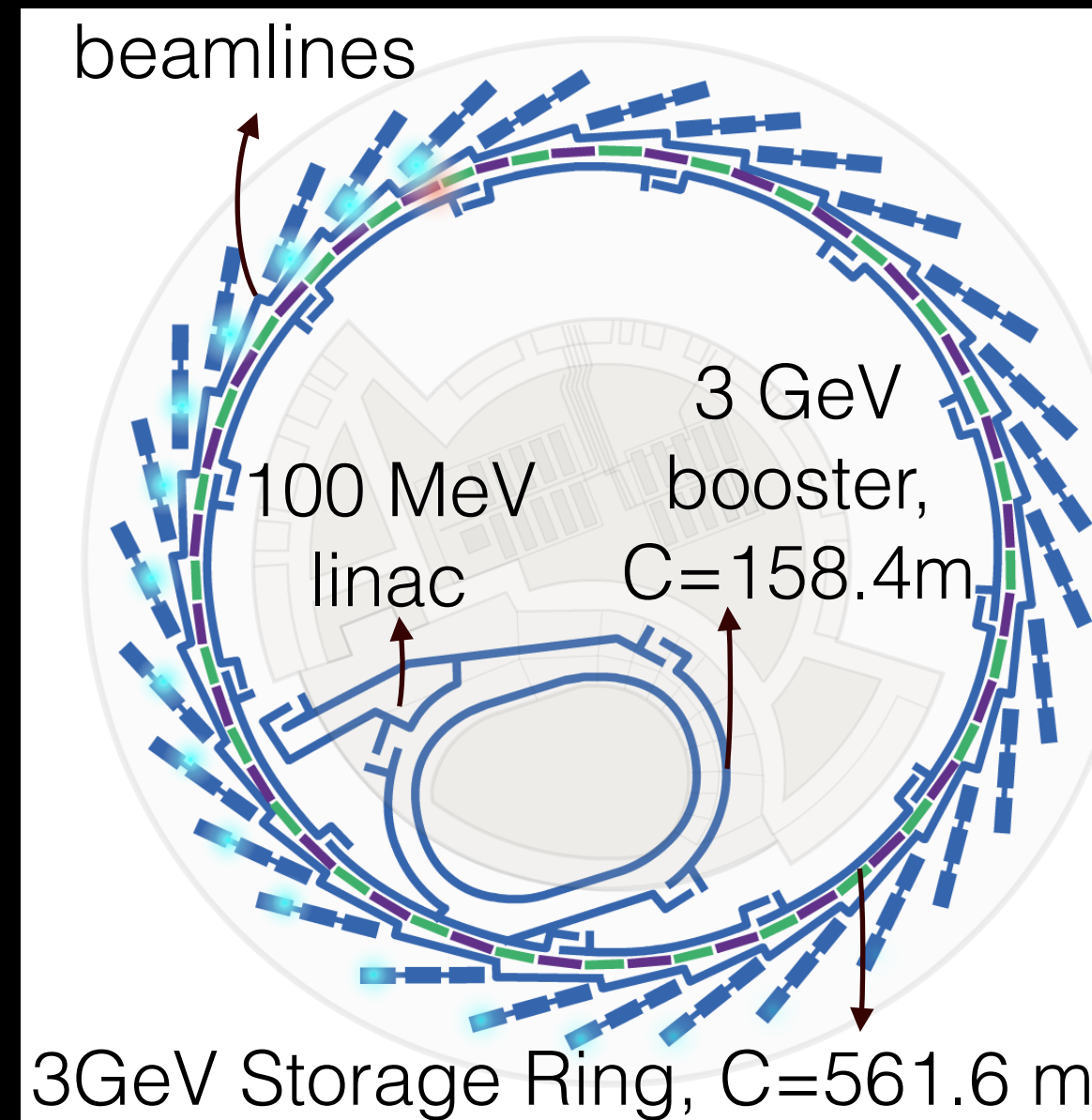


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Protein modelling: Synchrotron light allows scientists to solve 3D structure of proteins e.g. the Chikungunya virus. Image credit: Voss et al., Nature (2010) 468, 709 (via Synchrotron Soleil, France)

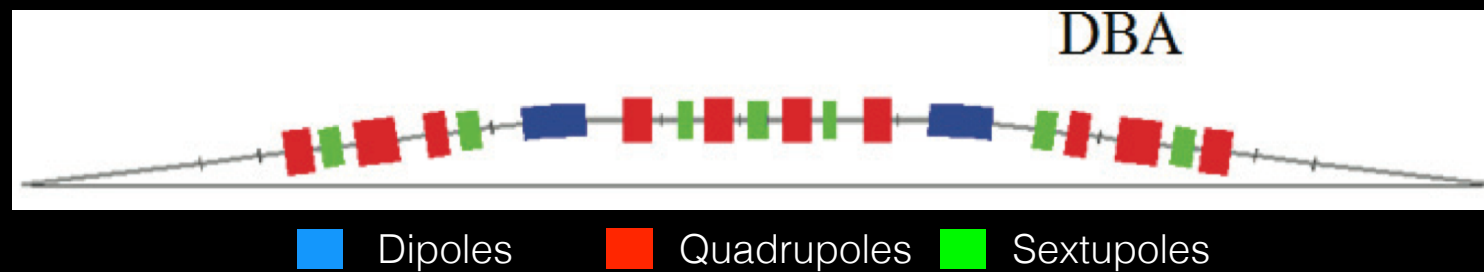
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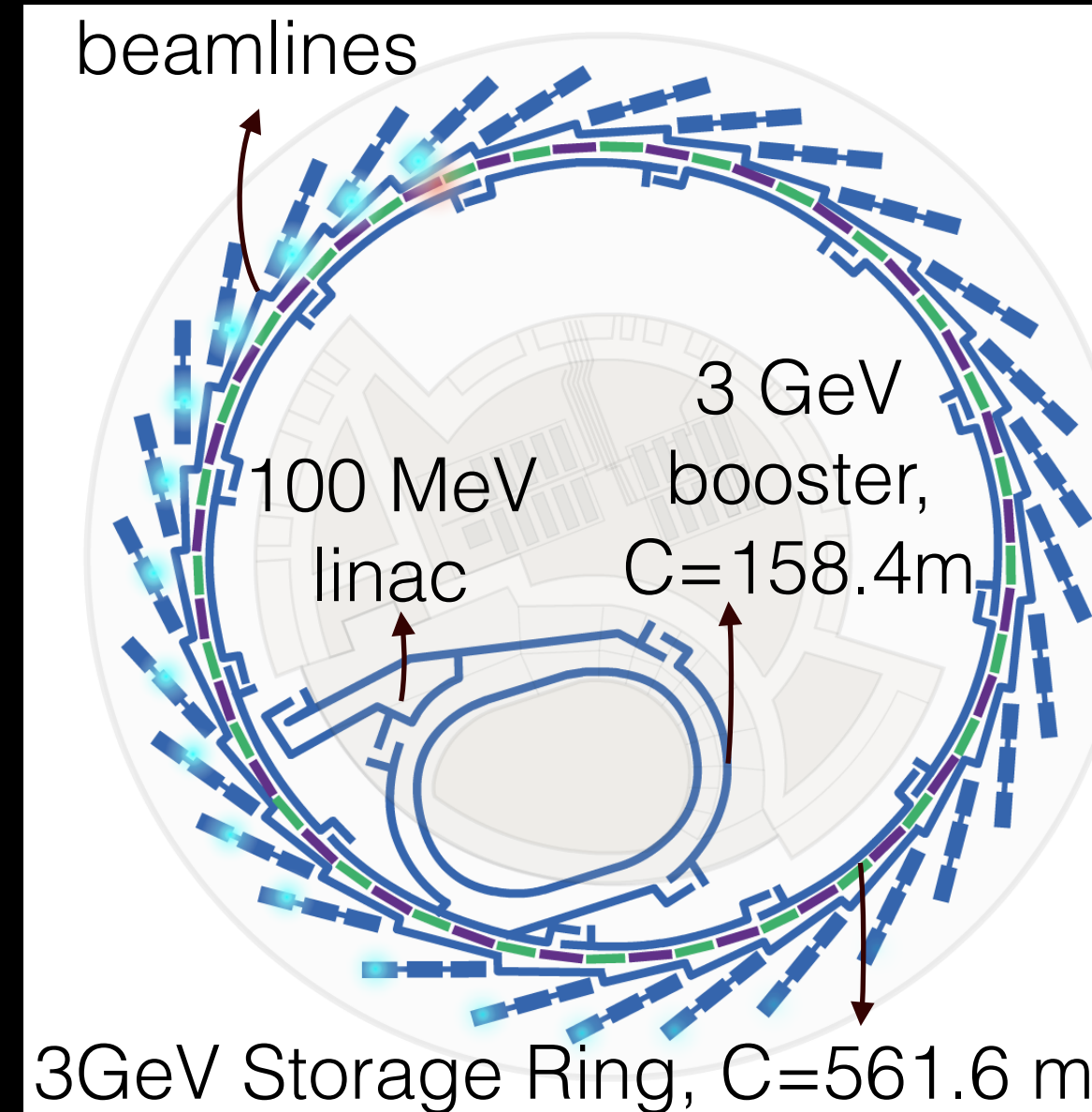


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- Diamond Light Source (DLS) consists of 24 cells; straight sections+Double Bend Achromat (DBA)

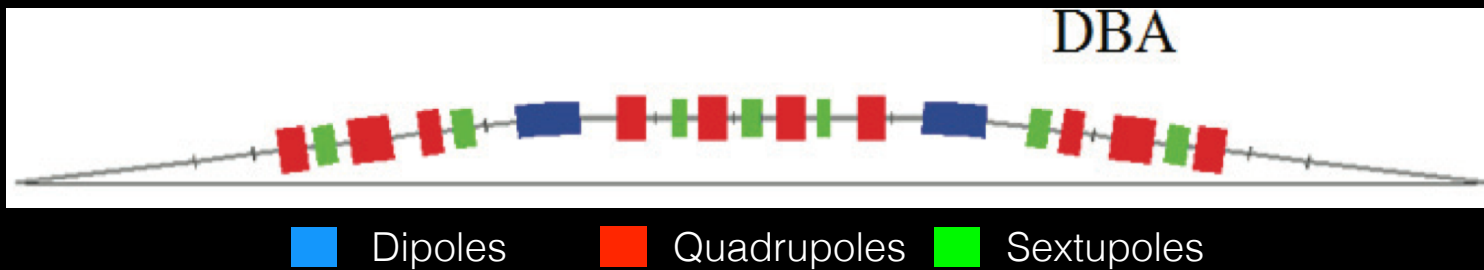


DBA: bending section uses 2 dipoles; zero dispersion at entrance and exit of cell



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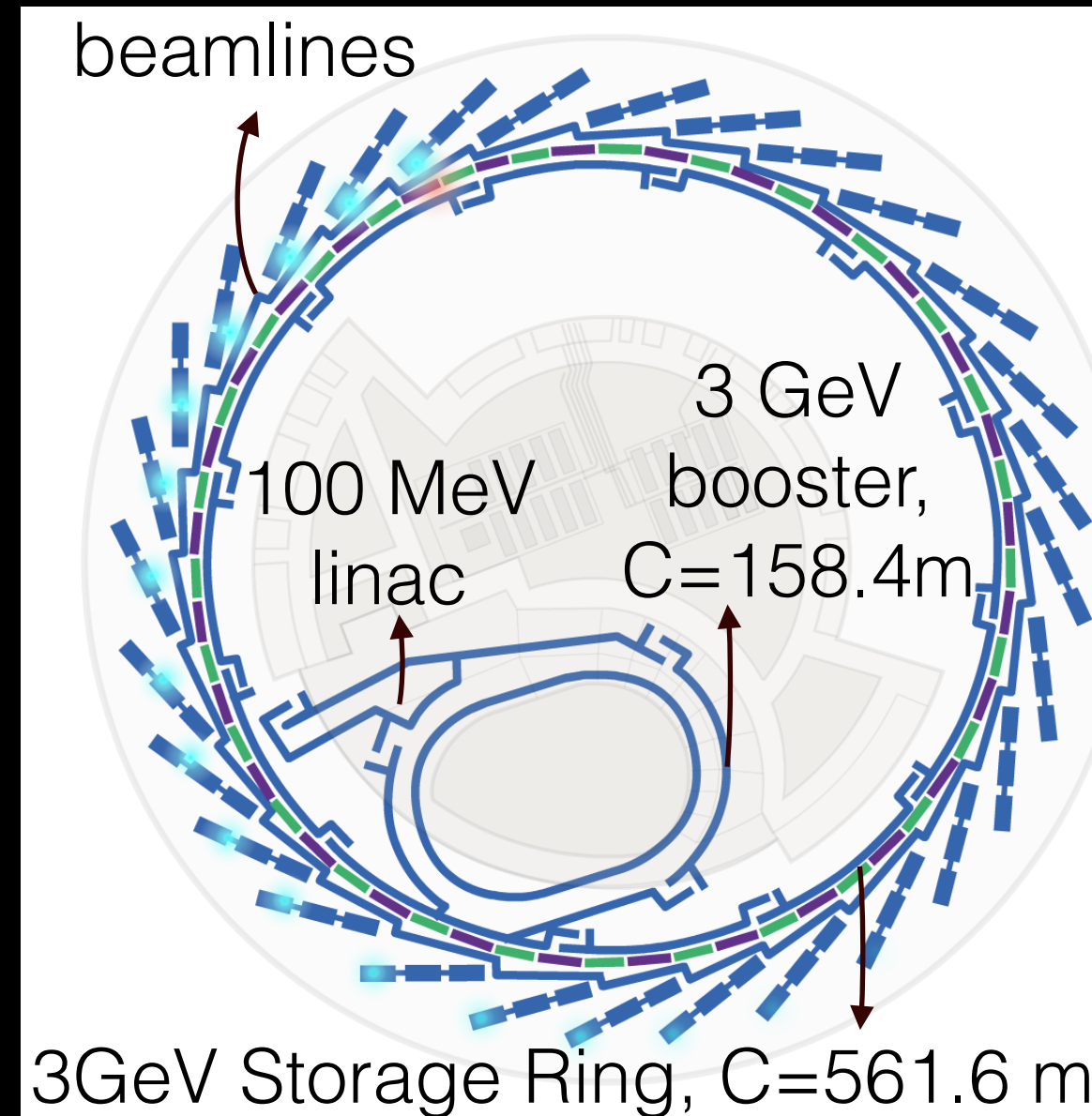
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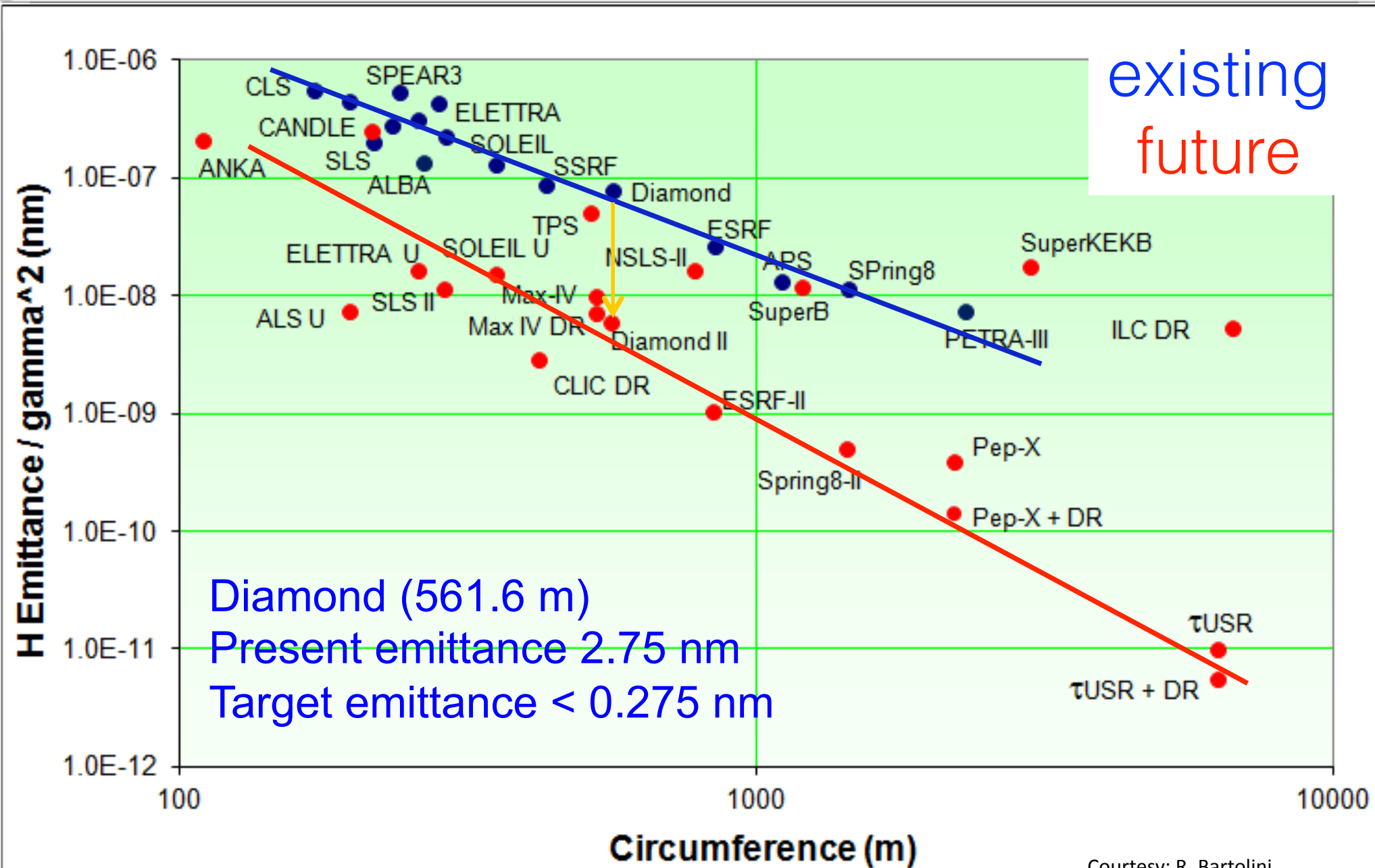
DBA: bending section uses 2 dipoles; zero dispersion at entrance and exit of cell

Operates since 2007 with nominal parameters:

- 2.7 nm H emittance (2nd best behind Advance Photon Source)
- 8 pm V emittance in 2012 (<2 pm world record in fall 2009)
- 300 mA reached in 2008



# Motivation



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- Upgrades SR facilities aim lower  $\varepsilon$ ; increase brilliance and transverse coherent factor

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Original DBA



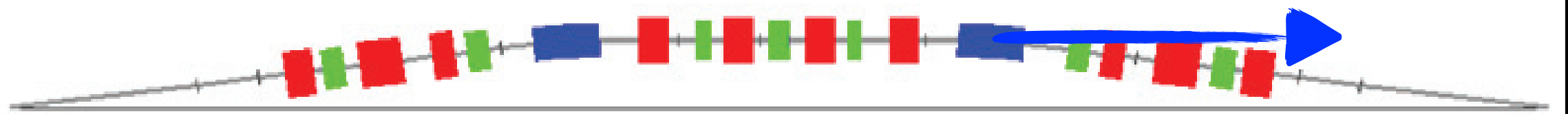
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Original DBA

BM Beamline





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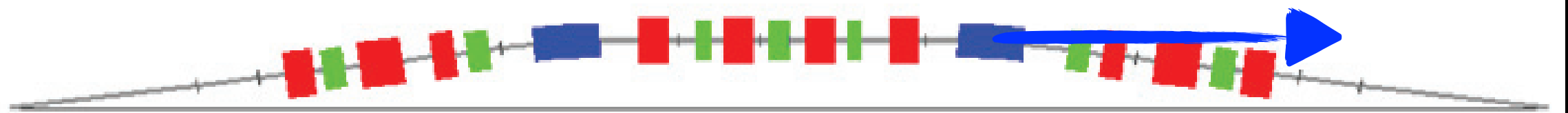
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Original DBA

BM Beamline

DDBA

3.35 m

3.4 m

3.35 m

DDBA

■ Dipole ■ Quadrupole ■ Sextupole

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BM Beamline

DDBA

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3.4 m

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Modified 4BA (DDBA): **additional straight** in mid of arc +  $\varepsilon$  reduction!

Origin

**Double-DBA (DDBA or modified 4BA):  
doubles capacity and reduces emittance by factor of 10!**

DDBA

3.35 m

3.4 m

ID Beamline

3.35 m

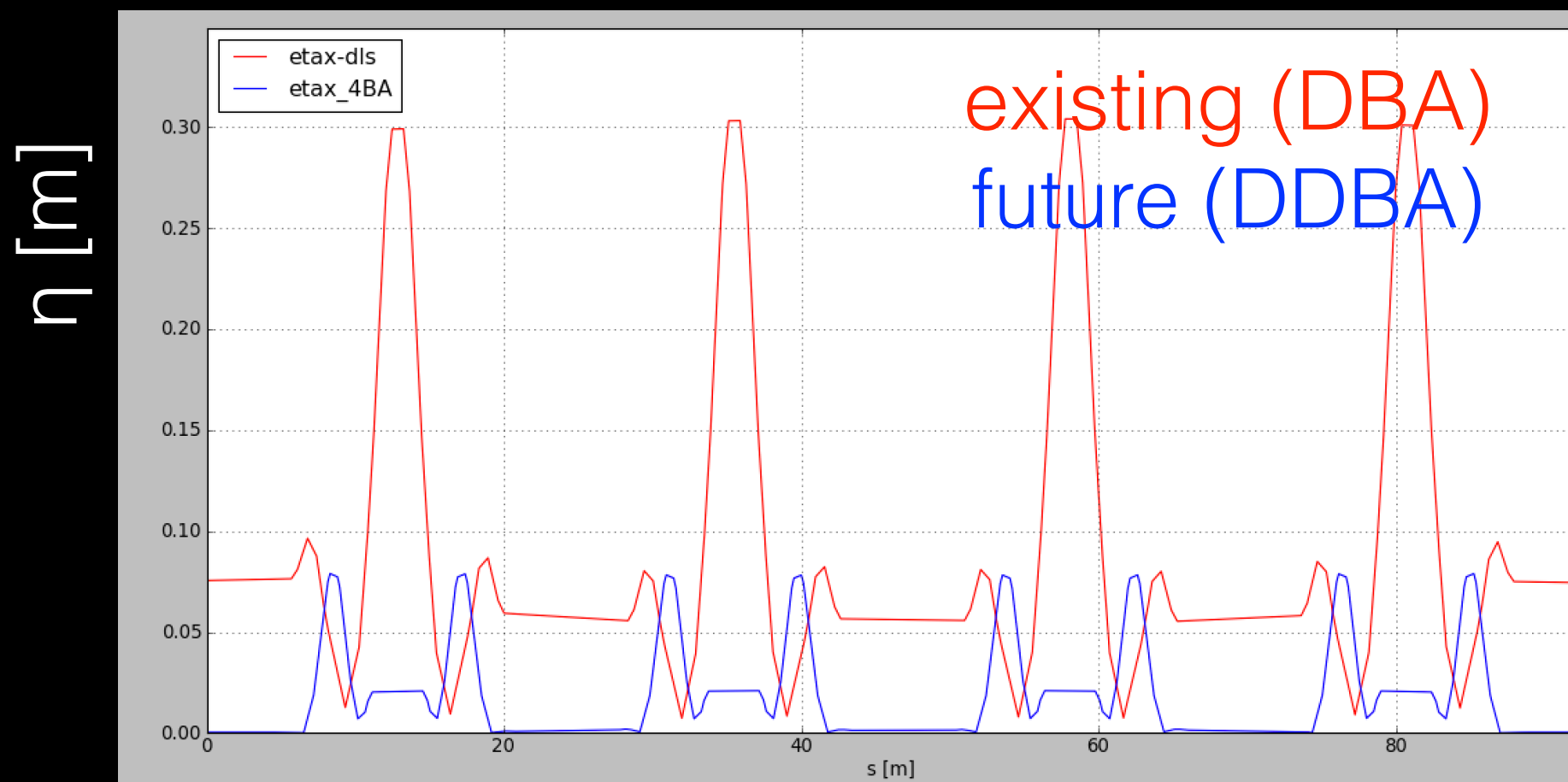
Insertion Device

DDBA

■ Dipole ■ Quadrupole ■ Sextupole

# Upgrade

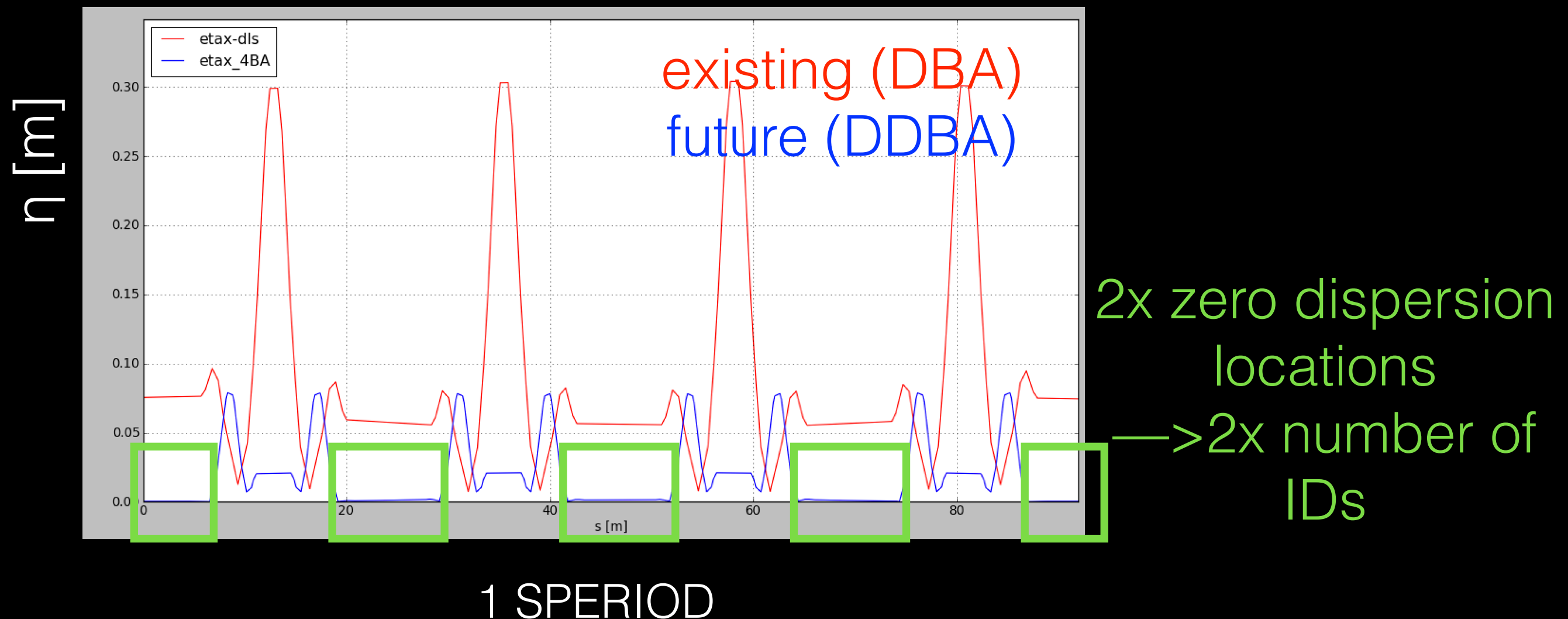
Full lattice upgrade: replacement of existing 24 DBA cells with 24 DDBA cells



1 PERIOD

# Upgrade

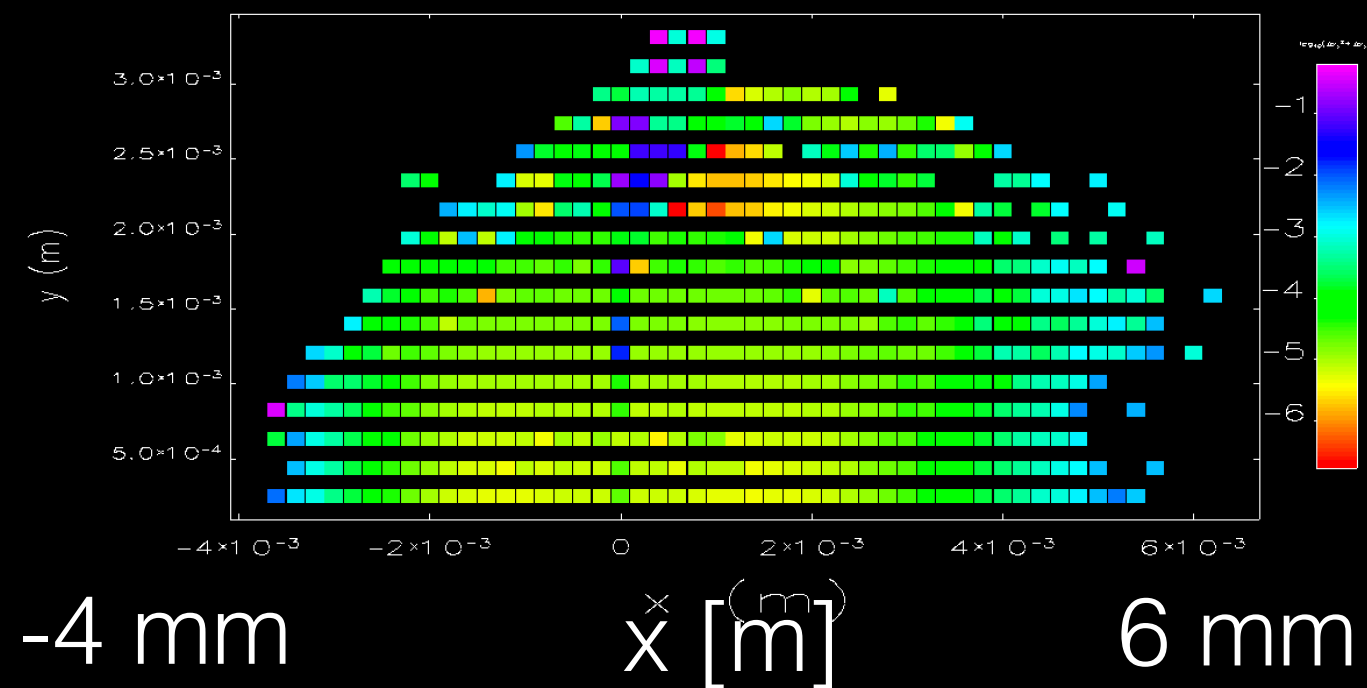
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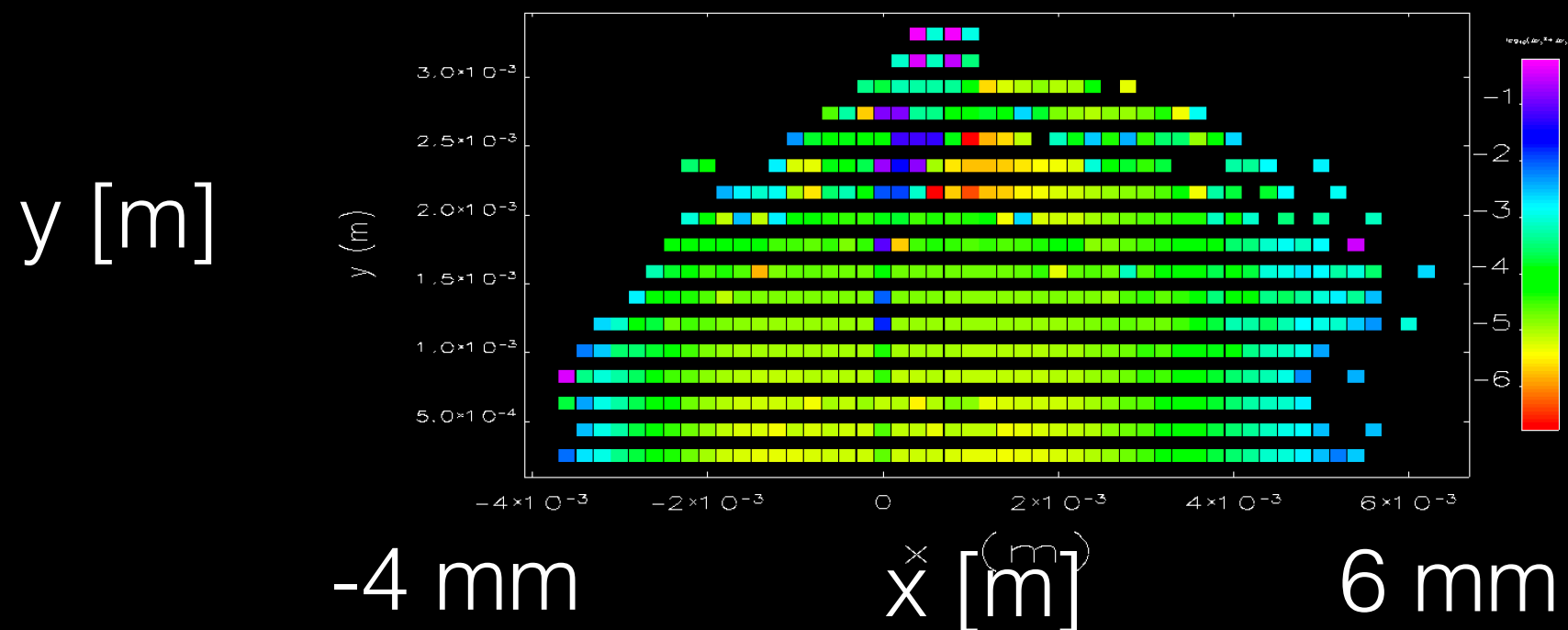
Although emittance of this lattice reduced by factor of 10,  
DA only (-4, 6)mm

$y$  [m]



# Upgrade

Although emittance of this lattice reduced by factor of 10,  
DA only (-4, 6)mm



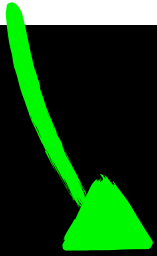
Goal: increase DA to  $\pm 10$  mm; cancel non-linear  
sextupole terms with phase advance manipulation





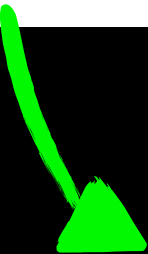


start here





start here



increase DA by  
cancelling non-  
linearities with  
phase advance  
manipulation



start here

The diagram consists of a white rectangular box at the top left containing the text 'start here' in red. A red arrow points from the bottom of this box to a text block below it. A black curved arrow points from the text block to the text 'large beta' on the right.

increase DA by  
cancelling non-  
linearities with  
phase advance  
manipulation

large  
beta



start here

increase DA by  
cancelling non-  
linearities with  
phase advance  
manipulation

large  
beta

need  
strong quads

```
graph LR; A[start here] --> B[large beta]; B --> C[need strong quads]; C --> D[large natural chromaticity]; A -- red arrow --> E[increase DA by cancelling non-linearities with phase advance manipulation];
```

start here

large  
beta

need  
strong quads

large  
natural  
chromaticity

increase DA by  
cancelling non-  
linearities with  
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start here

increase DA by  
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large  
beta

need  
strong quads

large  
natural  
chromaticity

correct  
with  
strong  
sextupoles

start here

increase DA by  
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phase advance  
manipulation

large  
beta

need  
strong quads

large  
natural  
chromaticity

strong  
non-linearities

correct  
with  
strong  
sextupoles



start here

large  
beta

need  
strong quads

large  
natural  
chromaticity

correct  
with  
strong  
sextupoles

strong  
non-linearities

DA reduction

increase DA by  
cancelling non-  
linearities with  
phase advance  
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start here

large  
beta

need  
strong quads

exit code error:  
wish me luck

large  
natural  
chromaticity

correct  
with  
strong  
sextupoles

strong  
non-linearities

DA reduction

increase DA by  
cancelling non-  
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UNIX PEOPLE ARE HAPPY



Things I enjoy doing  
outside work







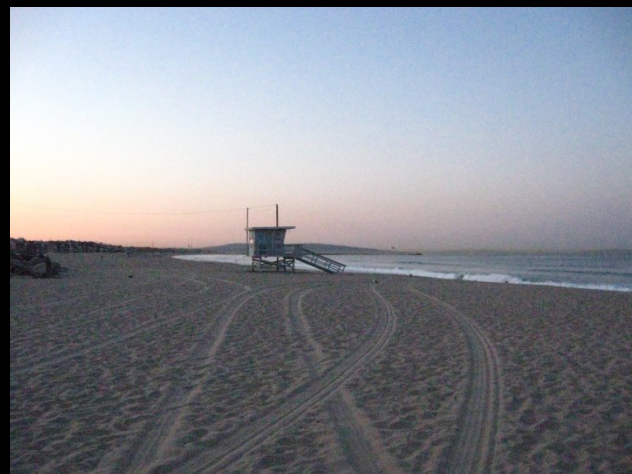




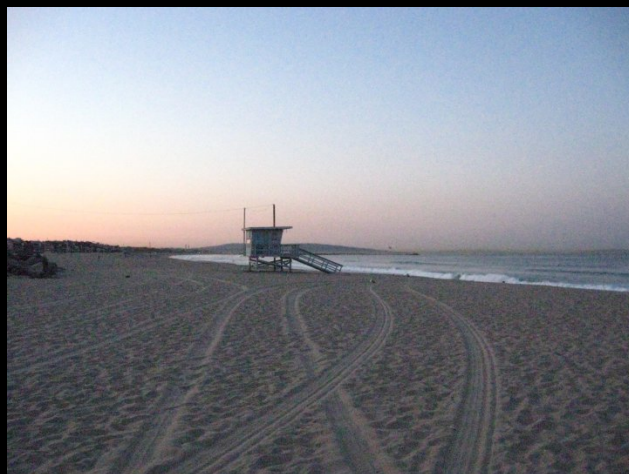




















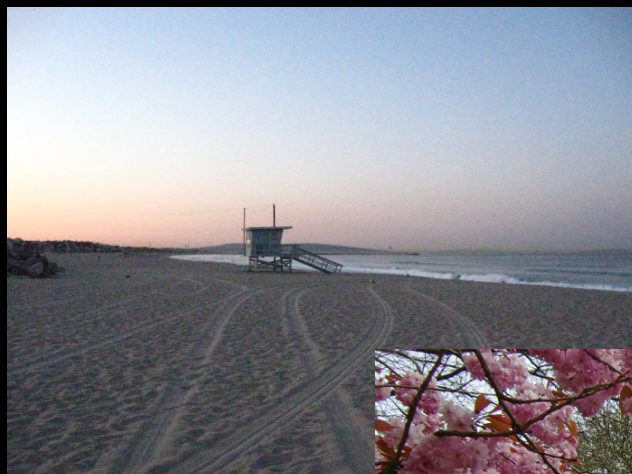


















































Thank you very much!

—Any questions?

# References

- [1] “Interim Design Report”, The IDS-NF Collaboration, arXiv: 1112.2853, IDS-NF-020, BNL-96453-2011, CERN-ATS-2011-216, EURONU-WP1-05, FERMILAB-PUB-11-581-APC, RAL-TR-2011-018, e-Print: arXiv:1112.2853 [hep-ex], Dec 2011
- [2] “Ionisation Cooling Lattices for the Neutrino Factory”, Androula Alekou, Mar 2012, PhD Thesis, MAP-DOC-4334
- [3] “Bucked coils lattice: A novel ionisation cooling lattice for the neutrino factory”, A. Alekou, J. Pasternak, Aug 2012, JINST 7 (2012) P08017, DOI: 10.1088/1748-0221/7/08/P08017
- [4] “Conceptual design and modelling of particle-matter interaction cooling systems for muon based applications”, D. Stratakis and H. Sayed et al., Phys.Rev.ST Accel.Beams 17 (2014) 071001
- [6] “The mass-hierarchy and CP-violation discovery reach of the LBNO long-baseline neutrino experiment”, LAGUNA-LBNO Collaboration, Jan 2014, arXiv:1312.6520 [hep-ph]
- [7] “Design Options of a High-Power Proton Synchrotron for LAGUNA-LBNO”, A. Alekou et al., IPAC-2013-THPWO081, May 2013
- [8] “High-Power Proton-Synchrotron Collimation Studies”, IPAC-2014-MOPRI109-June 2014
- [9] “The Double-double Bend Achromat (DDBA) Lattice Modification for the Diamond Storage Ring”, Richard Walker et al., Conference: C14-06-16, p.MOPRO103